

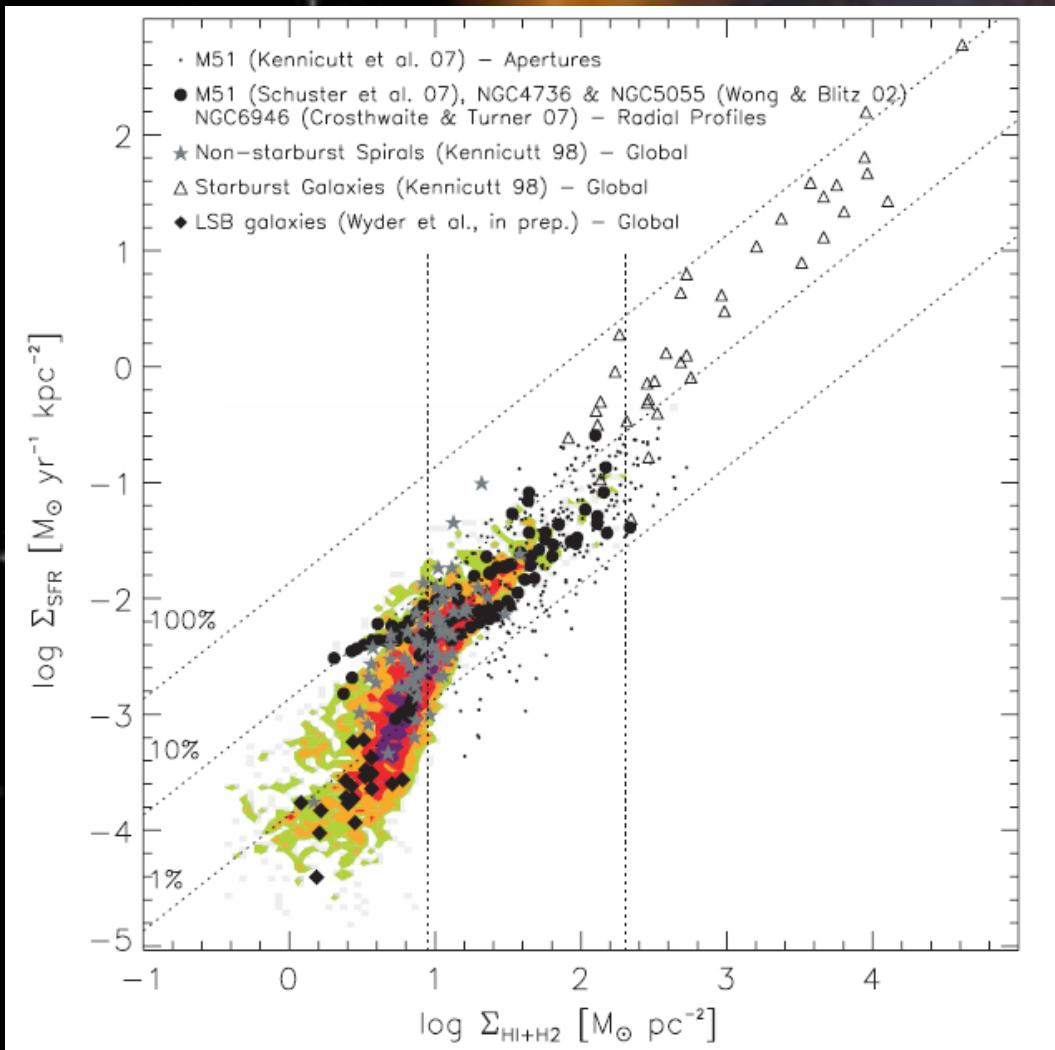
# SUNBIRD Multi- $\lambda$ KINEMATICS

Moses Mogotsi

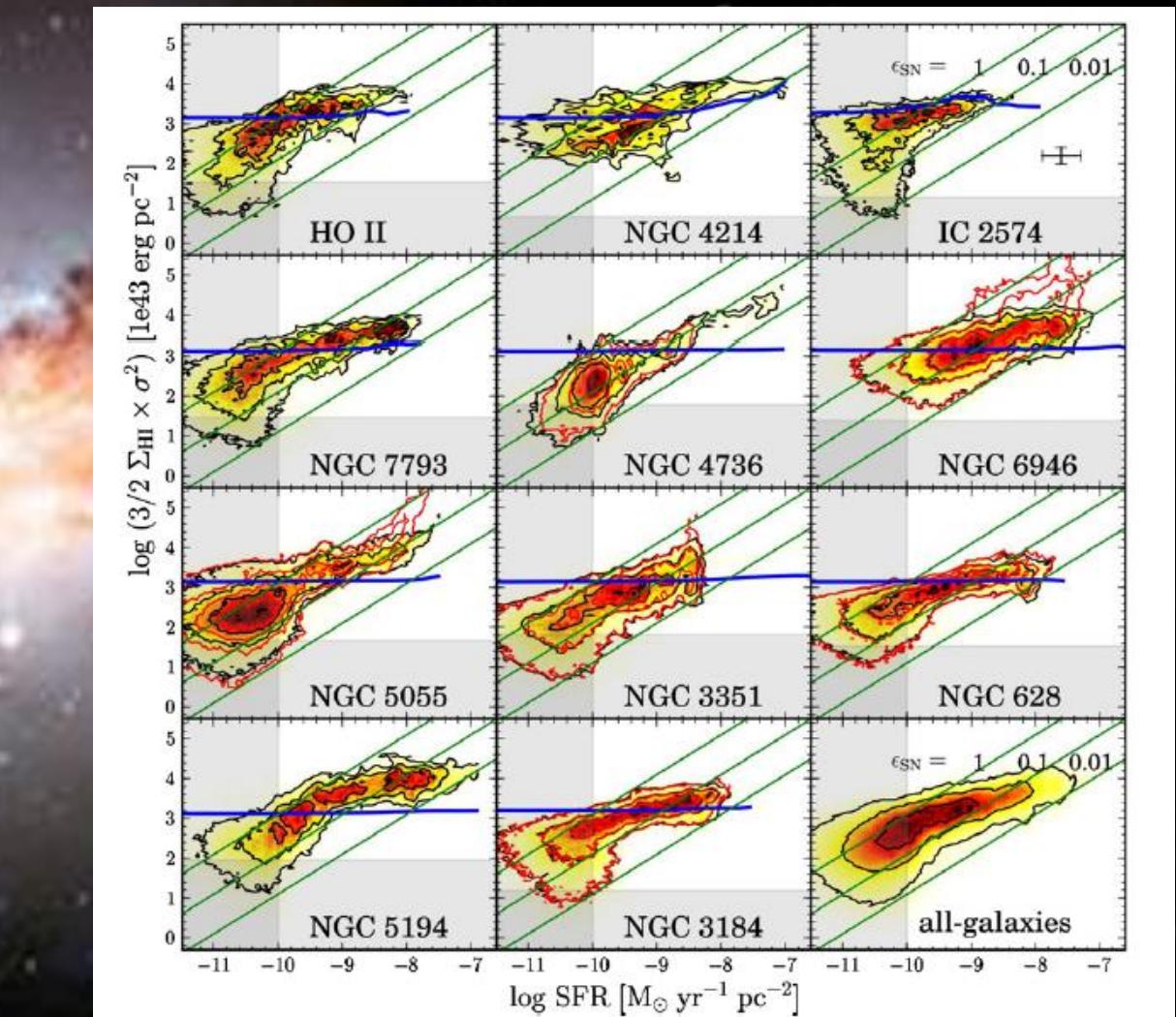
+ Väisänen P.<sup>1</sup>, Ramphul R.<sup>1</sup>, Tafere M., Kuncarayacti H.<sup>2</sup>, Kotilainen J.<sup>2</sup>, Kool E.<sup>3</sup>,  
& SUNBIRD Team

<sup>1</sup>South African Astronomical Observatory (SAAO); <sup>2</sup>University of Turku;

<sup>3</sup>Macquarie University

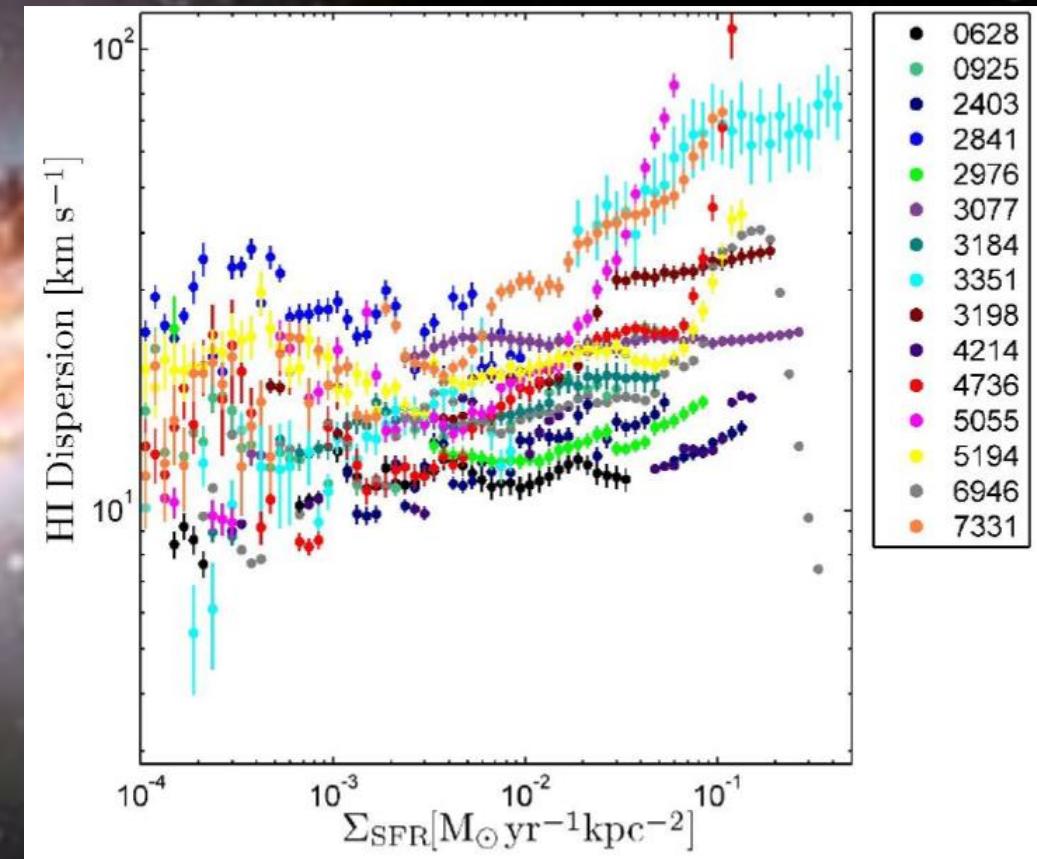
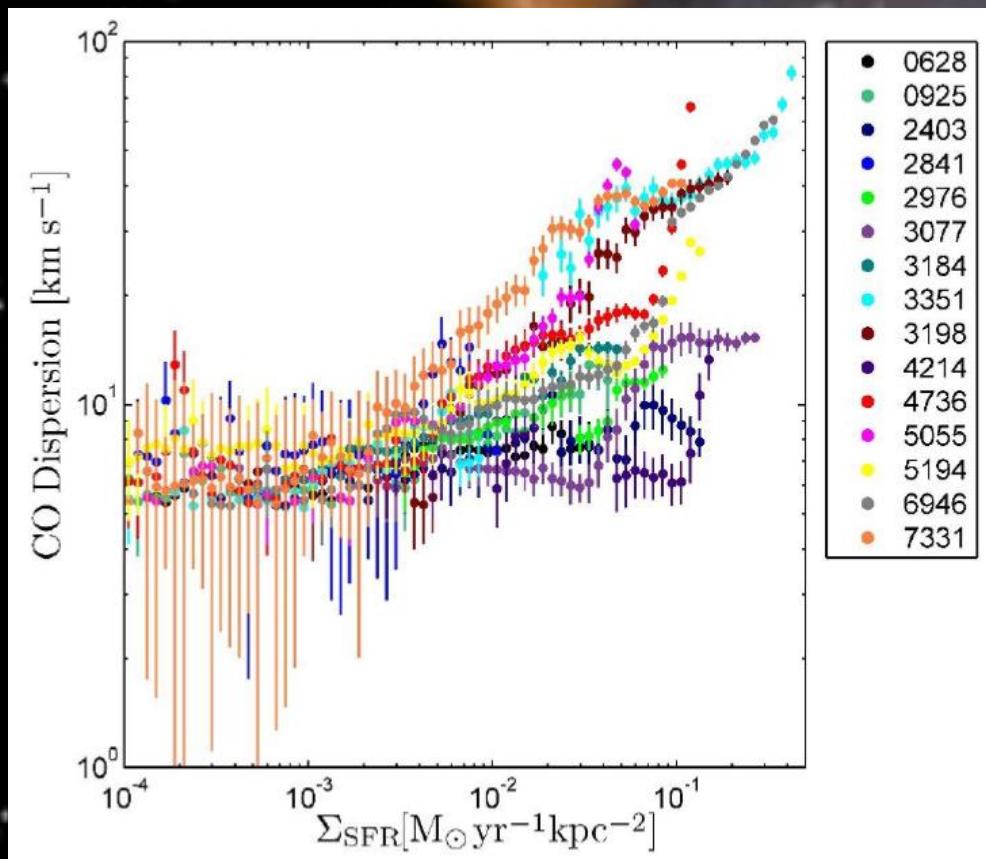


Bigiel+ 2008

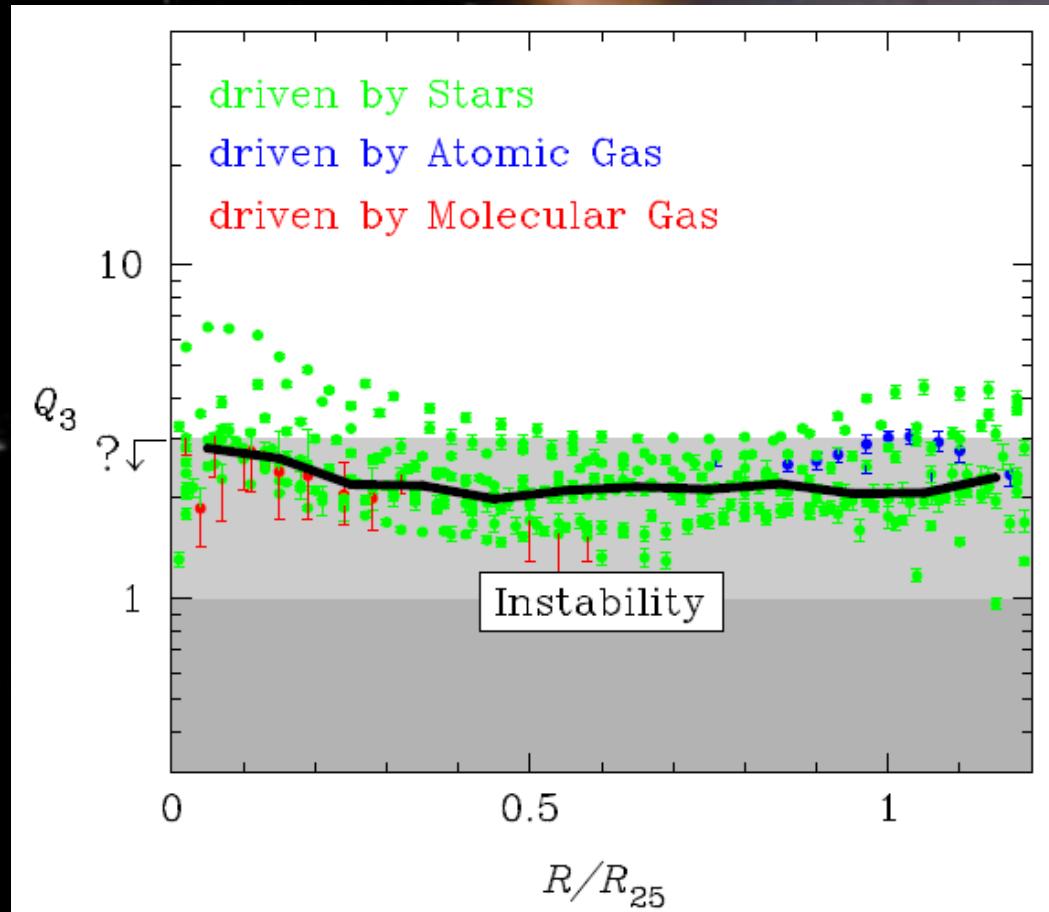


Tamburro+ 2008

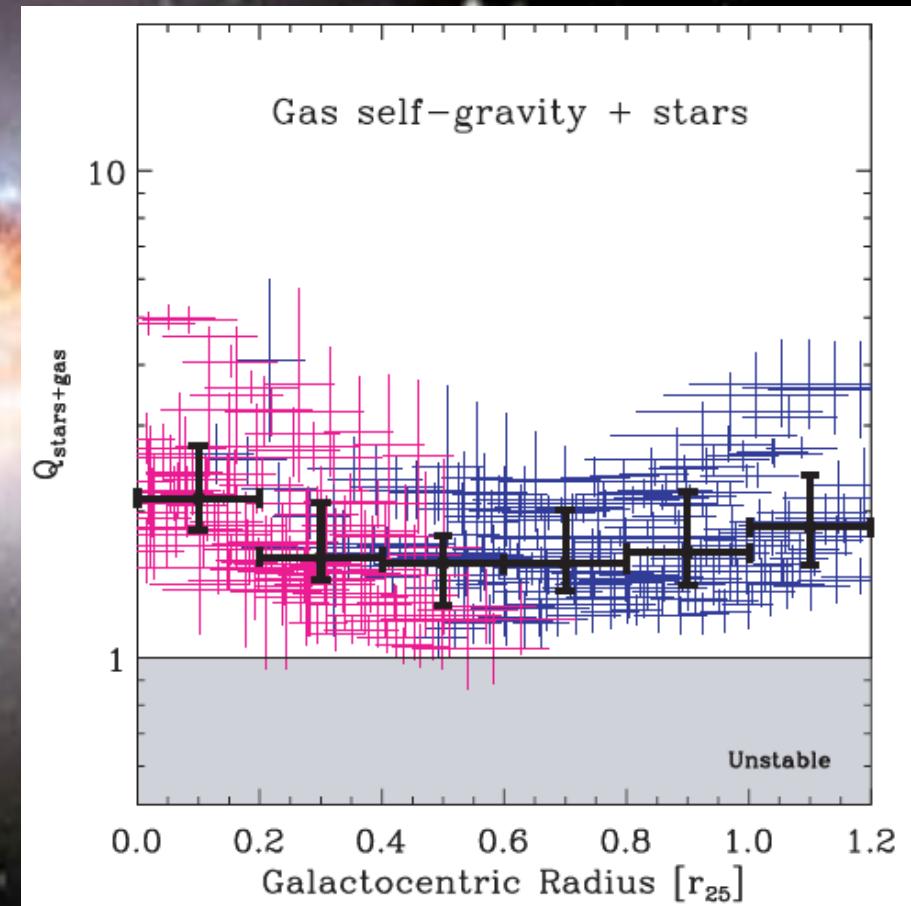
# Feedback?



# Disc Stability

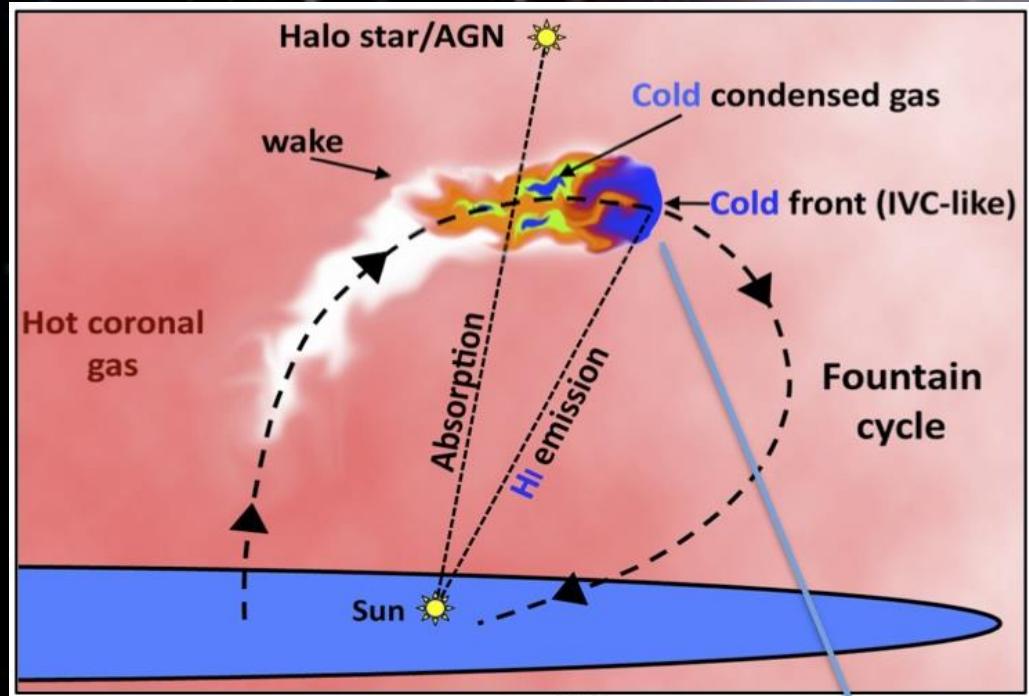


Romeo & Mogotsi 2017



Leroy+ 2008

# Feedback

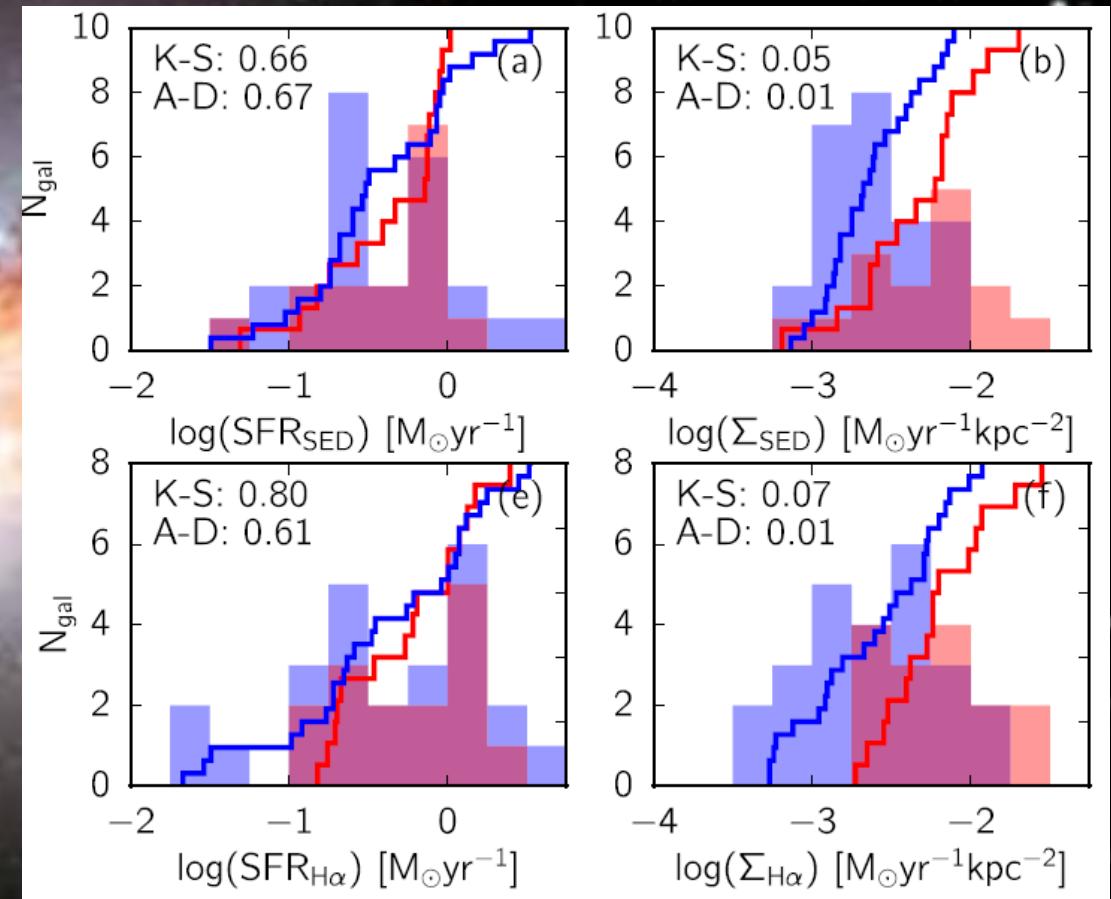


Fraternali+ 2014

Hubble Team NASA/ESA

# Galaxies are Windy...

- Wind-dominated (red) versus non-dominated (blue)
- Find winds across large ranges of SFR
- Wind dominated galaxies tend to only found at high SFR



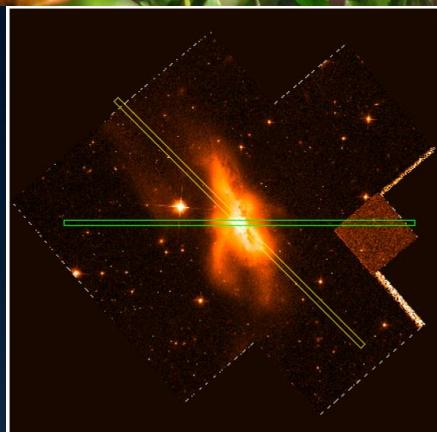
Ho+ 2016

# LIRGS/ULIRGS

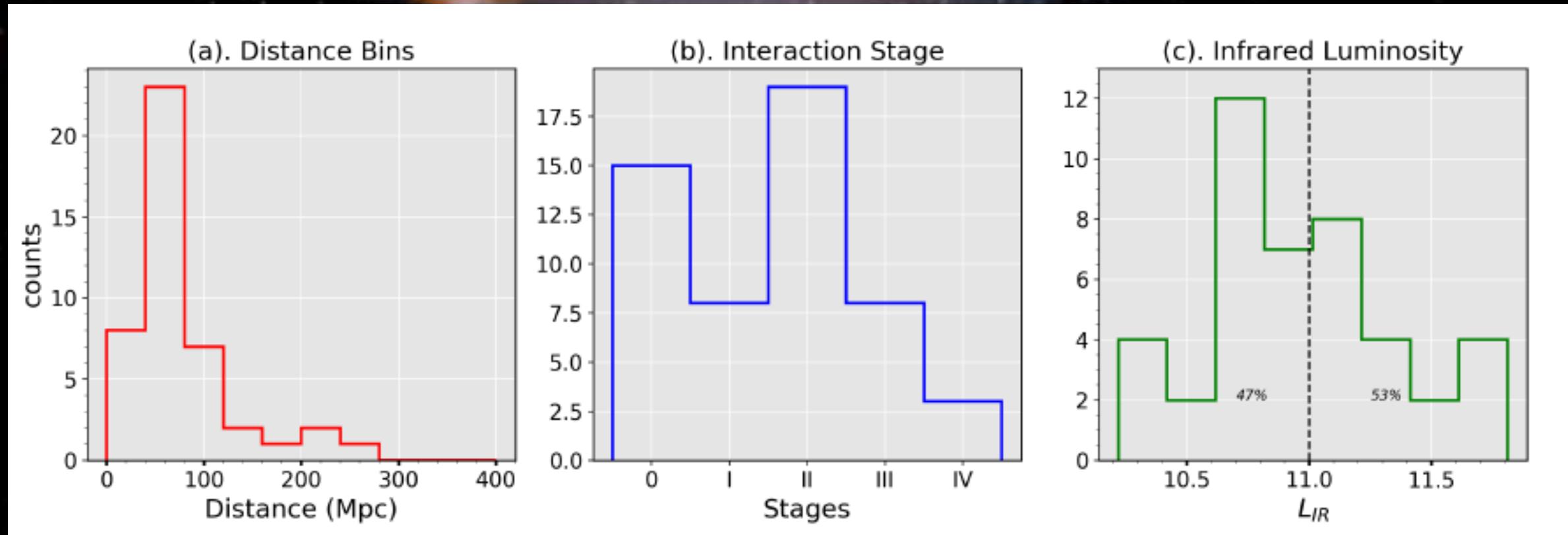
- Luminous Infrared Galaxies (LIRGS) and Ultraluminous Infrared Galaxies (ULIRGS)
- LIRGS:  $10^{11}L_{\text{sol}} < \text{LIR} < 10^{12}L_{\text{sol}}$  Sanders & Mirabel (1996)
- ULIRGS:  $10^{12}L_{\text{sol}} < \text{LIR} < 10^{13}L_{\text{sol}}$
- Gas Rich
- Most show signs of interaction or merger (higher fraction of mergers at higher luminosity)

# SUNBIRD Kinematics Observations

- SUperNovae and StarBursts and in the InfraReD
- SALT RSS Observations of starbursts, LIRGS, ULIRGS...
- 54 Targets
- 155 Observations
- PG900 : Stellar Populations
- PG1800, PG3000 : Kinematics



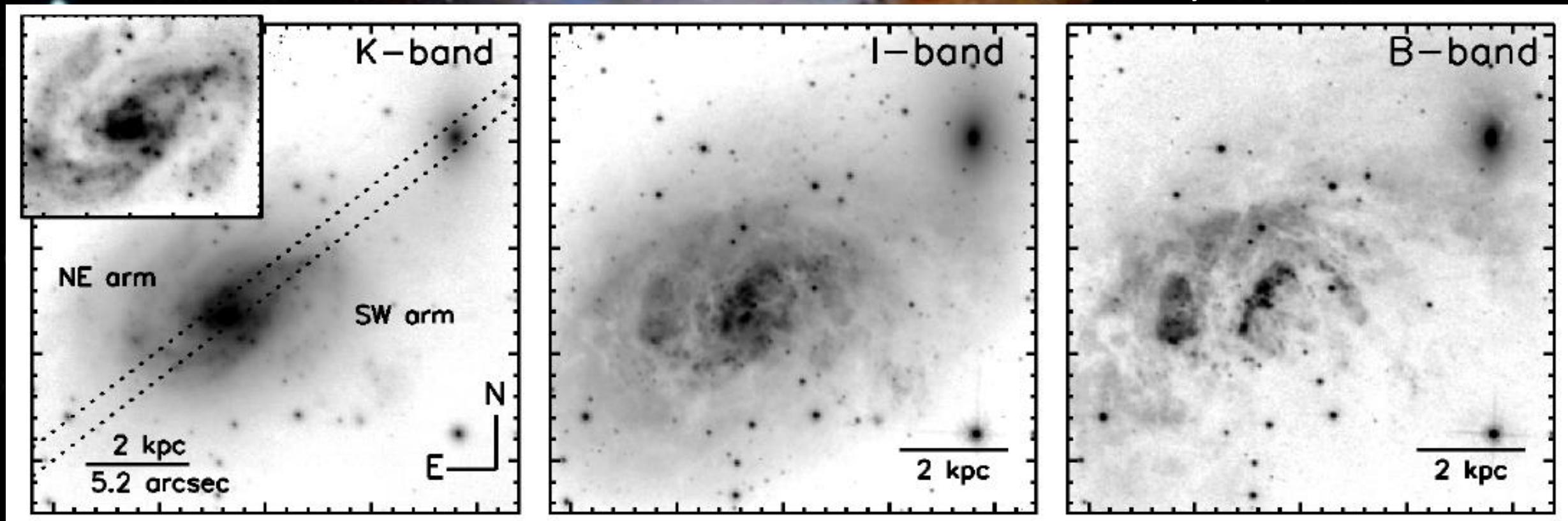
# SUNBIRD Observations (SALT)



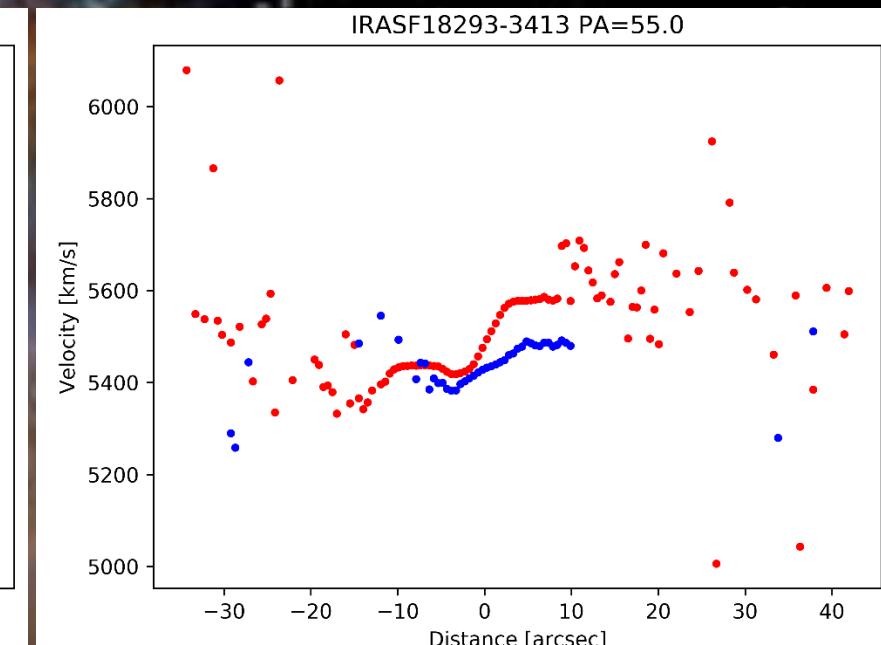
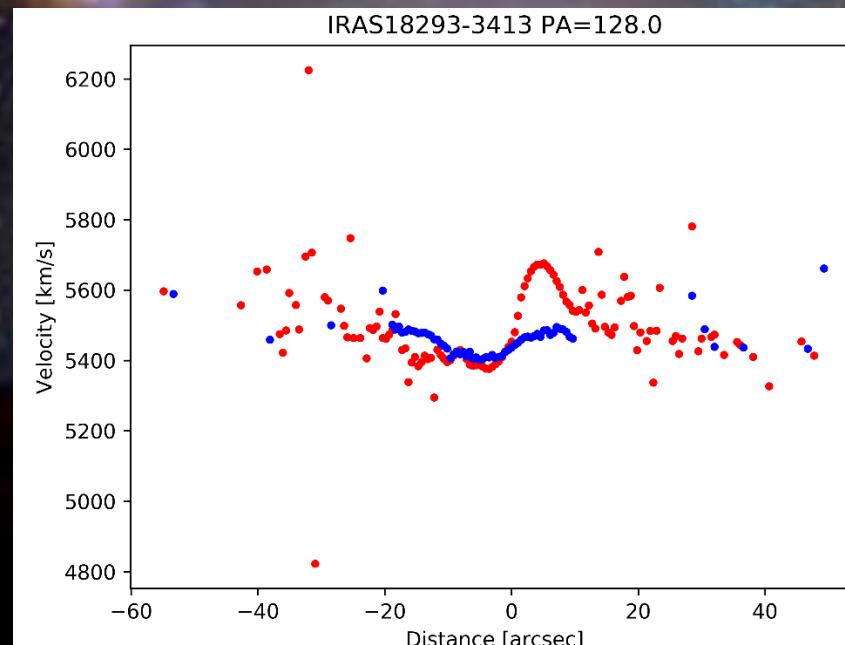
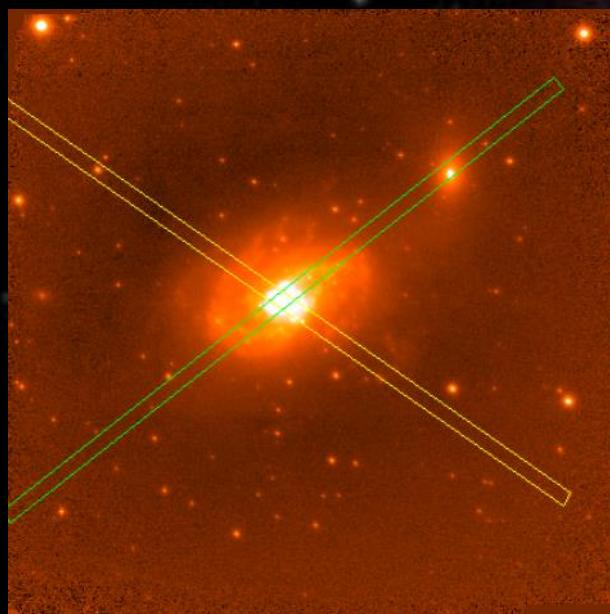
Rajin Ramphul PhD Thesis

# IRAS18293-3413

- D=75.7Mpc
- Stellar Mass =  $10^{10.63}$  Msol
- SFR =  $10^{2.04}$  Msol/yr



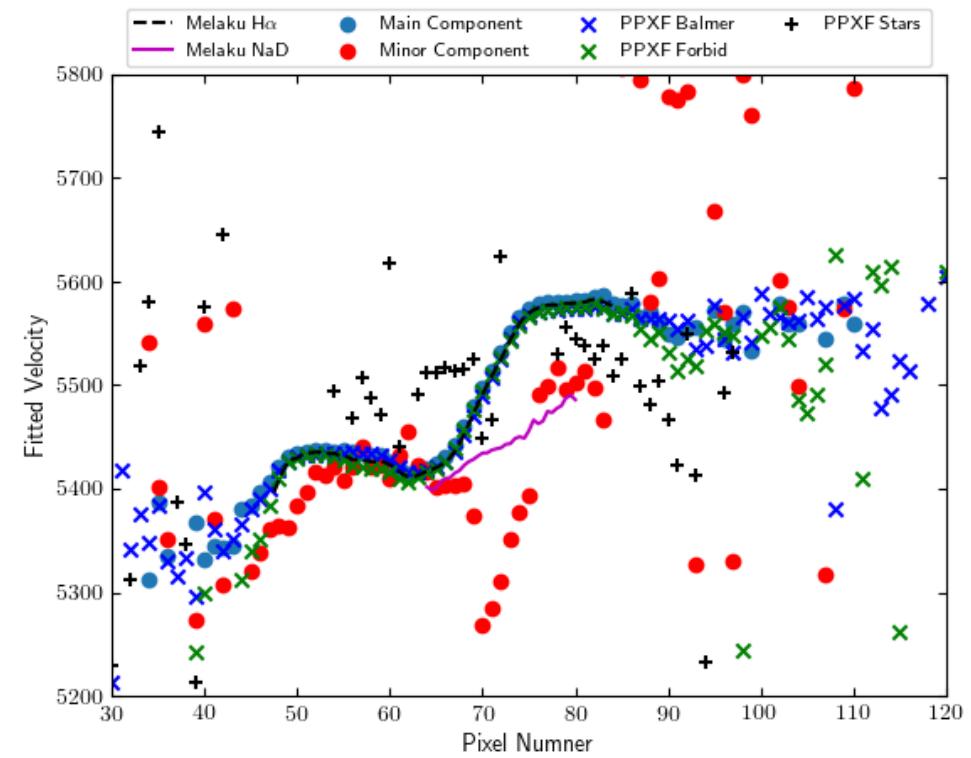
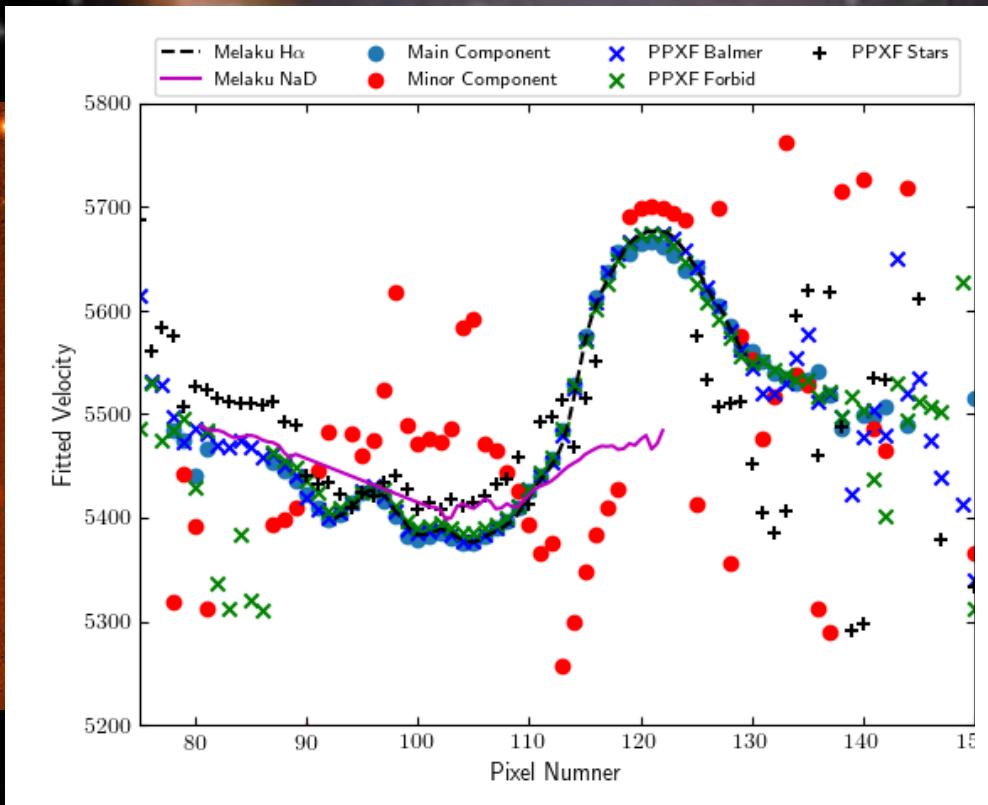
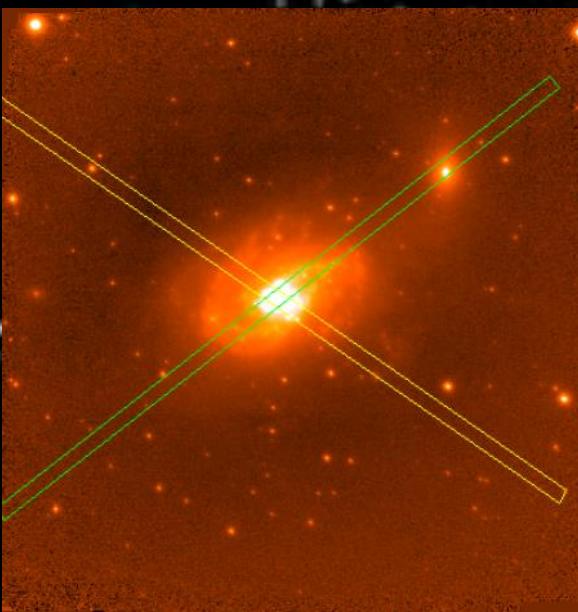
# IRAS18293-3413 SALT



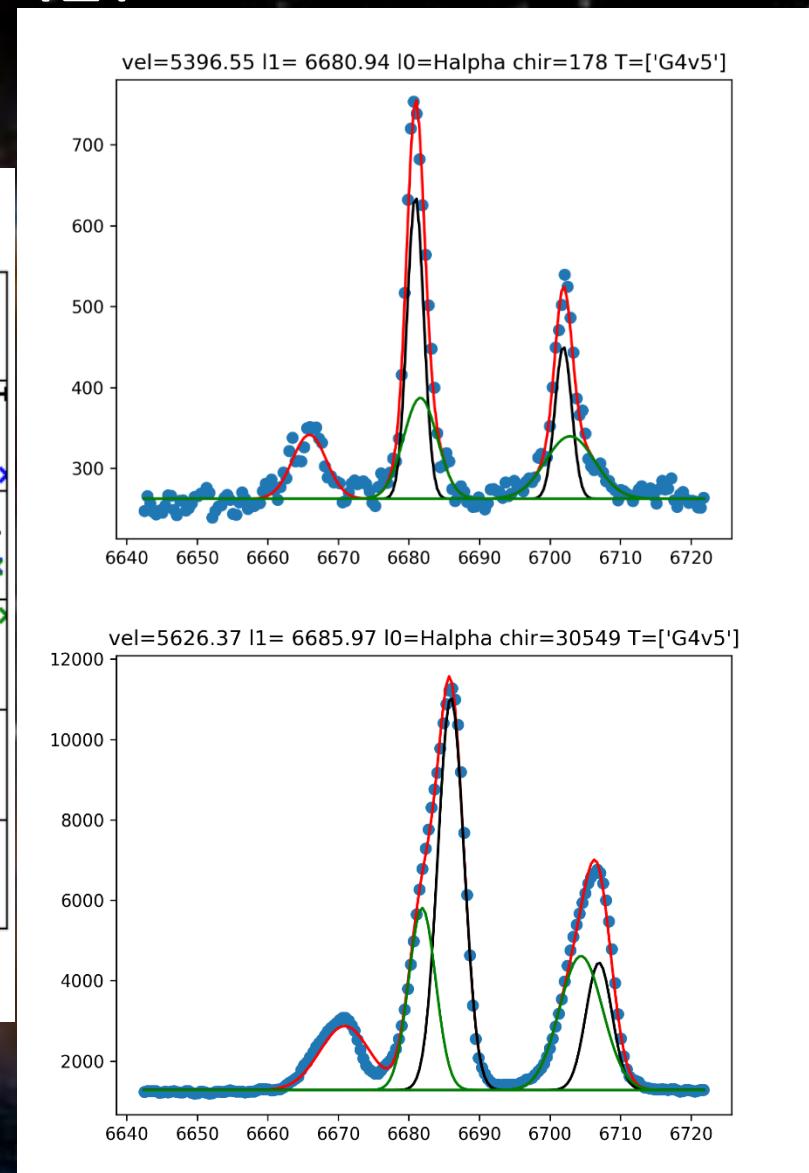
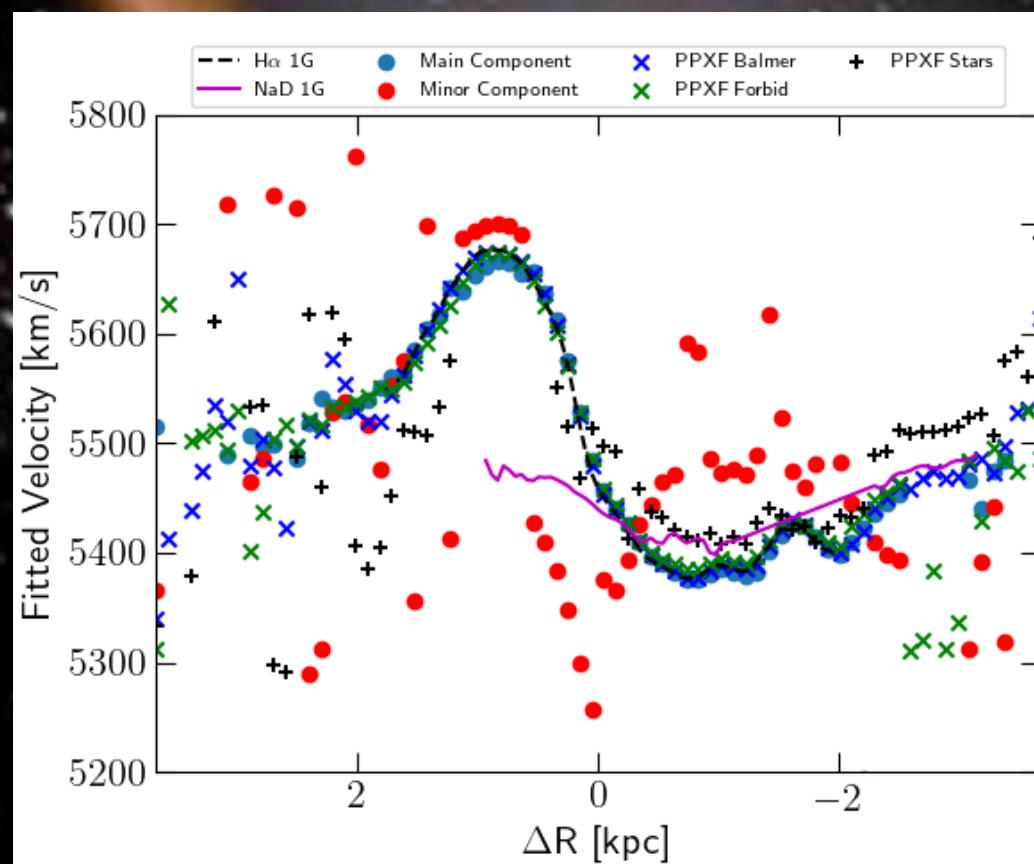
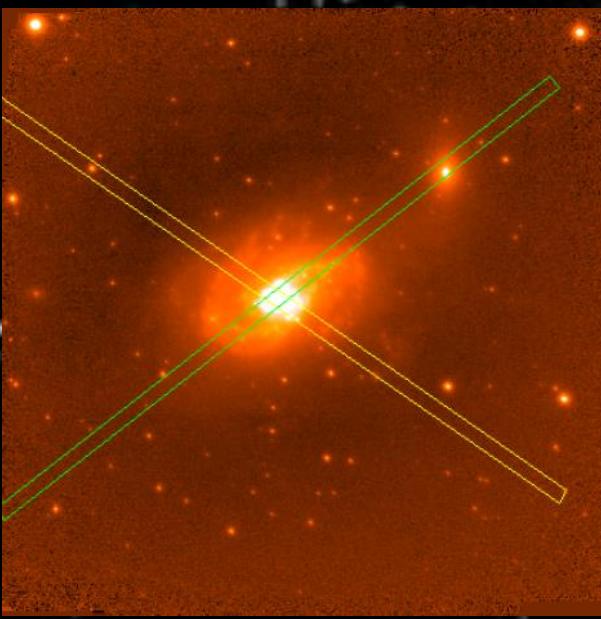
# IRAS18293-3413 SALT

PA=128

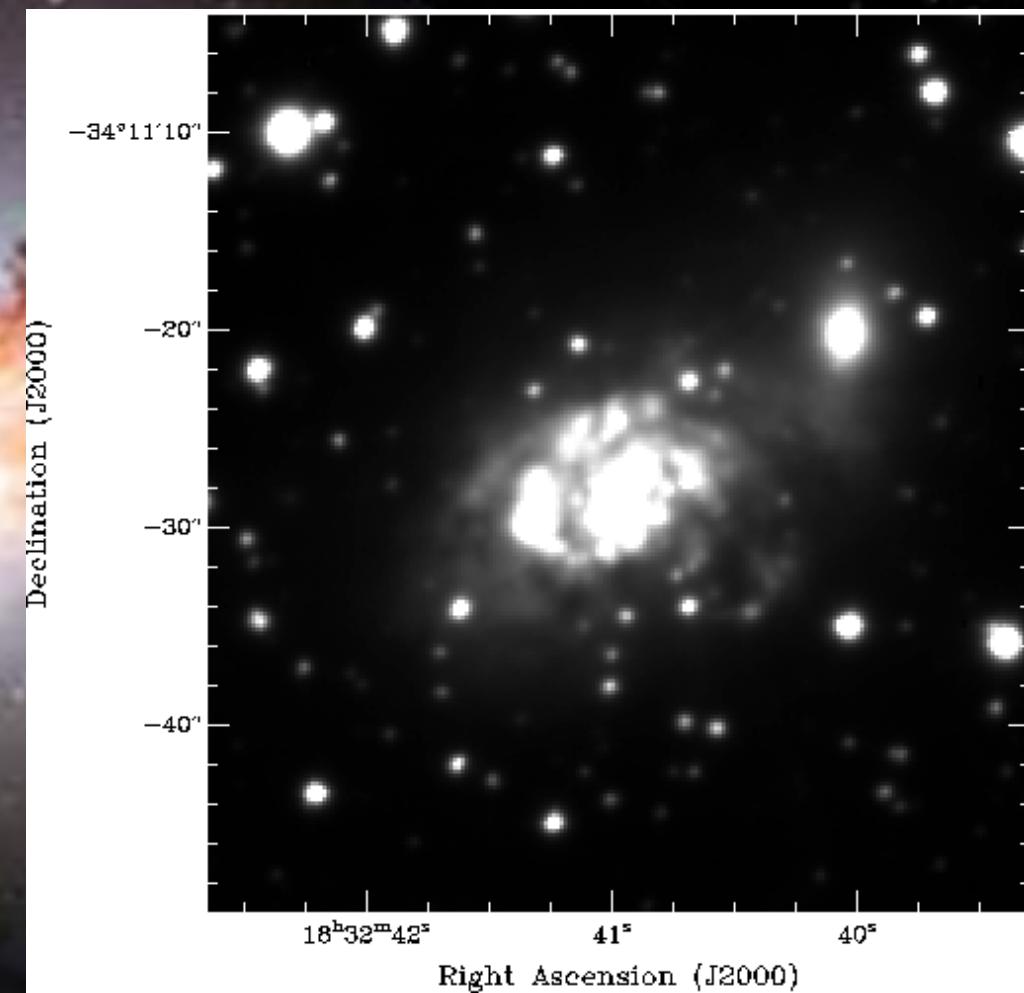
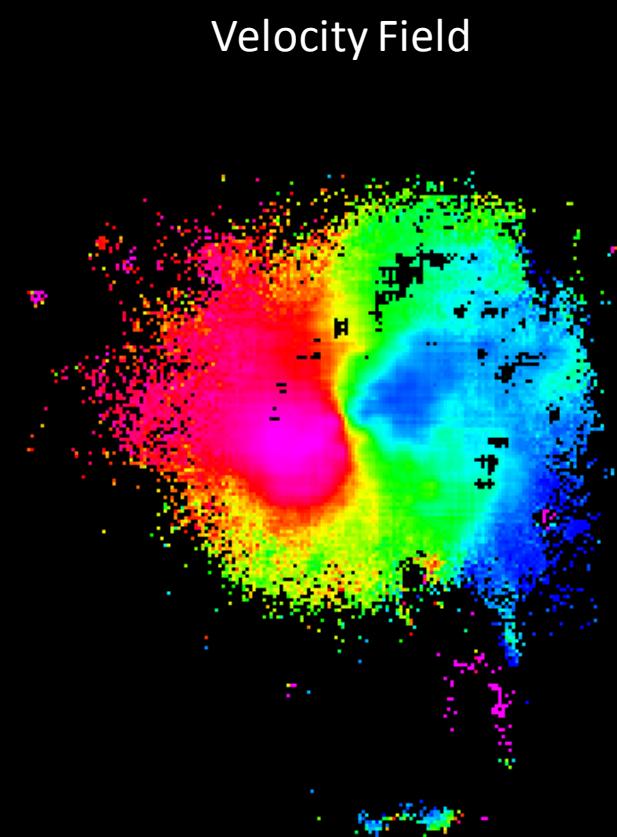
PA=55



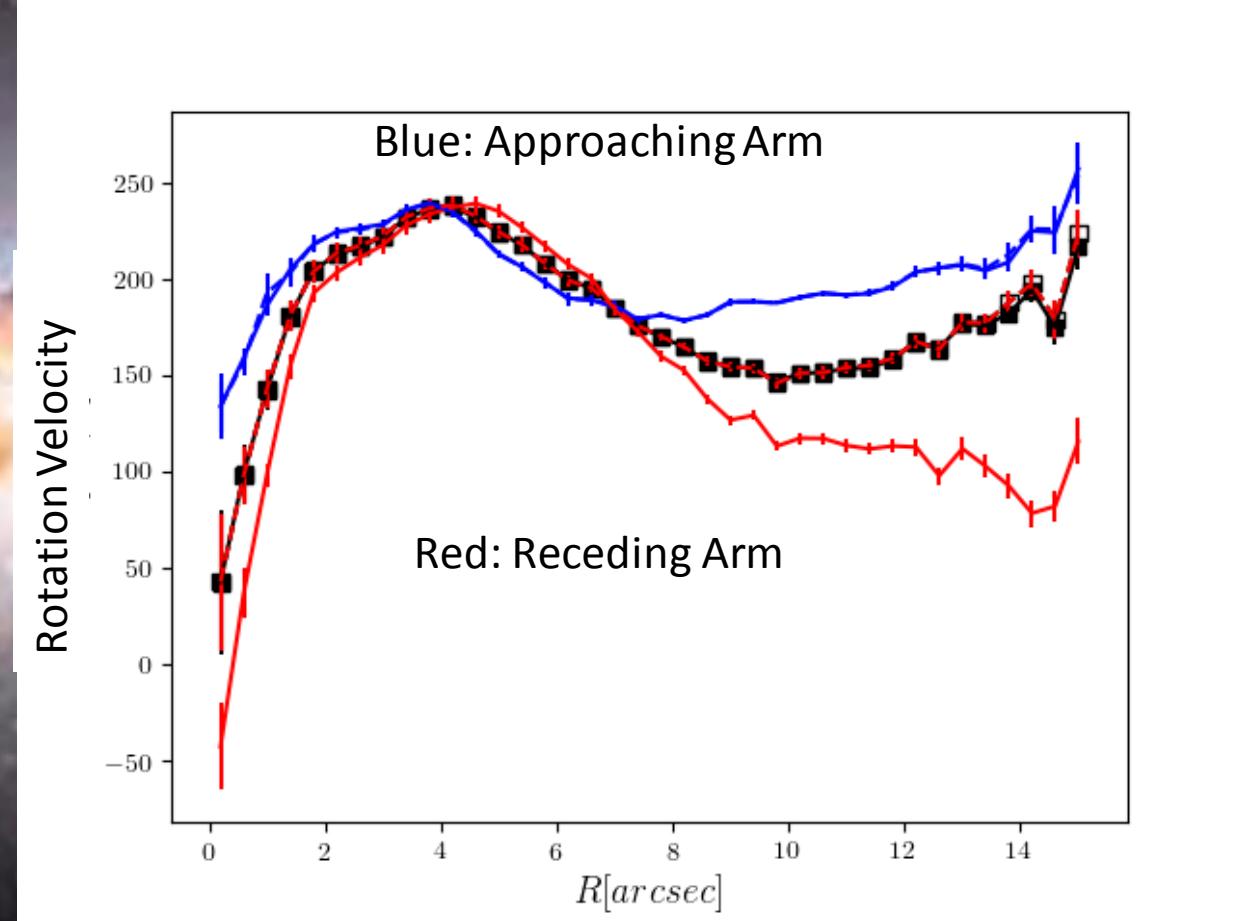
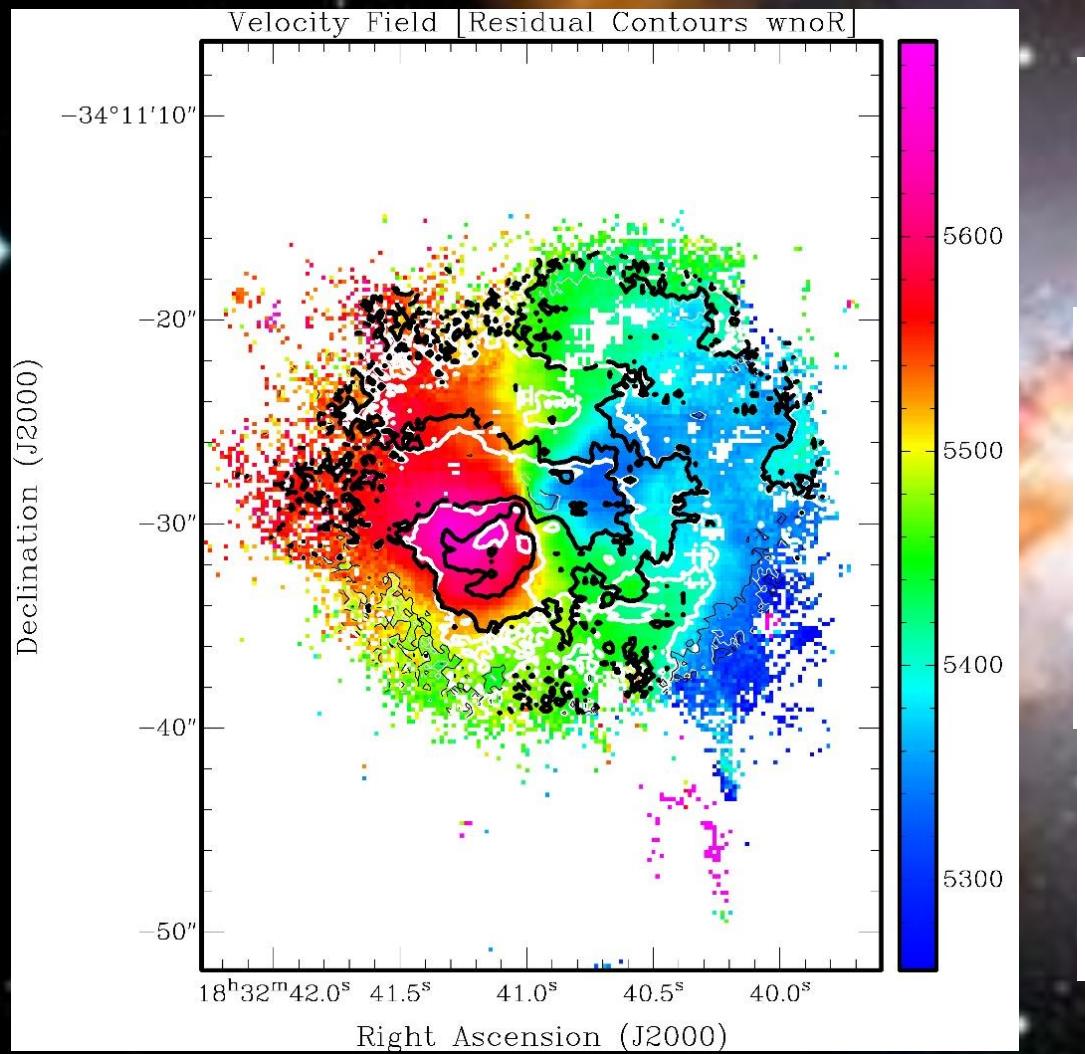
# IRAS18293-3413 SALT



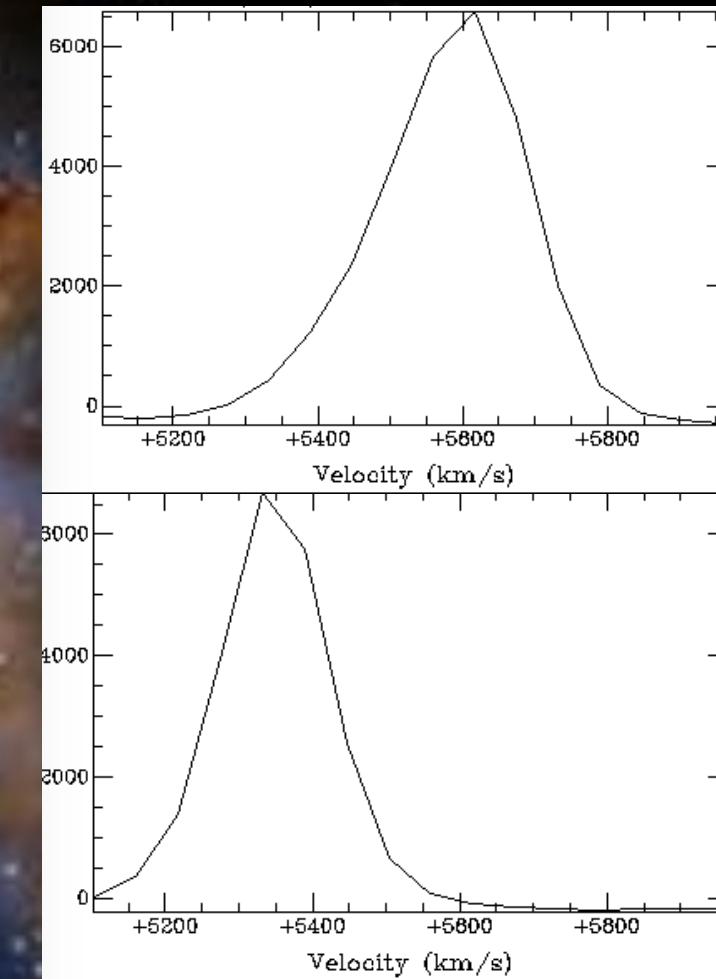
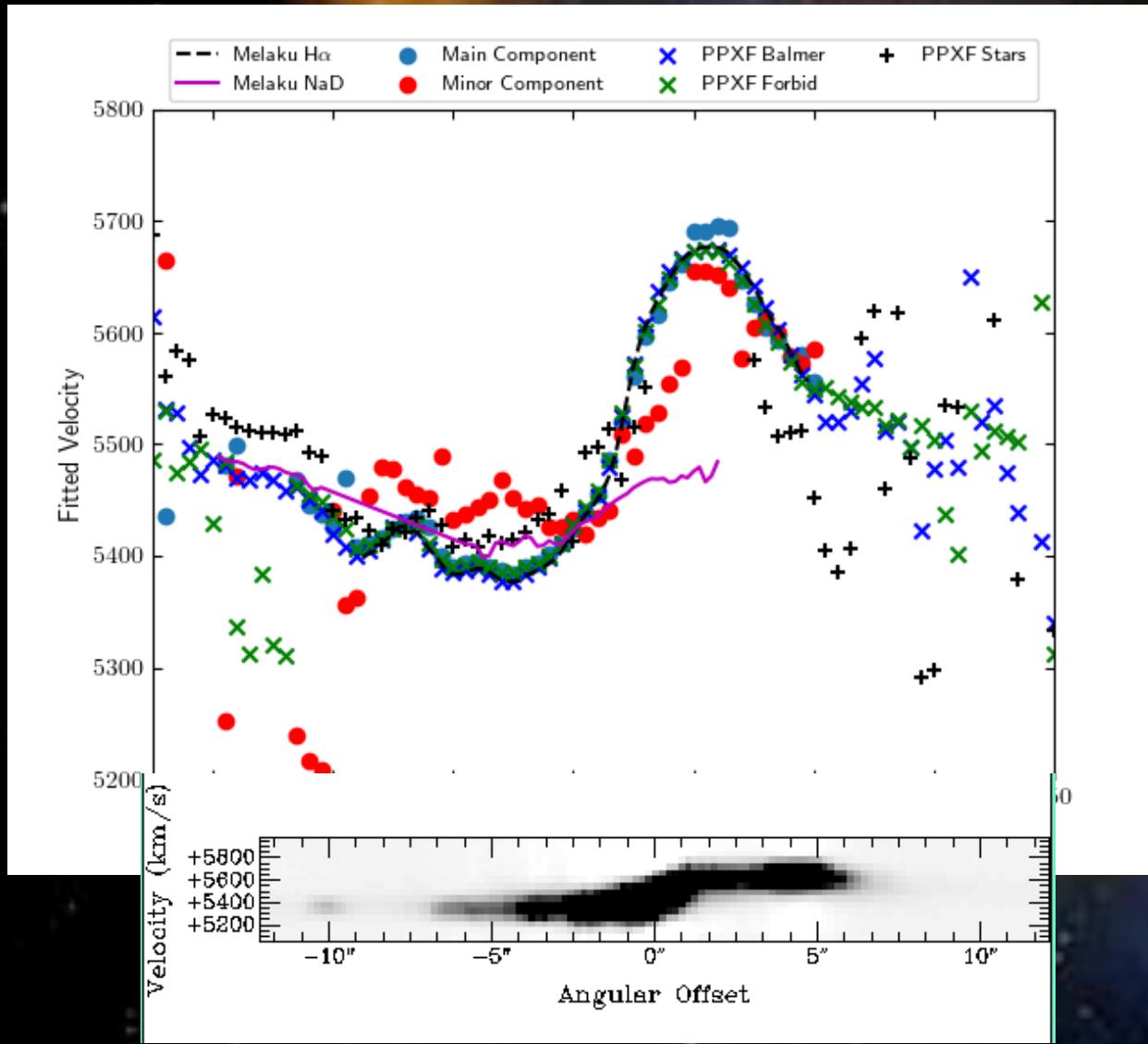
# IRAS18293-3413 MUSE follow-up



# IRAS18293-3413

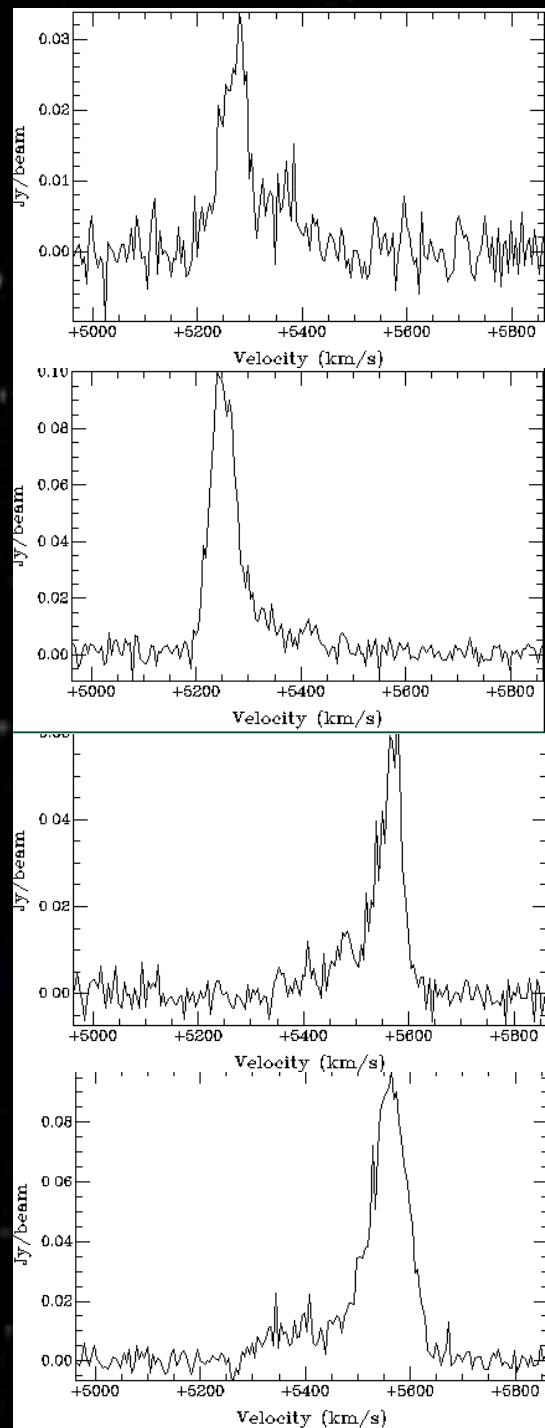
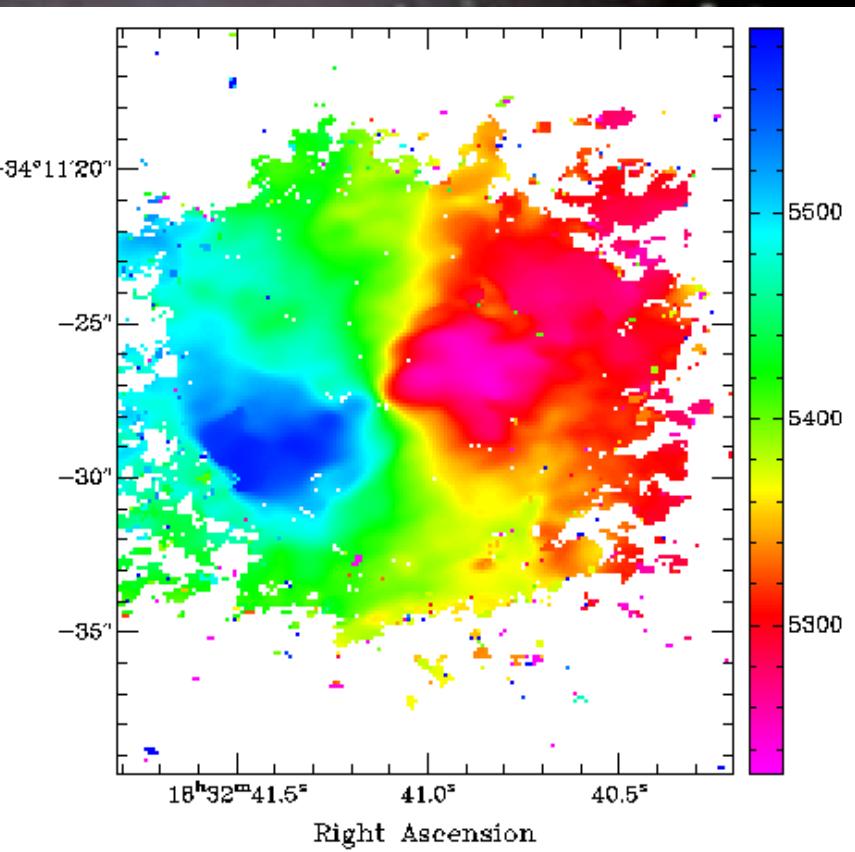
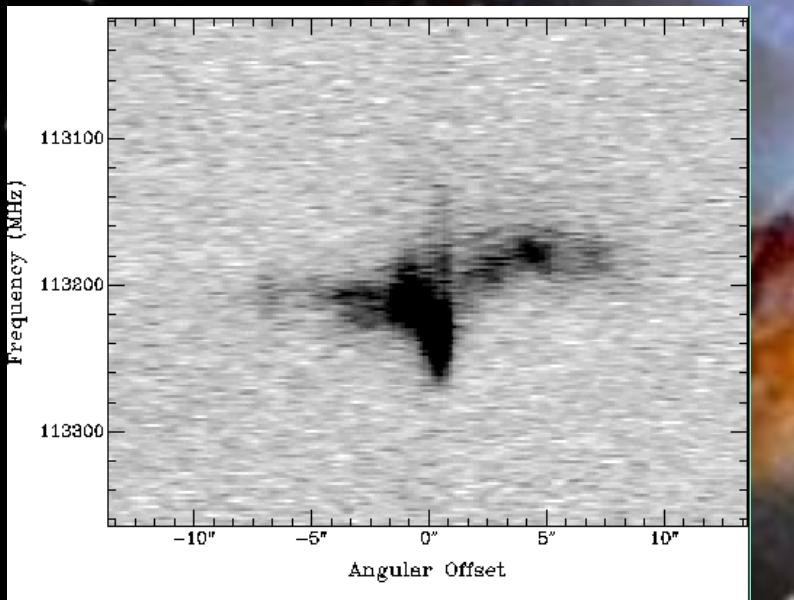


# IRAS18293-3413 SALT + MUSE

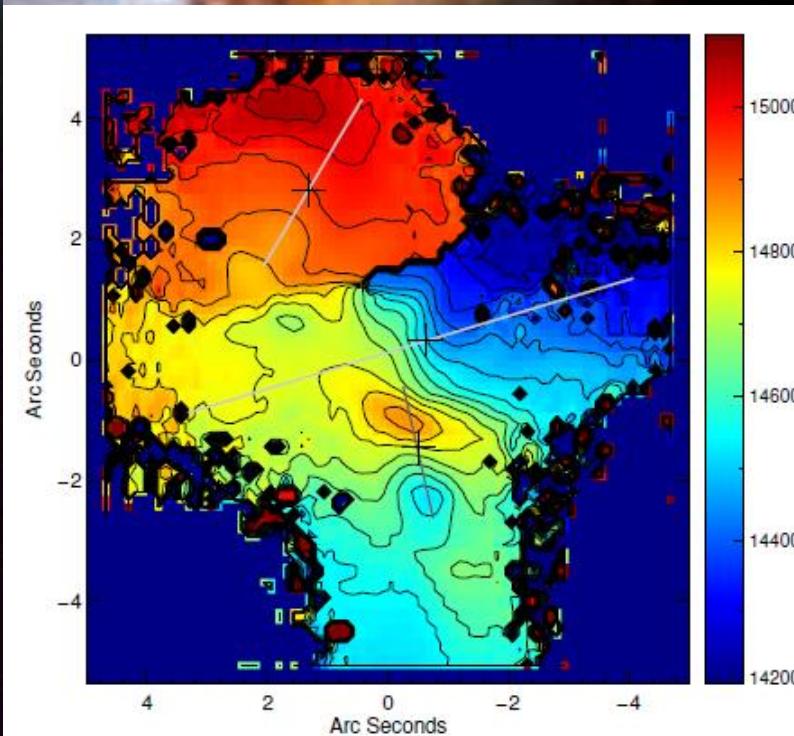
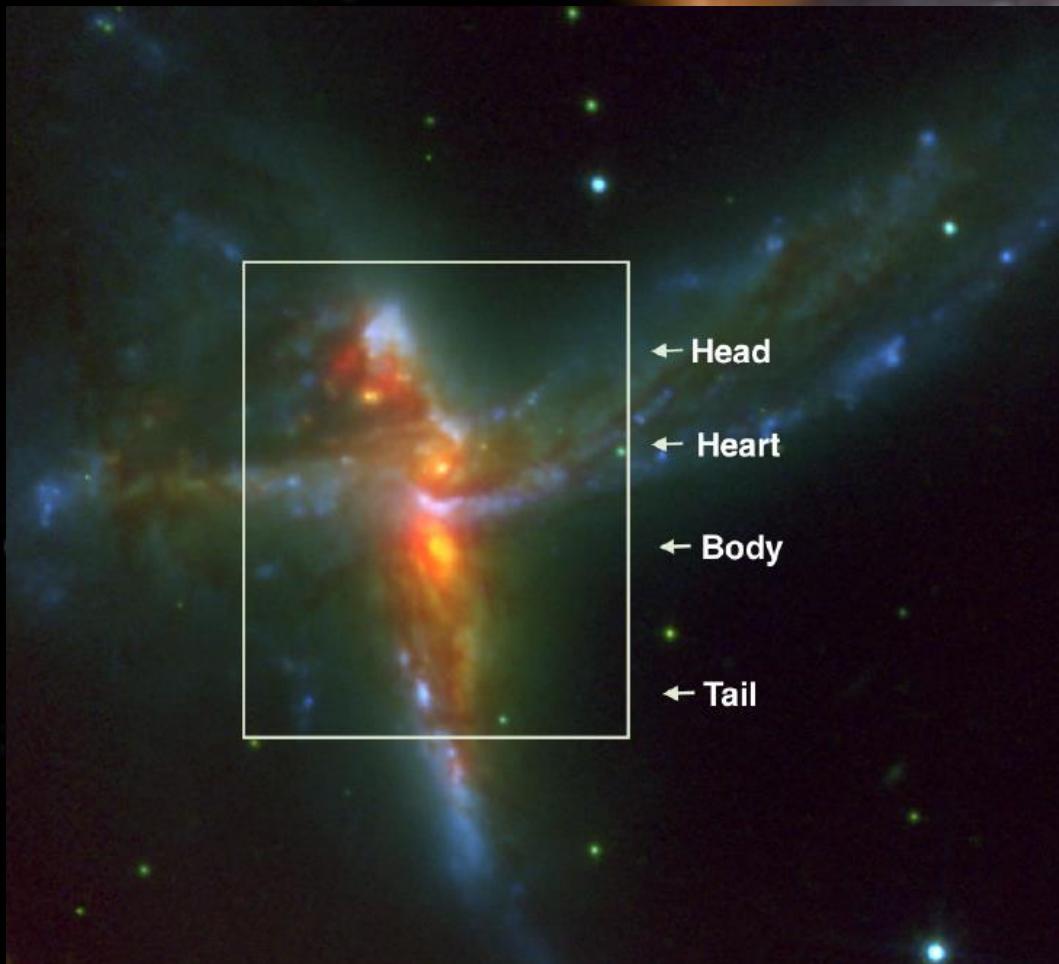


# IRAS18293-3413 CO

Minor Axis velocity excess



# IRAS19115-2124 (aka The Bird)

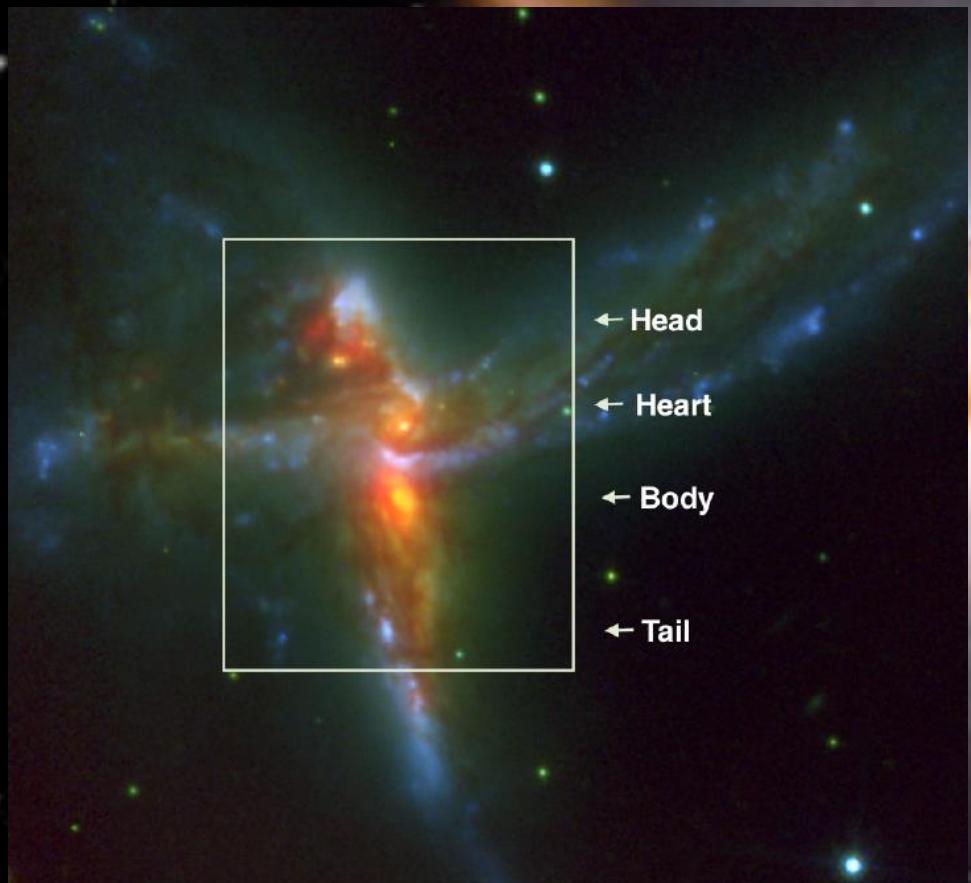


Vaisanen+ 2017

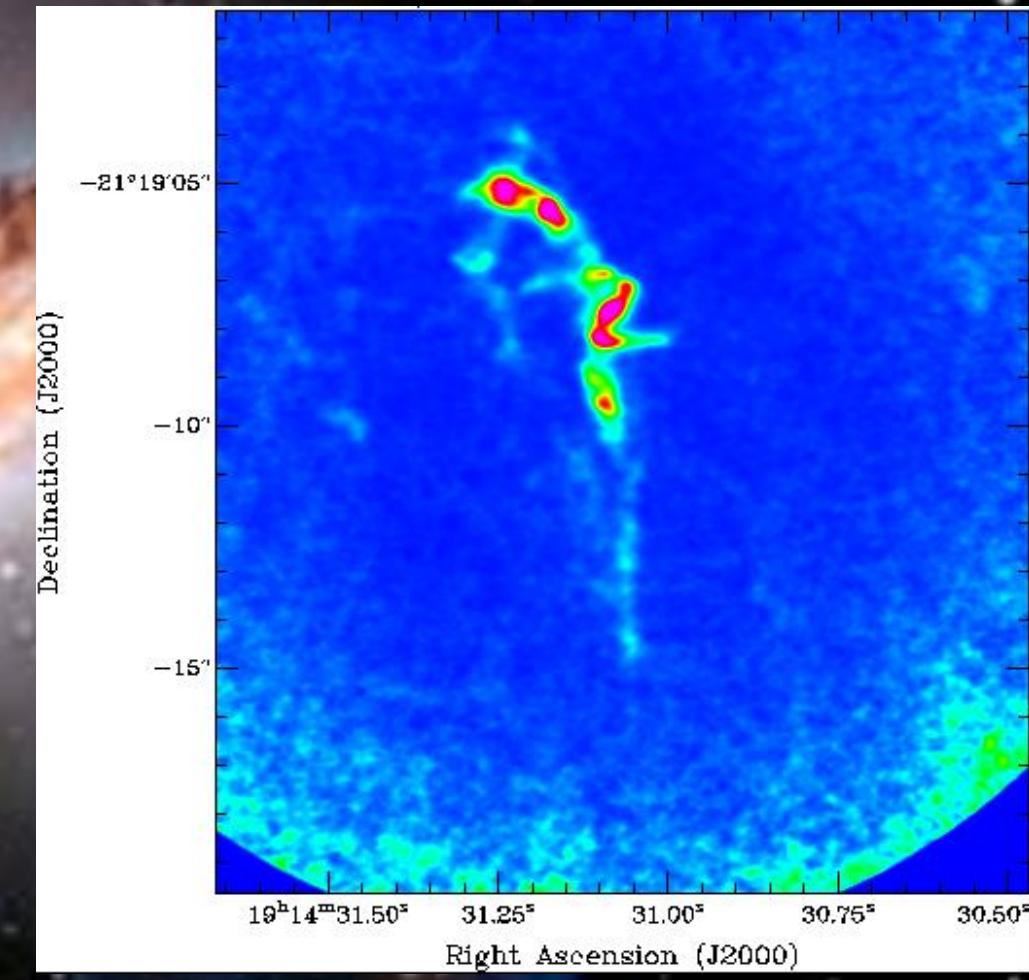
D=200 Mpc  
SFR  $\sim 190 M_{\text{sol}}/\text{yr}$



# The Bird: CO (J=3-2)

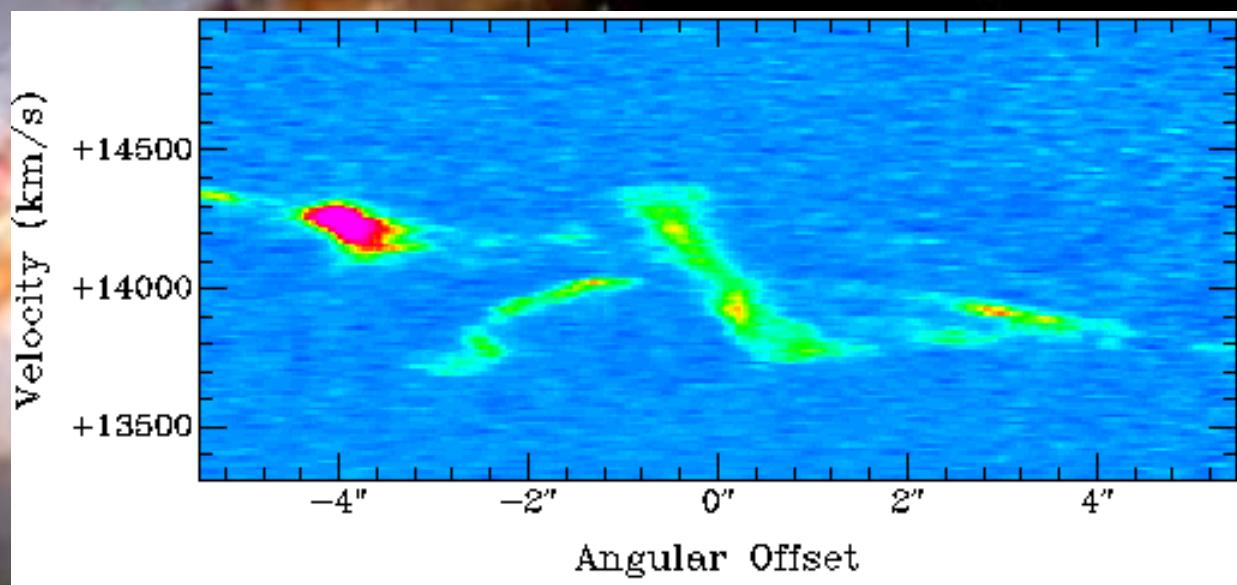
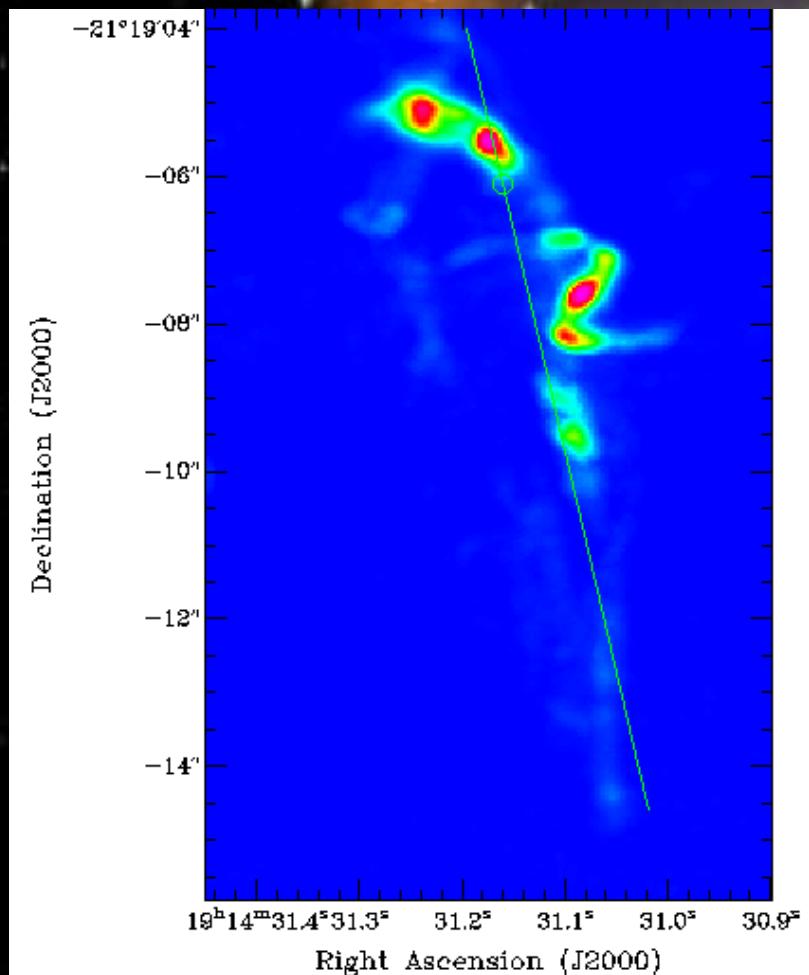


K-Band + HST I & B Bands

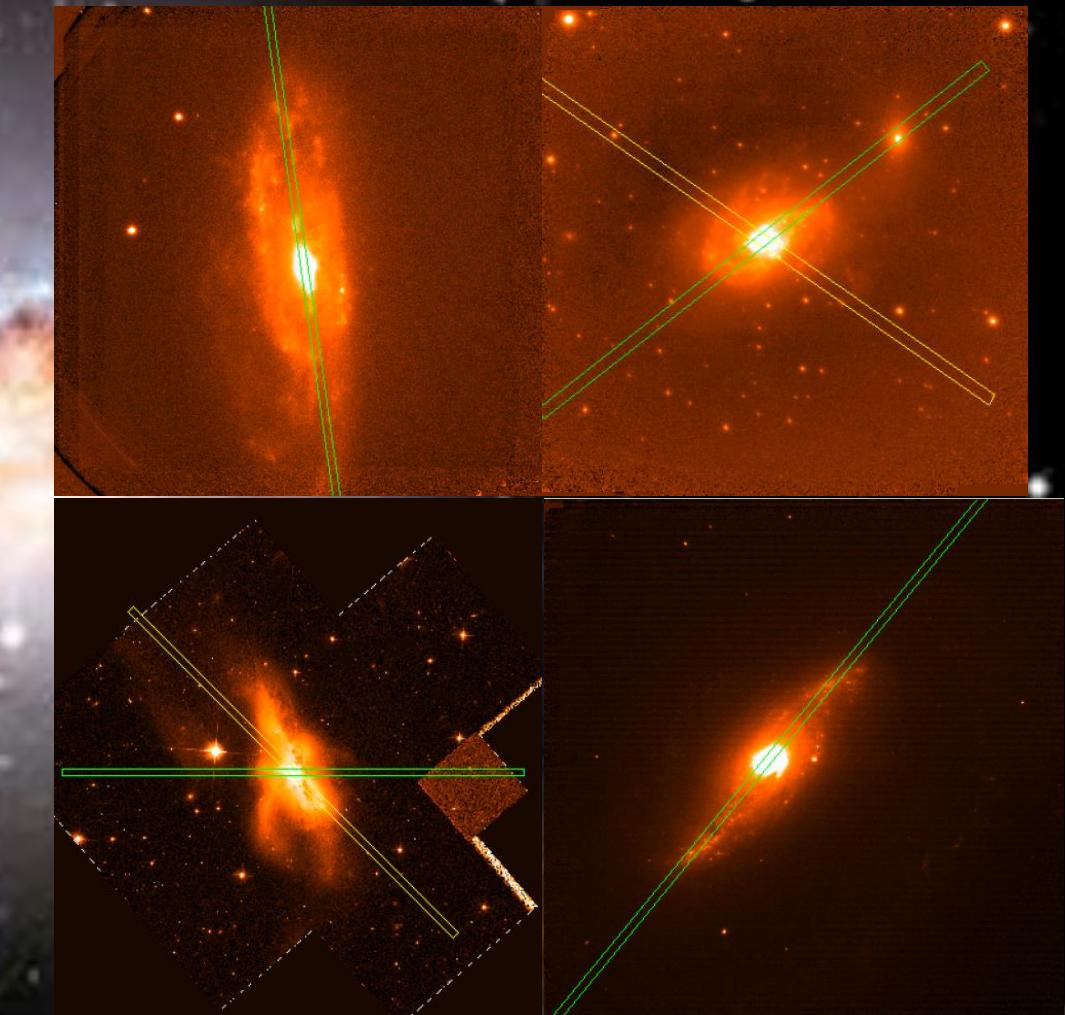


ALMA CO (J=3-2)

# The Bird: CO (J=3-2)

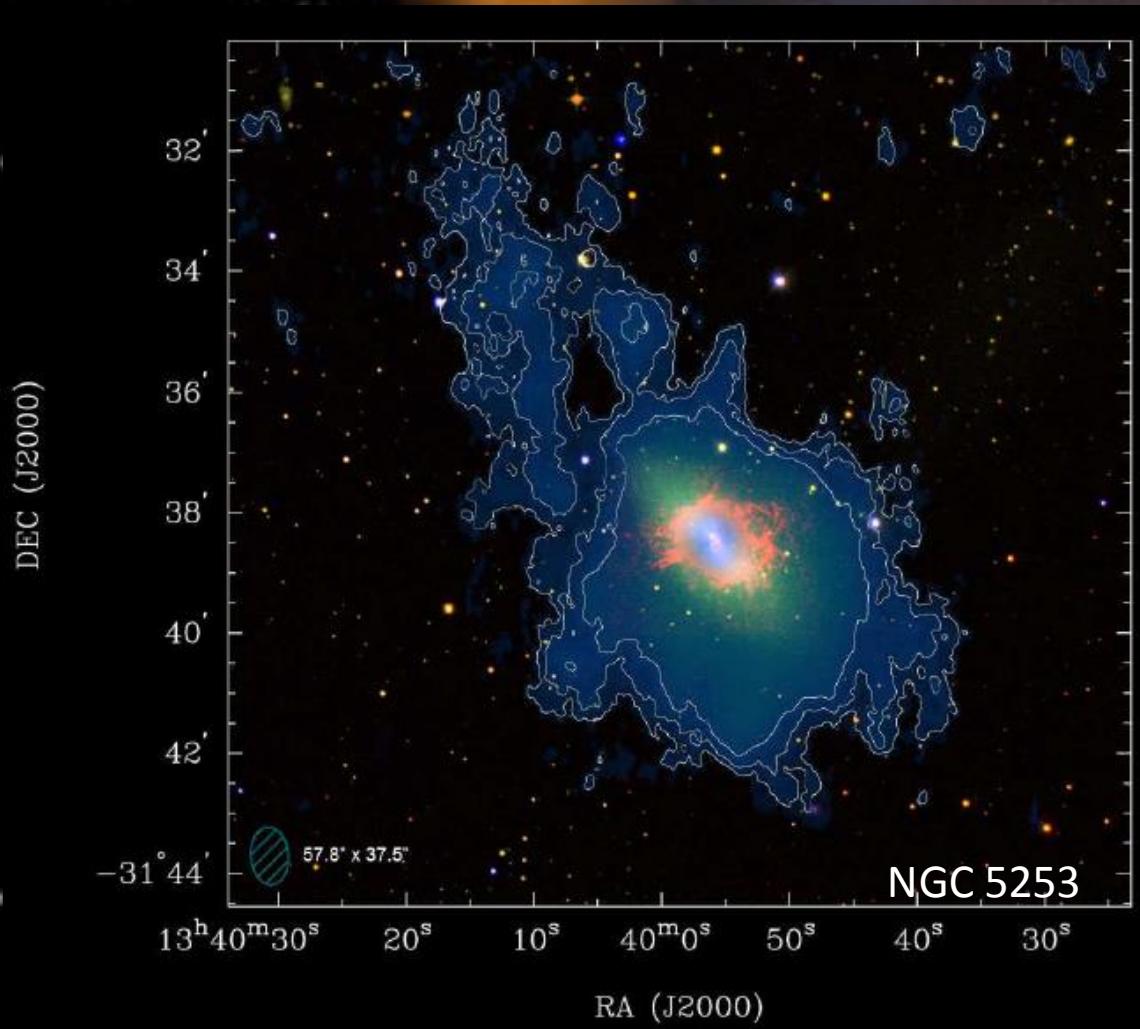


# HI Future... environments



de Blok+ 2018

# HI Future: Accretion versus Outflow



Lopez- Sanchez + 2011

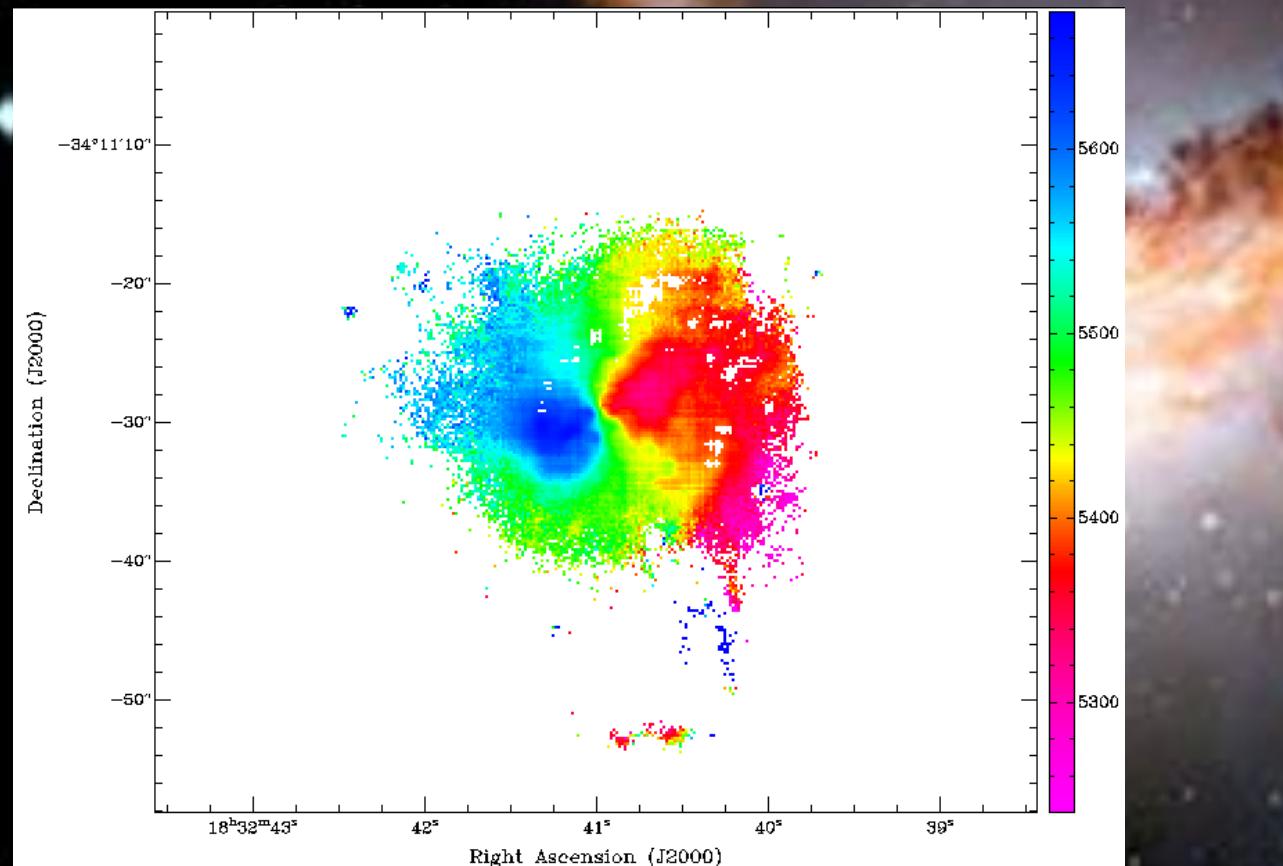


MeerKAT

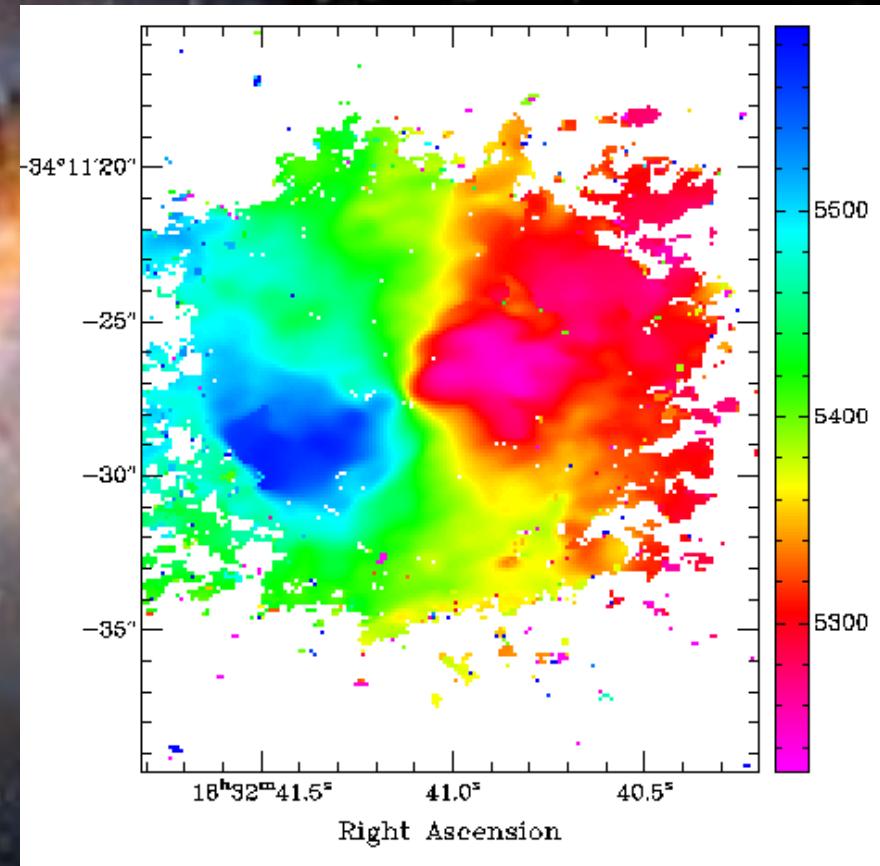


# IRAS18293-3413 MUSE + ALMA

MUSE H $\alpha$  Velocity Field



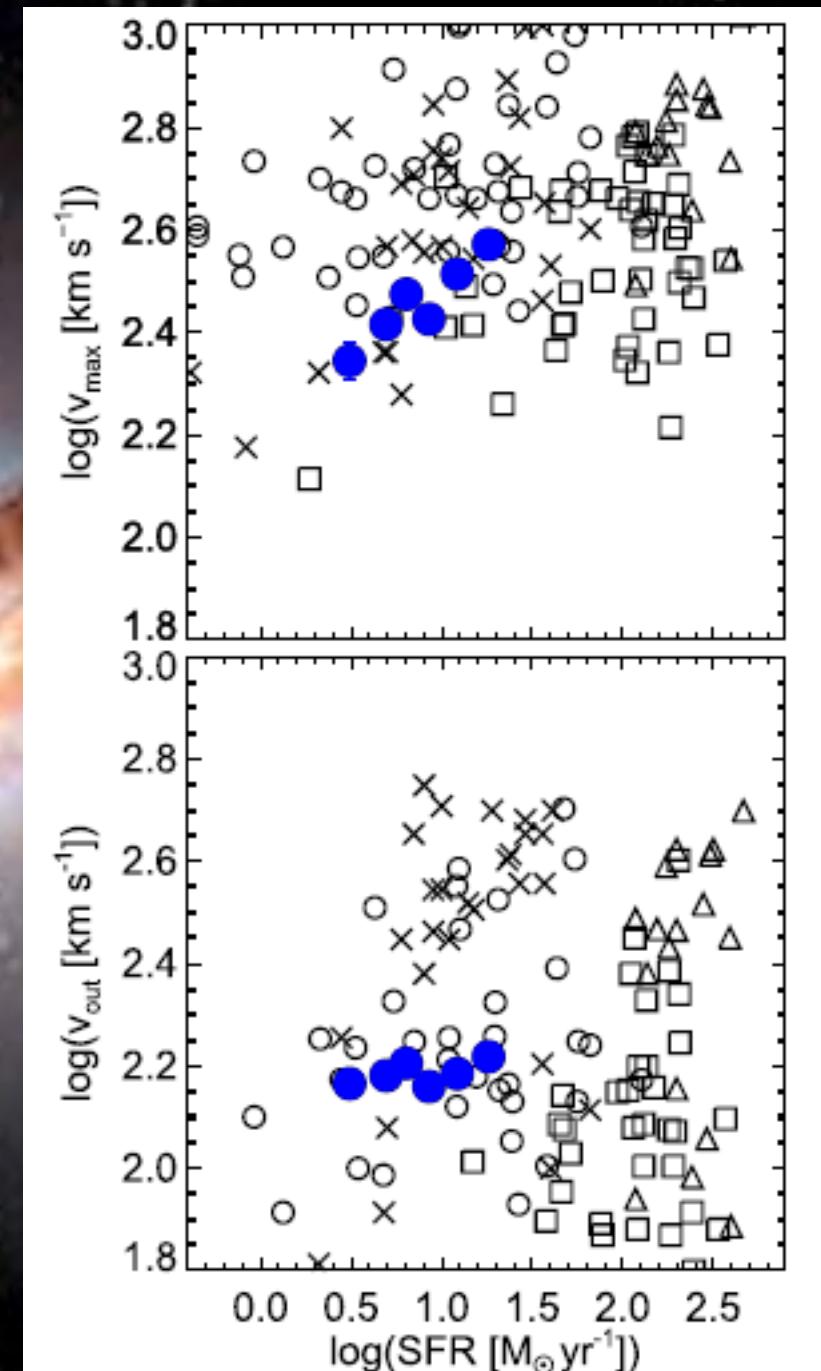
ALMA CO(J=1-0) Velocity Field



# Outflow Measurements

- Maximum Outflow Velocity :  
Max velocity at 90% of flux
  - Correlation
- Bulk Outflow Velocity : Mean  
velocity of outflow component
  - No Correlation

Sugahara+ 2017



# IRAS18293-3413

