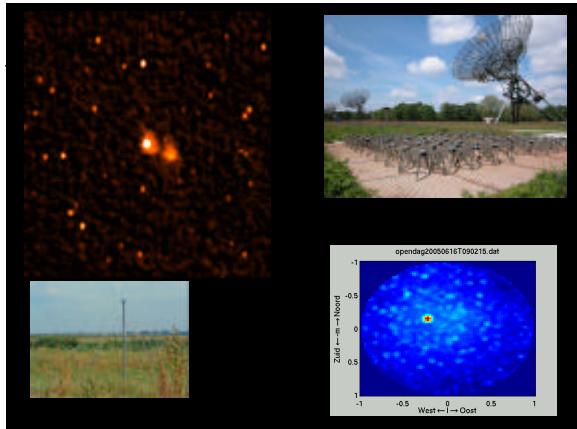


Early history and near future

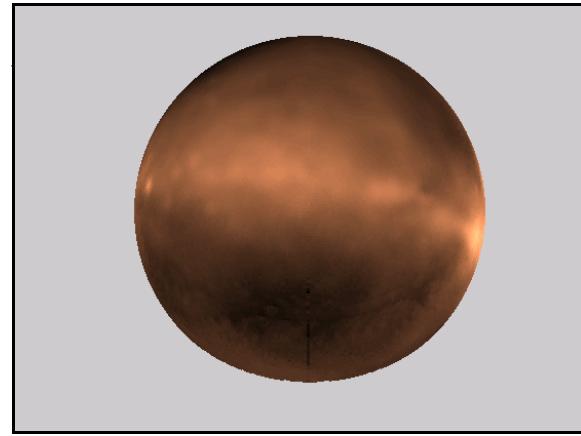
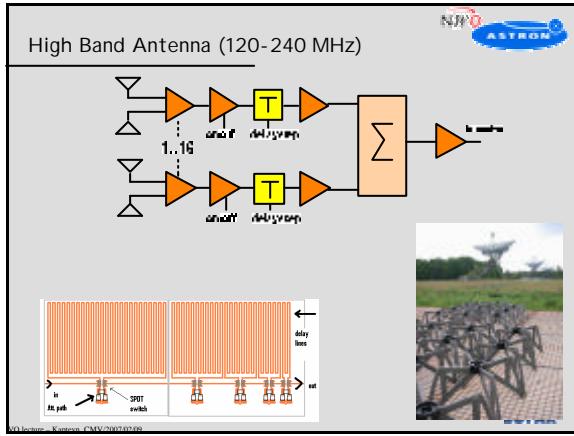
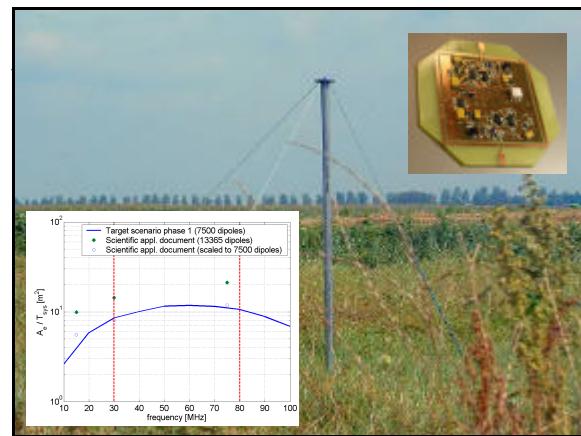
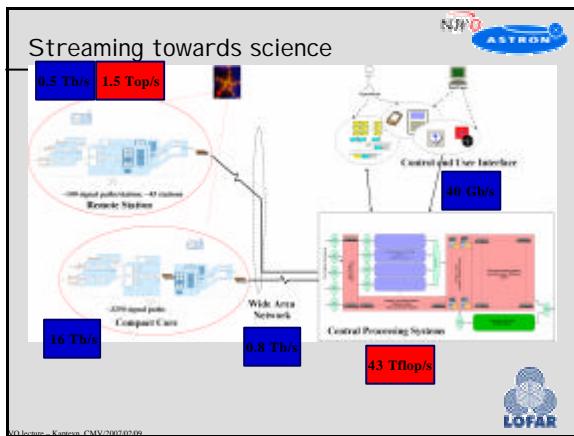
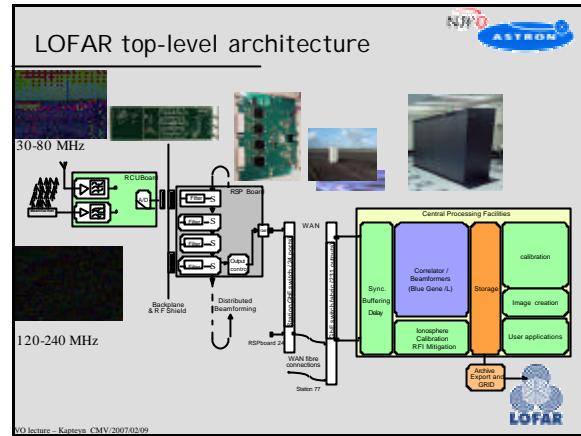
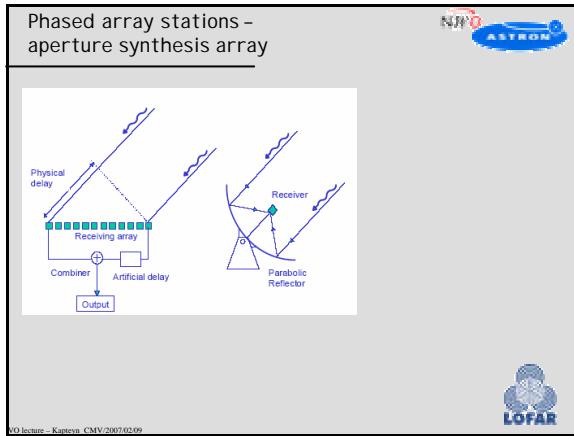
- Start of radio-astronomy:
 - Grote Reber, Karl Jansky @ 20MHz (~1930)
- Next large new facility:
 - LOFAR @ (20)-30-80 MHz and 120-240 MHz
- Why the gap?
 - Relatively easy to get some signal out
 - Quite difficult to get further than ~300 sources



LOFAR in summary

- Large distributed radio telescope
 - 32 central (24 Gbps) + 45 remote (2 Gbps) phased array antenna stations, each ~4 soccerfields in size
 - Full scale aperture synthesis array, extends 100 km
 - ~ 10 long-distance stations being discussed (E-LOFAR)
- Two main bands
 - High Band ~ 7,500 tiles 120 – 240 MHz
 - ~ 120,000 m² / 25K @ 150 MHz
 - Low Band ~ 7,500 dipoles 10 – 80 MHz
 - ~ 375,000 m² / 45,000K @ 20 MHz
- Digital Radio 40 MHz processing bandwidth
 - Extreme agility in time/space/frequency
 - Large instantaneous sky coverage
- Fibre network, Software Correlator
- New calibration & RFI mitigation schemes







LOFAR Prototype Station (LOPES):
detection of nanosecond radio flashes from
ultra-high energy elementary particles



Falcke et al. (2005), *Nature*, Vol. 435, p. 313




Thunderstorm Events

- Does the Electric field of the atmosphere influence CR radio signal?
- For $E > 100$ V/cm E-field force dominates B-field:
 - Fair weather: $E = 1$ V/cm
 - Thunderstorms: $E = 1$ V/cm
- Select thunderstorm periods from meteorological data:
 - Clear radio excess during thunder storms
 - B-field effect dominates under normal conditions
 - >90% duty cycle possible





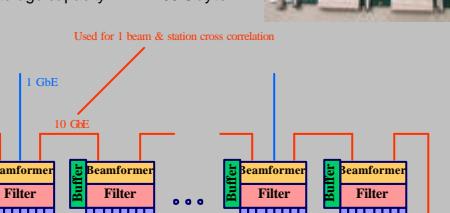


Station based processing

- Input data rate: ~ 460 Gbps
- Output data rate: ~ 2 Gbps
- Processing capacity: ~ 1.5 Tmuls/s
- Storage capacity: 96 Gbyte



Used for 1 beam & station cross correlation

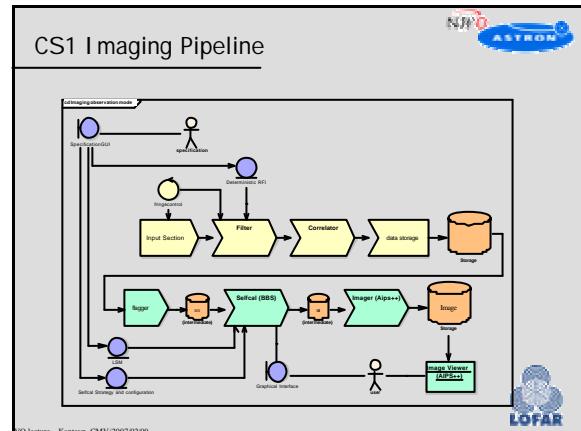
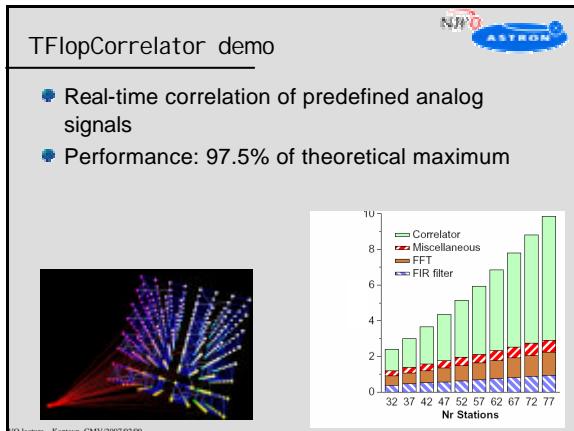
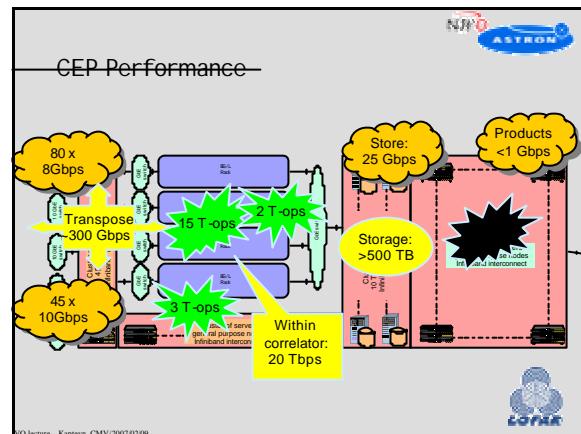
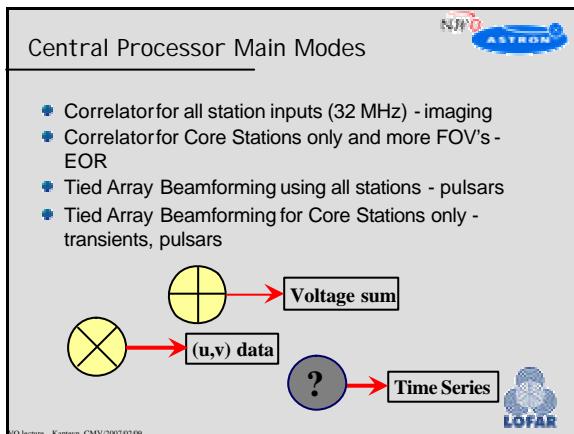
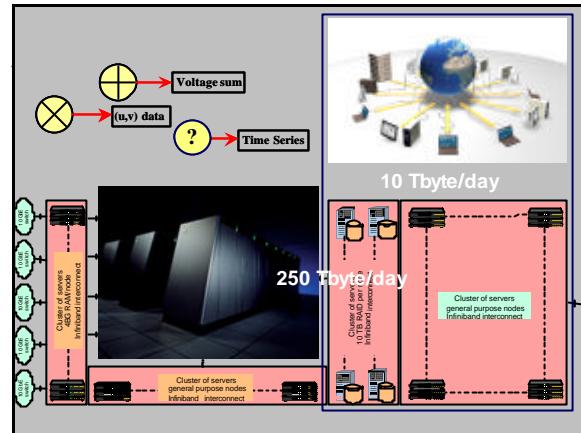
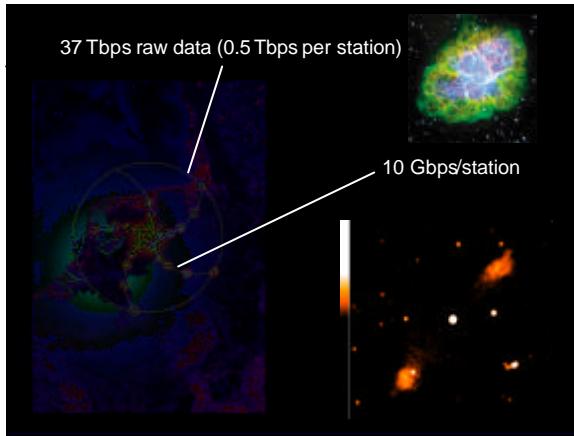


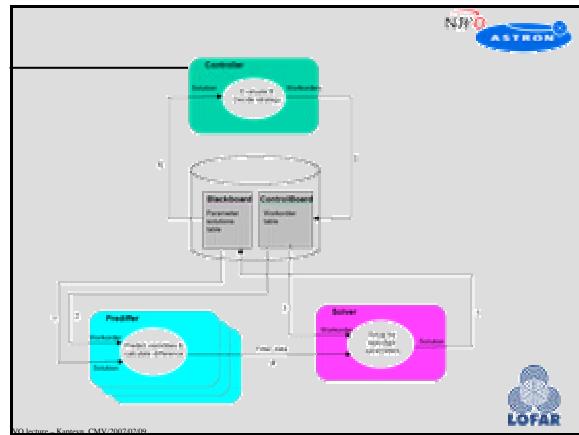
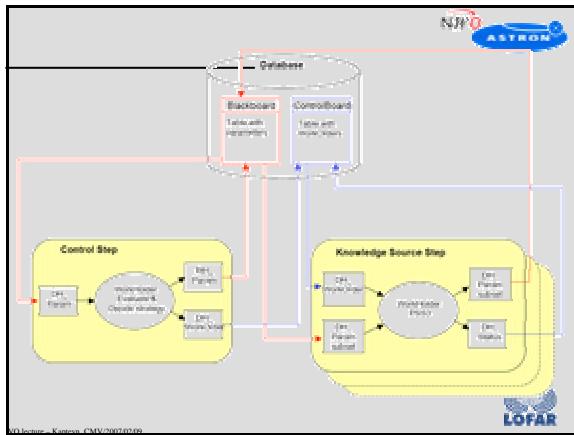
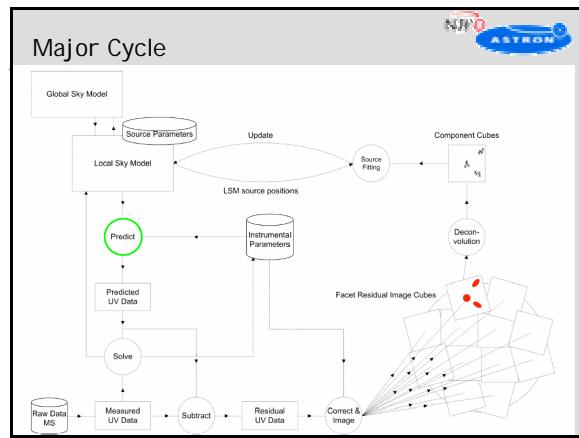
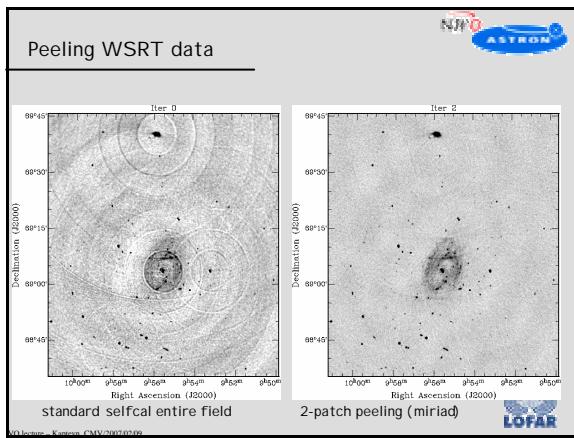
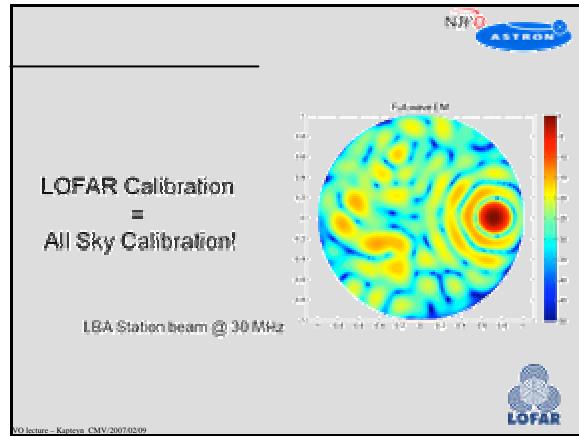
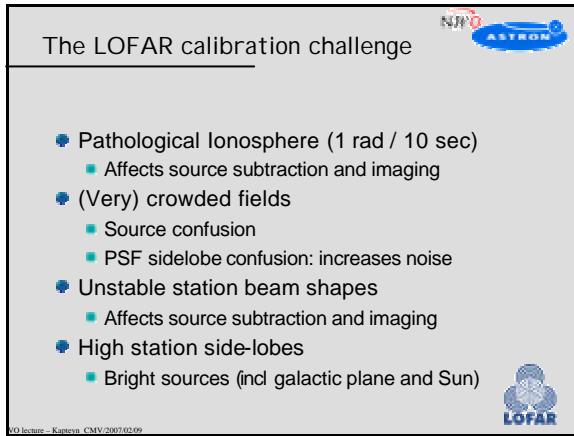
ITS Spatial Filter Experiment

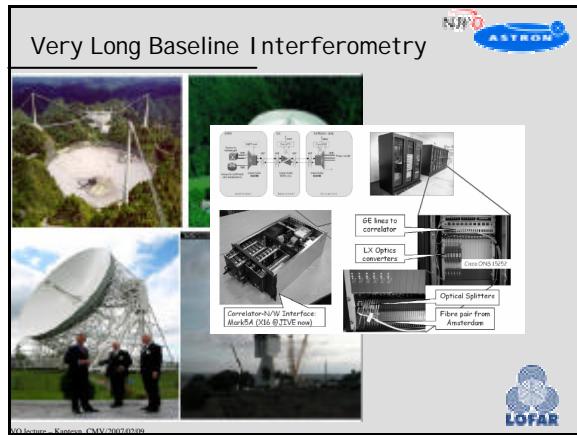
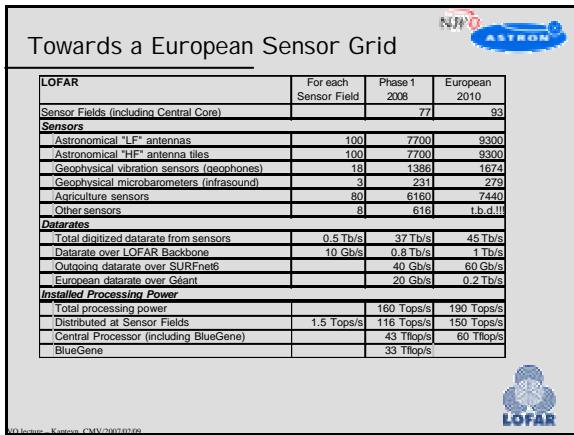
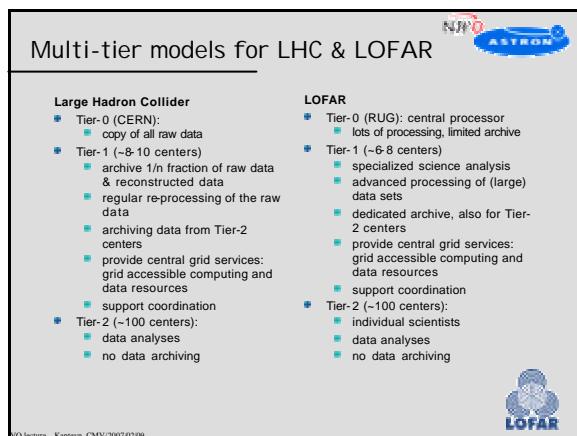
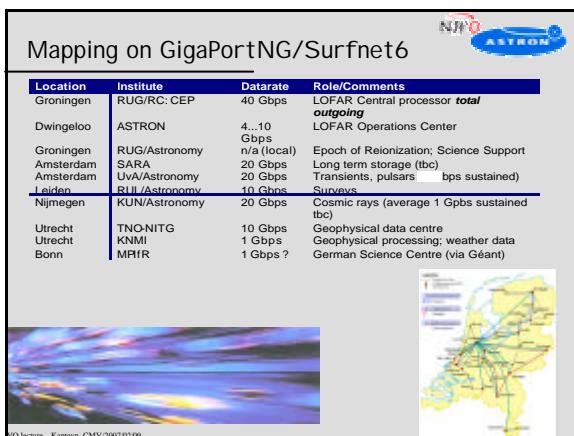
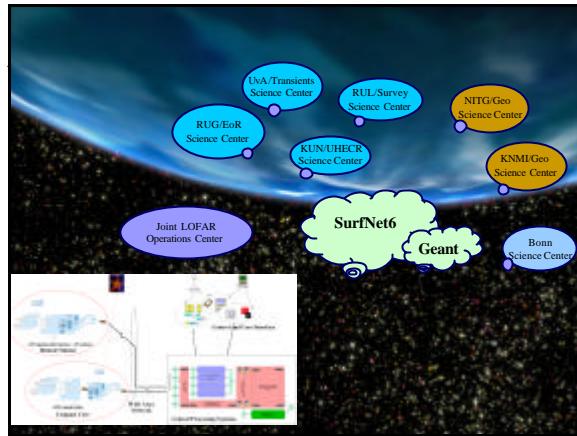
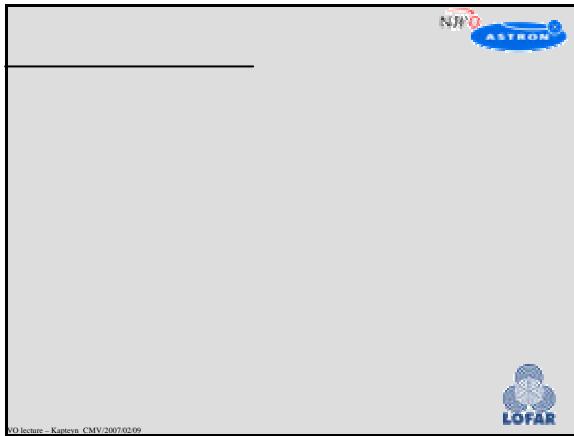
ITS sky map at 27.85 MHz, no RFI
Two strong sources (Cas A, Cyg A) visible

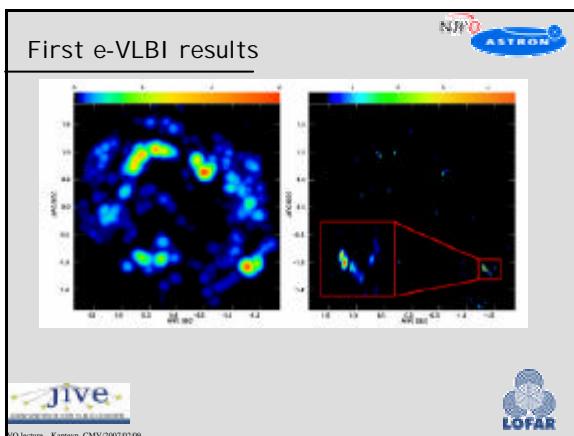
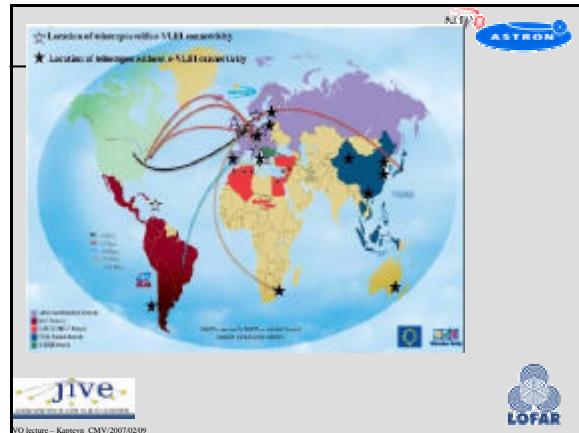
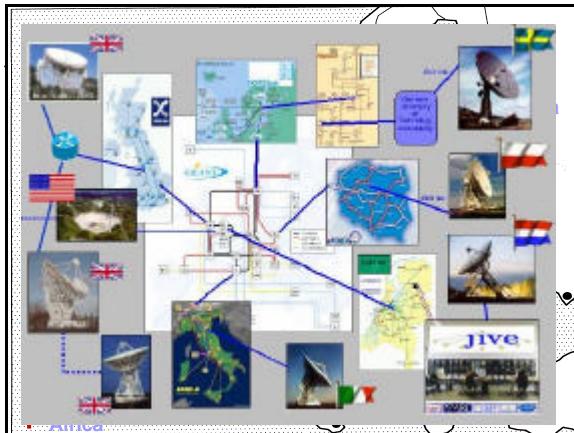
ITS sky map at 27.81 MHz, strong RFI at $(Az, el) = (-1.3, 0)$ rad

ITS sky map at 27.81 MHz, fixed null at $(Az, el) = (-1.3, 0)$ rad









LOFAR Performance

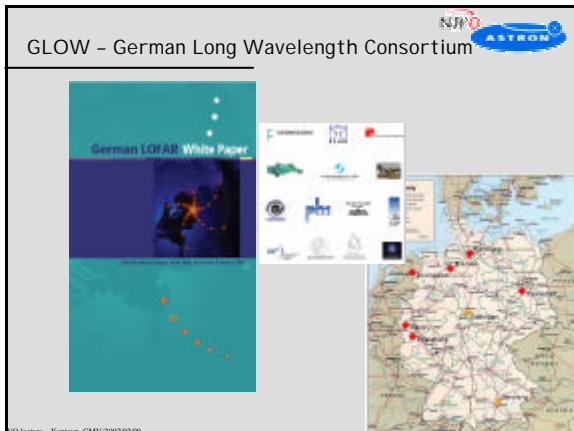
Frequency (MHz)	A_{eff} (m^2)	T_{sys} (in K)	dS in 1s (mJy)	dS in 10h (mJy)	dS in 100h (mJy)
30	3.3×10^5	23k	68	0.35	0.11
75	5.2×10^4	2450	46	0.24	0.07
120	3.3×10^5	820	2.4	0.013	0.004

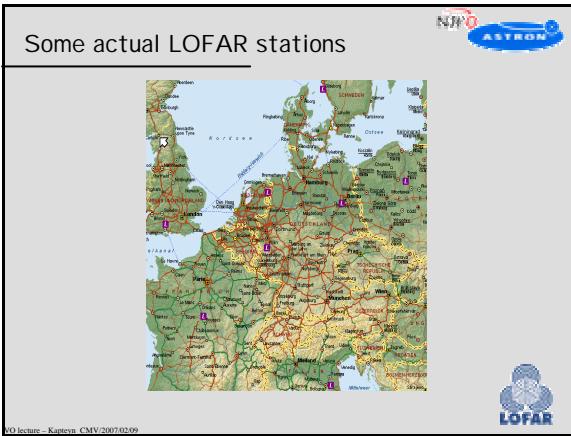
Approximate sensitivity per beam, with 4 MHz BW and for a single polarization

NRAO ASTRON

LOFAR

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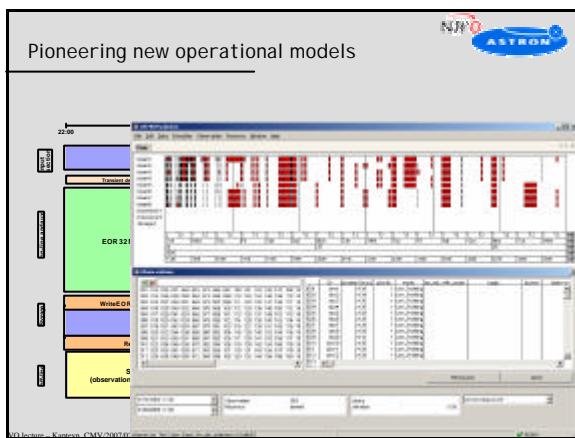


Baselines and Resolution

Site:	Latitude	Longitude	D(km)	240				115		80		30 MHz	
				125	261	375	999 cm						
CS-1	52,927		6,877	0				50	5,2	10,8	15,5	41,2	"
LOFAR 77								267	1,0	2,0	2,9	7,7	"
Effelsberg	50,533		6,883					422	0,6	1,3	1,8	4,9	"
Potsdam	52,400		13,067					595	0,4	0,9	1,3	3,5	"
Onsala	57,400		11,917					602	0,4	0,9	1,3	3,4	"
Chilbolton	51,144		-1,436					702	0,4	0,8	1,1	2,9	"
Nancay	47,383		2,200					786	0,3	0,7	1,0	2,6	"
Torun	53,017		18,617					997	0,3	0,5	0,8	2,1	"
Medicina	44,521		11,645										

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LOFAR



The path to the SKA

	Long Haul data network			Central processor	Output data rate
	Distance	Net data rate to processor	Technology		
LOFAR	160 km	800 Gb/s	8 GbE CWDM	43 T flop/s	10 TByte/day
E-LOFAR	2000 km	1 Tb/s	10 GbE	40 T flop/s	20 TByte/day
SKADS demo.	singl. tile	30 Gb/s	10 GbE	10 T flop/s	0.3 Tbyte/day
APEXIF	2 km	1.6 Tb/s	10 GbE	0.5 Pflop/s	200 Tflops/day
e-VLBI	9000 km	2 Tb/s	t.b.d.	1 Pflop/s	400 Tflops/day
SKA	>3000 km	10 Tb/s	t.b.d.	~ 10 Pflop/s	~ Poflops/day

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