



MeerChoirs: Unveiling the resolved HI in nearby Groups with MeerKAT (Part 1)

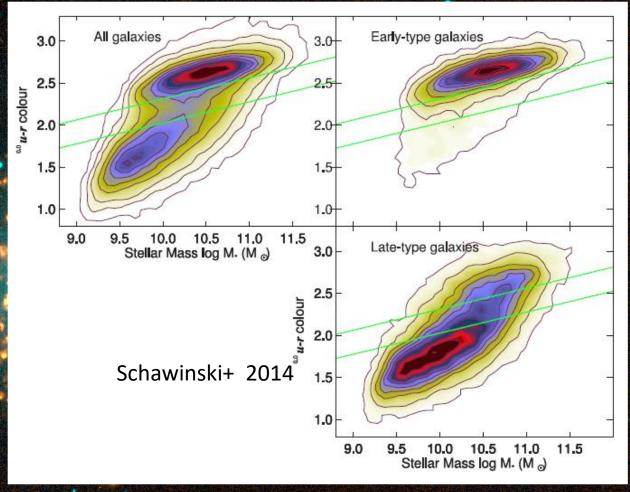
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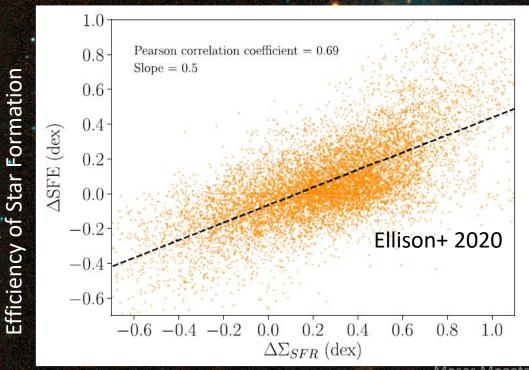
Galaxy Evolution

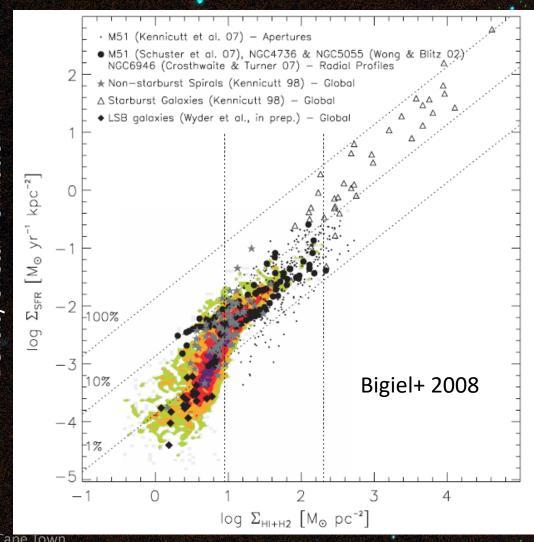
- Early Type Galaxies
 - Gas Poor
 - Low Star Formation Rate
- Late Type Galaxies
 - Gas Rich
 - Star Forming
- Evolve from Blue to Red



Galaxy Evolution – Star Formation

- Gas is needed to form stars
 - Need cold gas
 - How much gas is turned into stars?
 - How do galaxies maintain fuel?

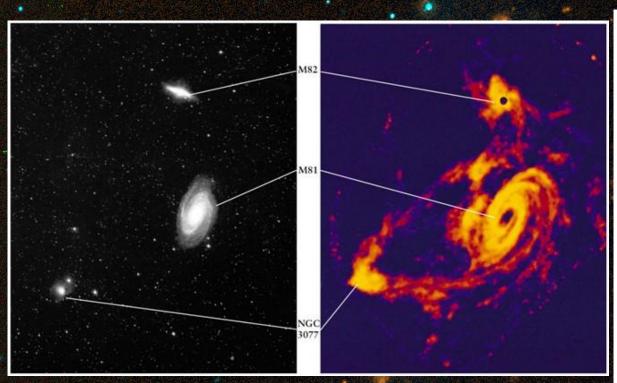


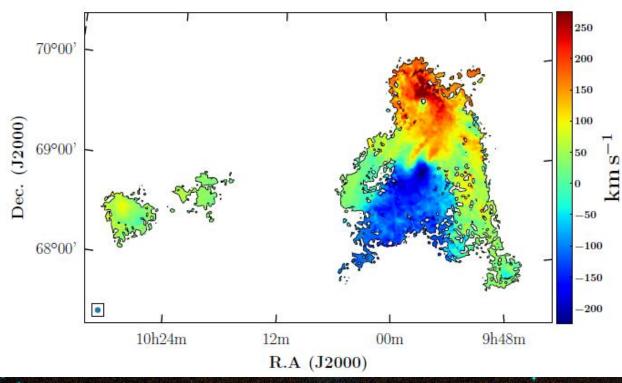


Distance away from Star Forming Sequence

Gas Density

Galaxy Environments





E. McCormick lecture

Sorgho+ 2018

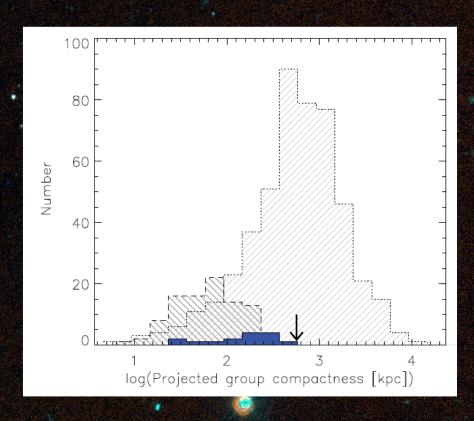
SINGG & Choirs

- HI (HIPASS) based sample
- SINGG: Hα, R, UV, observations (Meurer+ 2006, Hanish+ 2006, Wong 2006)
- Choir Groups
 - Multiple (4+) $H\alpha$ sources within a HIPASS detections (Sweet+ 2006)

Choir Groups Properties

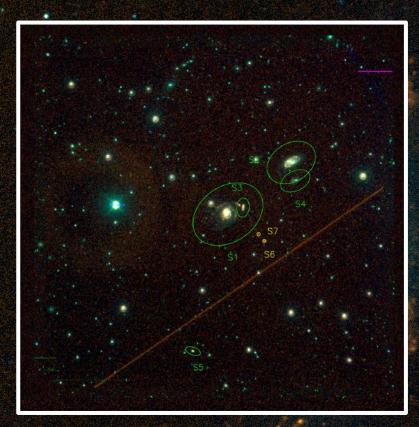
- Distance: 23 151 Mpc
- Group M_{HI}: 9.92 10.61 logM_{sol}
- Emission Line Group Members: 4-10

- Lots of star forming galaxies
- No bright ellipticals associated with groups

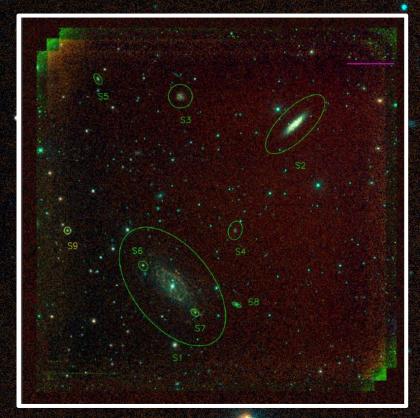


Sweet+ 2013

Choirs Groups

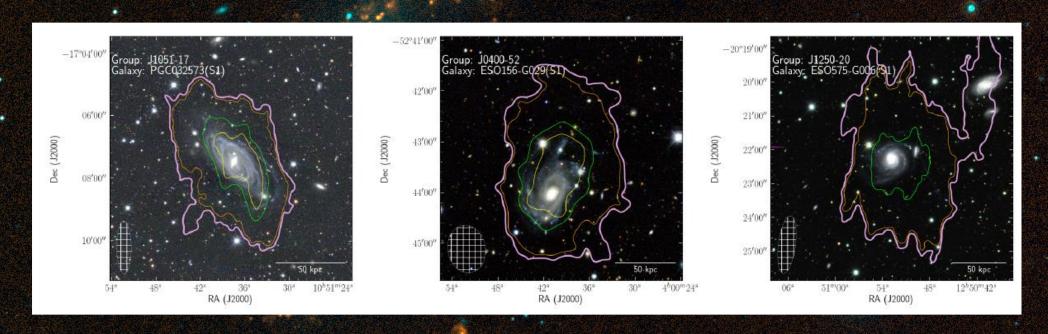


J1250-50, Sweet+ 2013

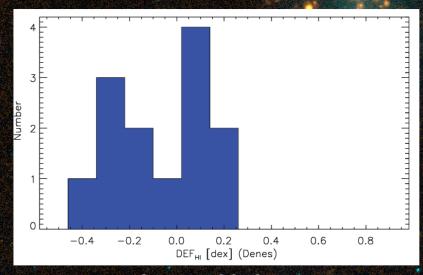


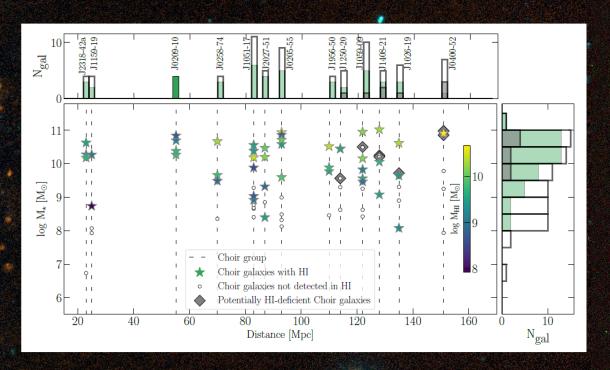
J1051-17, Sweet 2013

- Previous ATCA and VLA observations of *some* of the groups
- Dzudzar+ 2019, 2021

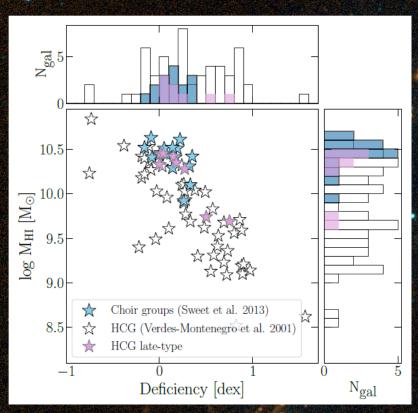


- Groups are mostly not HI deficient
- Seem to be in early stages of evolution



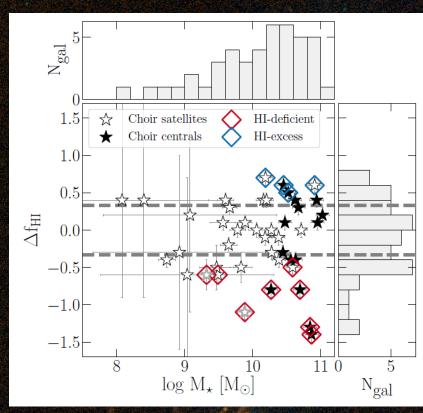


Dzudzar+ 2021

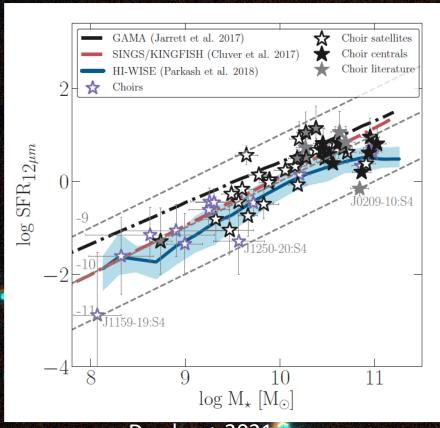


Dzudzar+ 2021

Dzudzar+ 2021



Dzudzar+ 2021



Dzudzar+ 2021

MeerChoirs

- Resolved HI studies of the galaxies and gas within Choir Groups with MeerKAT
- Study the evolution of galaxies and atomic gas within groups
- 32k, ~6 hours per group
- ~10-20" arcsec spatial resolution (2-5x better)
- Utilizing the combination of high spatial, spectral resolution, sensitivity and large field of view

MeerChoirs

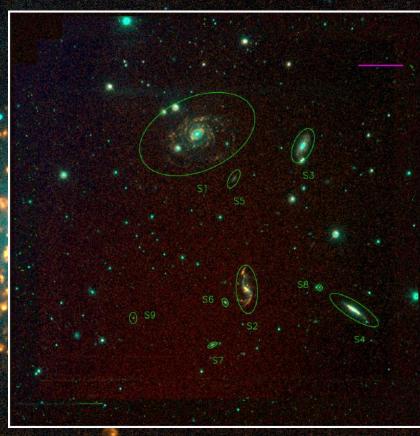
- MeerKAT Observations
 - Pilot of 6 Groups awarded Priority B Time
 - Allocated Priority B Time to observe 8/9 of the remaining groups
 - One group was observed by a different program

MeerChoirs Preliminary

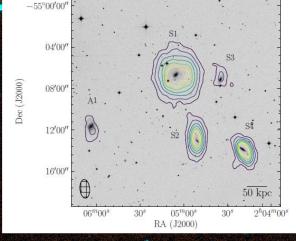
- Using the Calibrated data from MeerKAT pipeline
- CASA Imaging
- Via IDIA-Ilifu

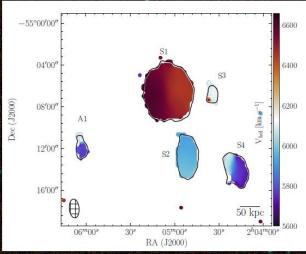
MeerChoirs

- J0205-55
- Two sub-groups,
 "Group a" in SINGG
- Optical detections: 9
- HI detections: 4



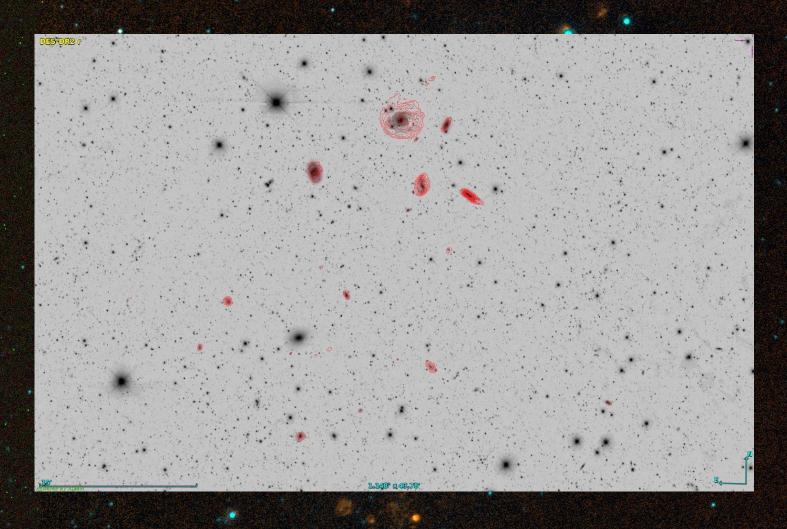
 $H\alpha$ + R, Sweet+ 2013





HI Dzudzar+ 2021

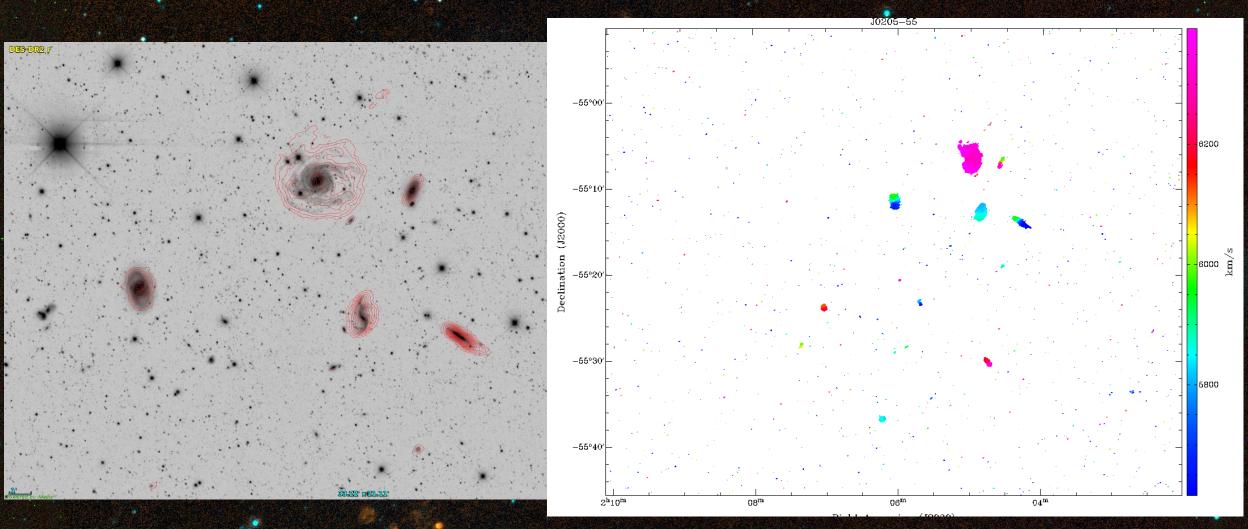
MeerChoirs Prelim Work



Preliminary Shallow Imaging

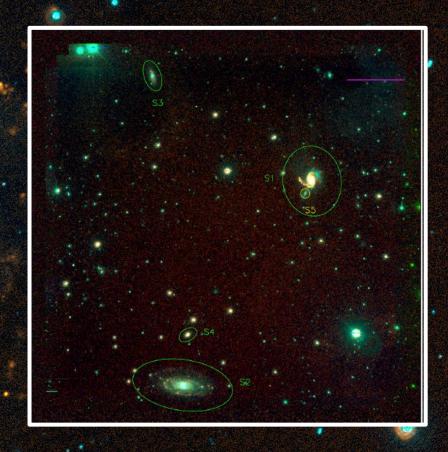
- 5 Sources detected as before
- 2 new HI detections in same FOV
- ~6/7 additional HI detections in group

MeerChoirs Very Preliminary



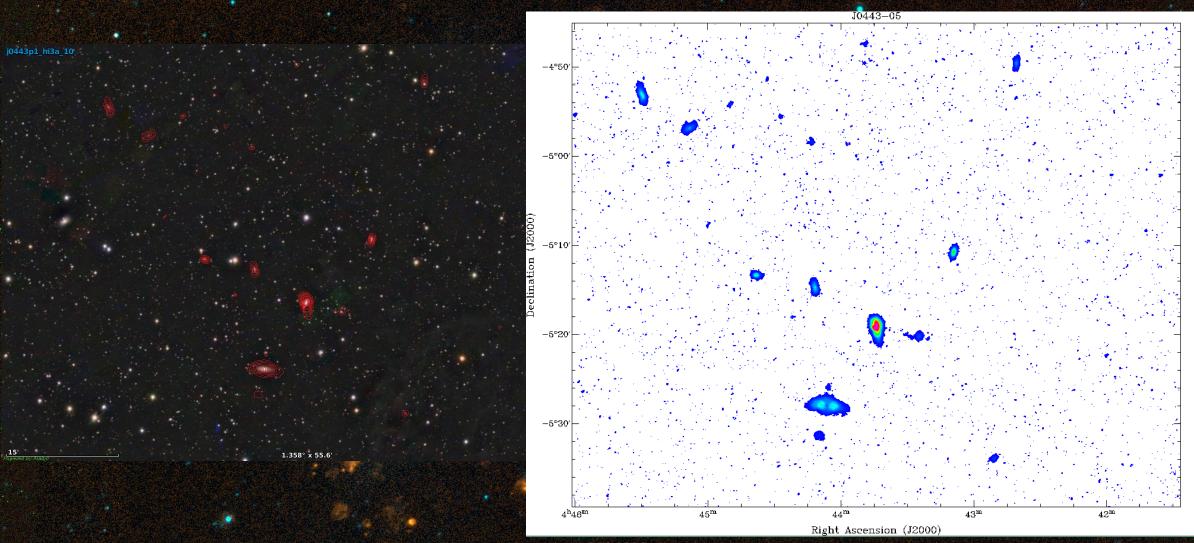
MeerChoirs

- J0443-05
- Optical detections: 5
- No resolved HI

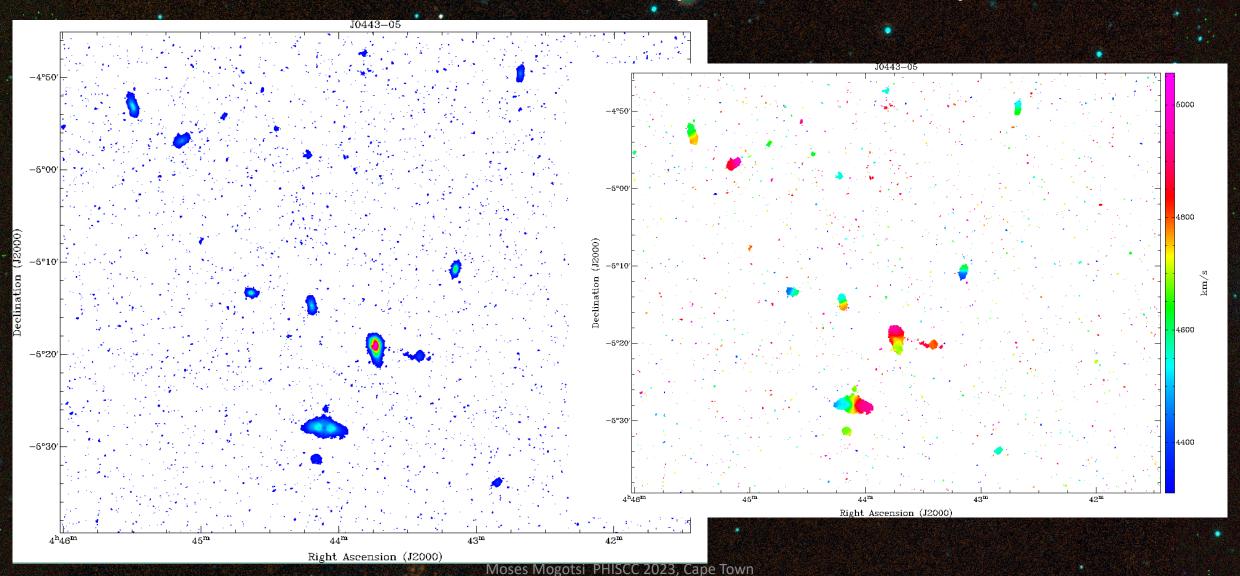


 $H\alpha$ + R, Sweet+ 2013

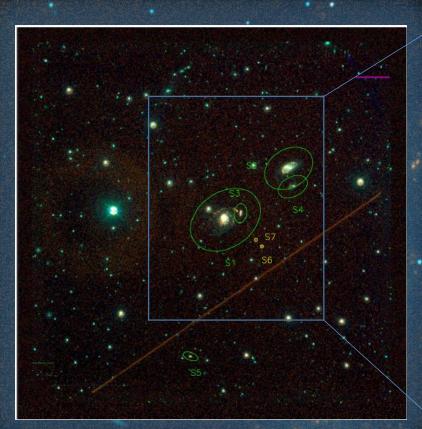
MeerChoirs... Very Preliminary



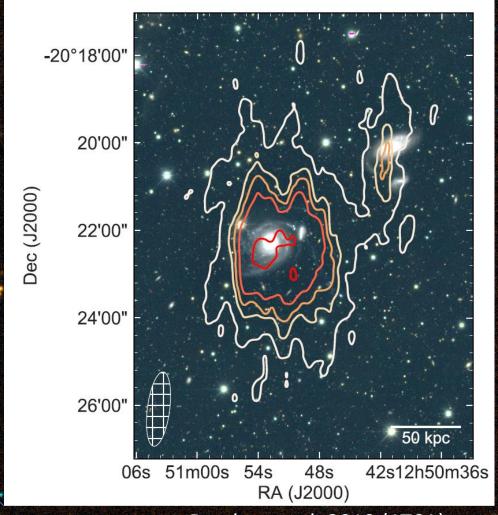
MeerChoirs... Very Preliminary



MeerChoirs

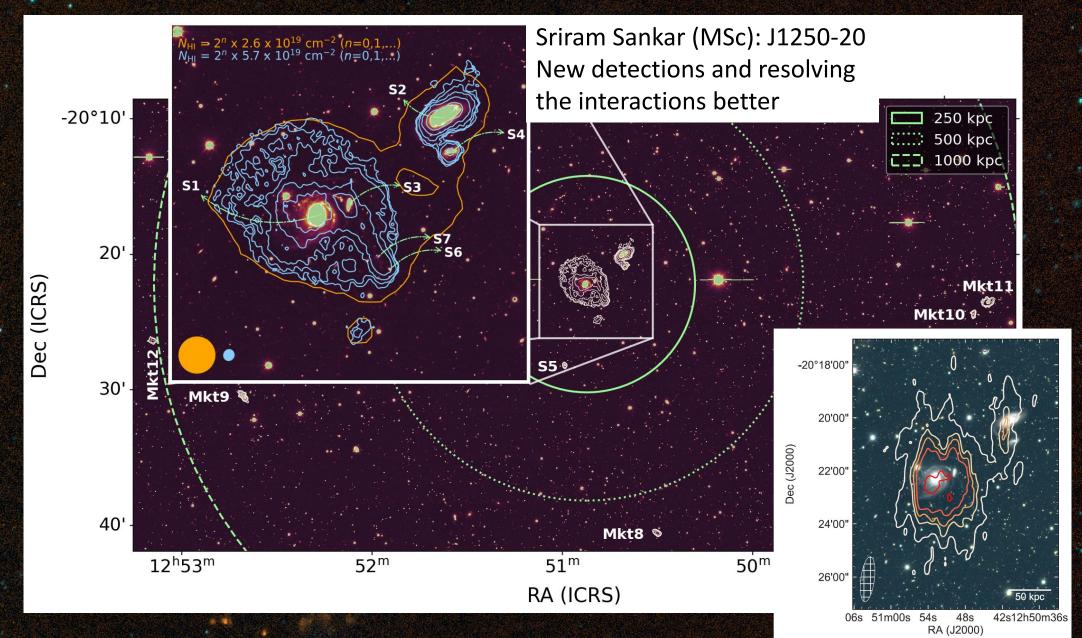


Sweet + 2019 (Optical SINGG image)



Dzudzar et al. 2019 (ATCA)

MeerChoirs Results



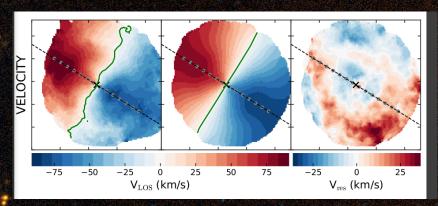
MeerChoirs Results



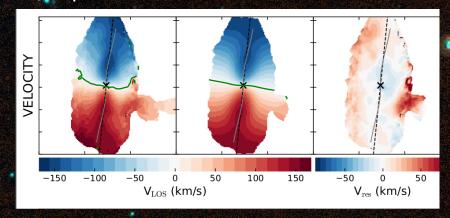
Sriram Sankar (MSc) J1250-20 Zoom in Optical + HI

MeerChoirs Exciting Results... see Sriram Sankar's Talk





J1403-06, Arp 271

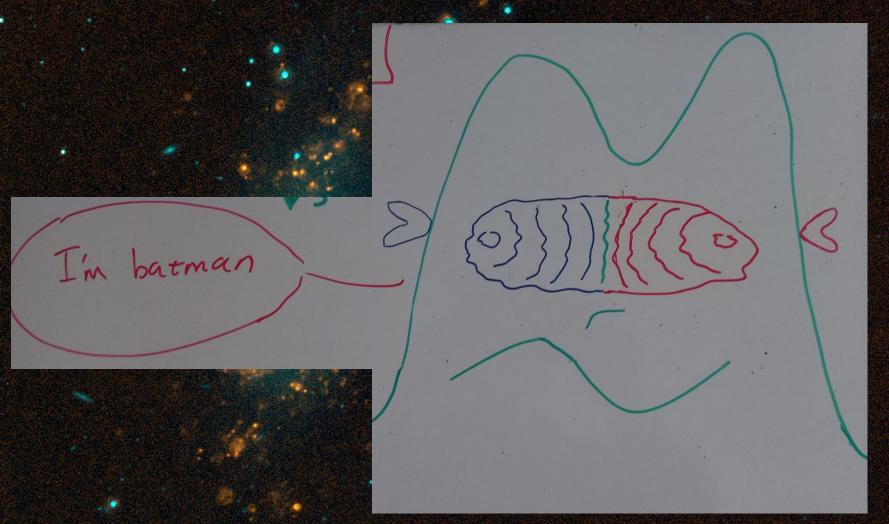


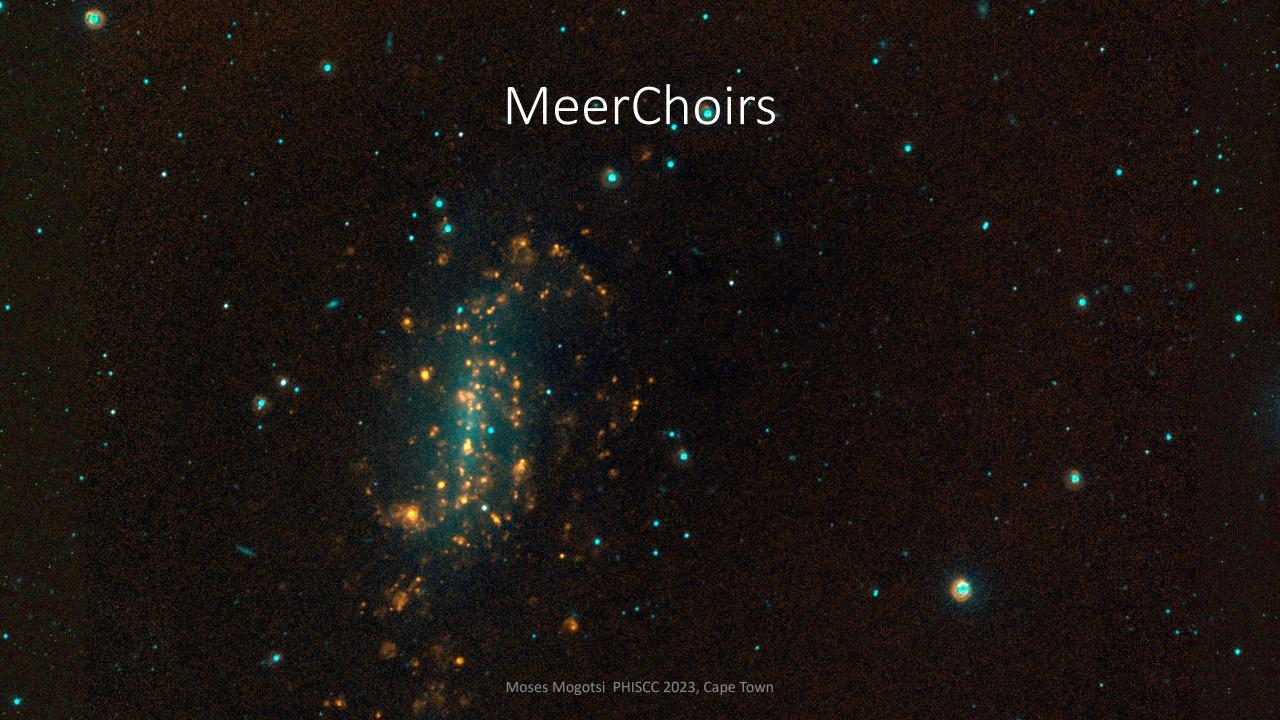
MeerChoirs... next

- Sriram to finish J1250-20 and J1403-06 initial analysis for MSc
- Refinement of HI imaging (Using Caracal?) and data reduction for the groups
- Re-reducing rest of the data
- Analysis of HI and continuum data
- Follow up any missing data
 - HI data for other interactions (LIRGs, Collisional rings)
 - Lots of work to do, Collaboration welcome



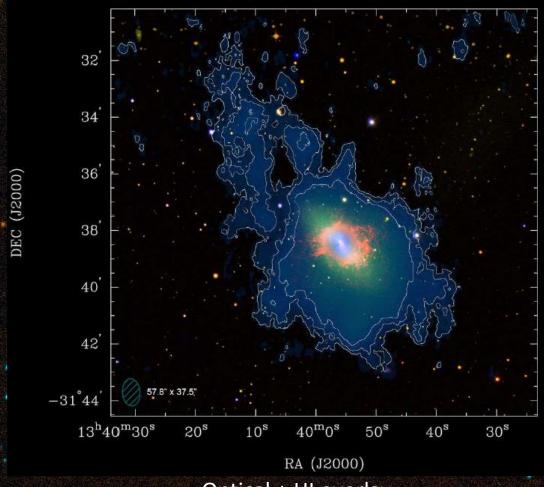
Thank You





NGC 5253

- Starburst
- HI and Ionized Gas Outflow
- Possible Infalling Cloud
- In the M83 group



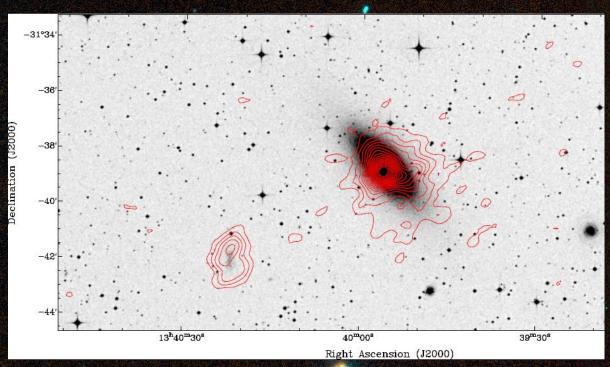
Optical + HI overlay; Lopez-Sanchez + 2012

NGC 5253

MeerKAT-16 Observations

 Detect anomalous gas, planned to propose for deeper and higher resolution observations from full MeerKAT array

HI Overlayed on Optical



With help from G. Jozsa, T. Randriamampandry

