

Gravitational Instability Theory

While it is hard, or perhaps even unfeasible, to encrypt a fully self-consistent analytical description of the formation of structure in the Universe from the pristine initial conditions of the random stochastic density field, we may identify various crucial evolutionary stages and seek to formulate appropriate approximations illuminating the essential aspects of these stages. Thus, we will follow the following action plan:

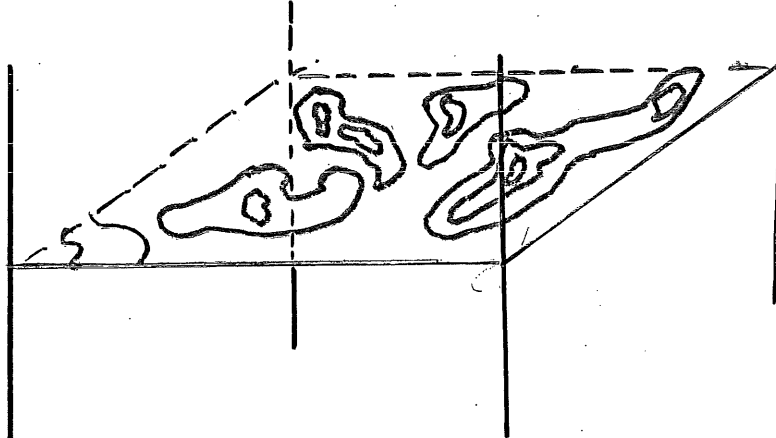
PLAN OF ACTION

STARTING CONDITIONS

- * a (stochastic) Gaussian Primordial Density Field.

fully characterized by:

Power Spectrum $P(k)$

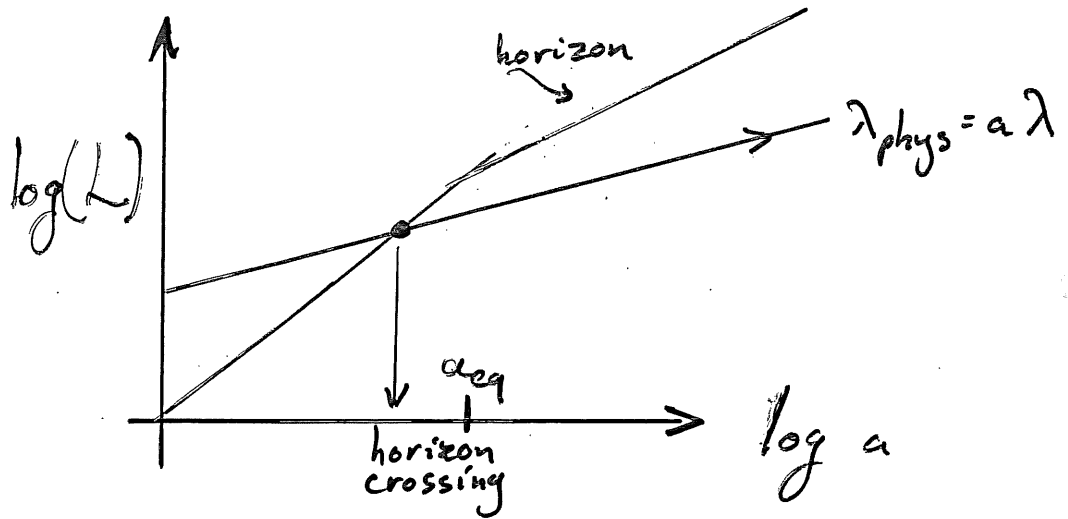


included:

- statistical description Gaussian density field
- mass perturbations
- protostructures & filtering

Stage 0

* From Superhorizon → Subhorizon



Stage 1

LINEAR PHASE:

Density fluctuations still small: $\delta \ll 1$.

① Pressureless "Dark Matter"

• Basic Linear Theory of Perturbation Growth

Basic description of perturbation evolution in early stages of Cold Dark Matter scenarios (current "paradigm").

CDM supposedly only interacts via gravity and weak force/interaction.

- CDM is the dominant factor (gravitationally) in the total mass budget.

$$\underline{\text{WMAP}}: \left. \begin{array}{l} \Omega_{\text{CDM}} \approx 0.27 \\ \Omega_{\text{baryon}} \approx 0.04 \end{array} \right\} \frac{\Omega_{\text{CDM}}}{\Omega_{\text{b}}} \approx 7$$

- \Rightarrow Basic understanding of influence Ω_c on evolution cosmic structure.

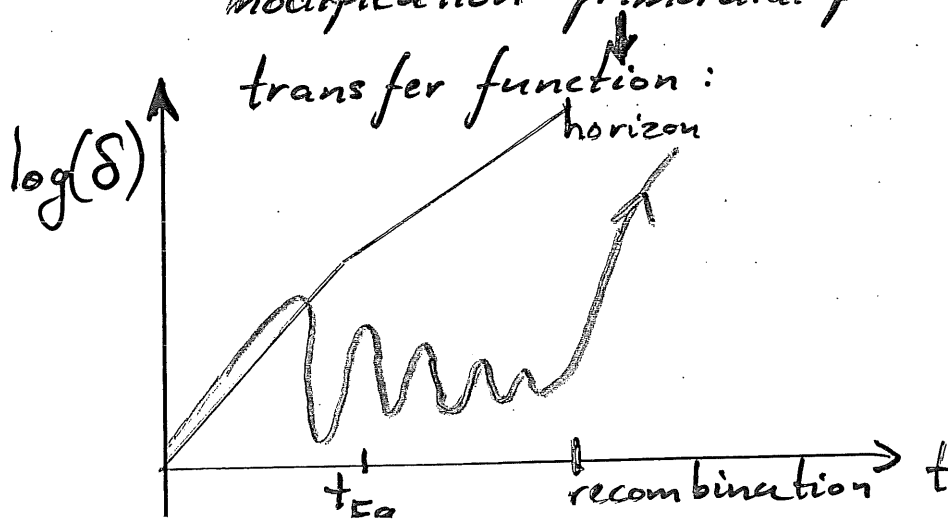
② Jeans Instability Theory

\downarrow
Linear fluctuation theory for a generic medium with pressure

- as when baryons dominate matter budget
- before recombination, when photon pressure plays an extremely important role:

\Rightarrow Essential for Analyzing CMB fluctuations.

- important, to understand spectrum of density fluctuations:
modification primordial power spectrum



included:

Physical Mass Scales:

- Jeans Mass
- Photon Diffusion: Silk Mass
- Meszaros Effect. (CDM)

Stage 2

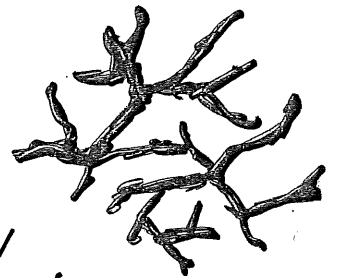
QUASI-LINEAR PHASE

- fluctuations start to decouple from Hubble expansion
- fluctuations change physical shape.
- imprint foamlike matter distribution.

Displacement particles (not yet crossing paths):

Linear Lagrangian formalism:

Zel'dovich approximation



- important tool to understand the generic properties of spatial pattern mass distribution.
- crucial for setting initial conditions N-body simulations

Stage 3

FULLY NONLINEAR EVOLUTION

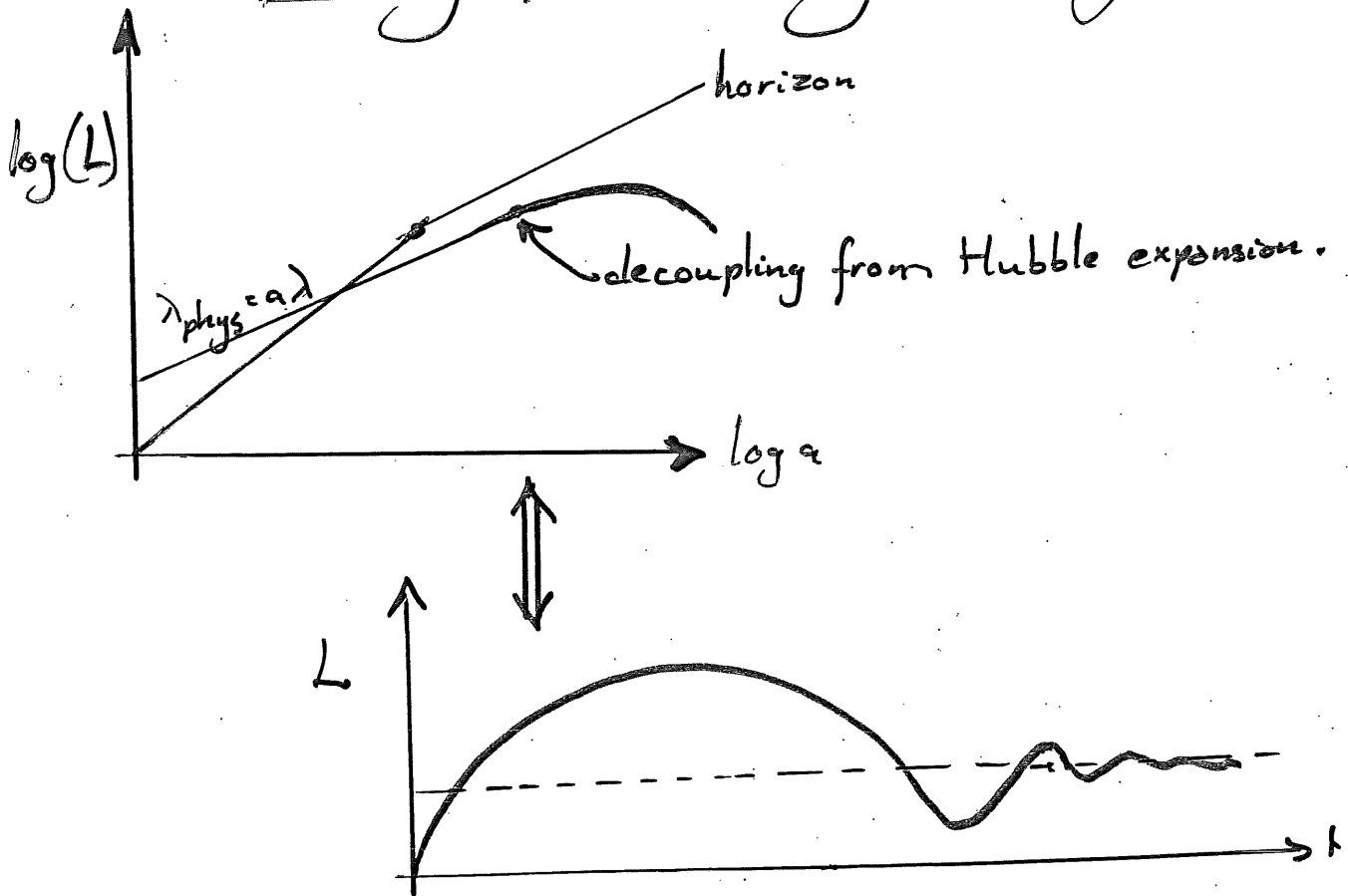
fully selfconsistent analytical treatment not possible.

- Nonlinear Models (approximation)

Models for highly symmetric configurations, used to understand generic behaviour and interpret more generic complex situations:

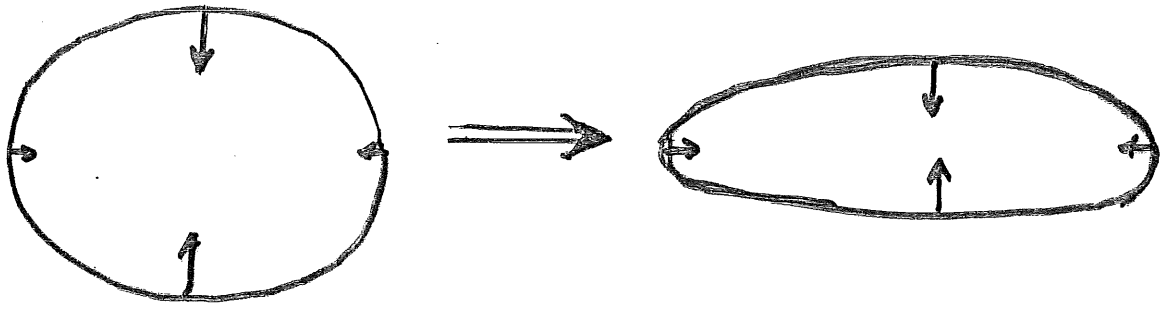
- ① Spherical Model

Template for generic collapse history of overdensity into "object":



- ② Ellipsoidal Model

Template for modified collapse history
shape evolution
of generic nonspherical density
perturbation

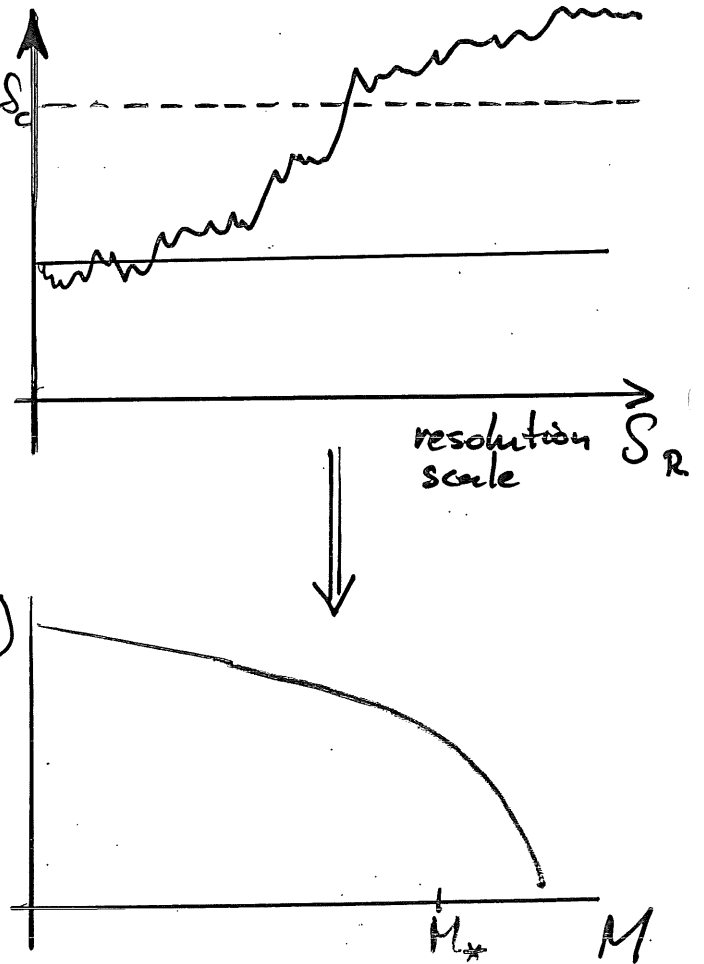


Fully Self-consistent and Complete Evolution
N-body Simulations

Hierarchical Structure Buildup:

③ Press-Schechter / Excursion Set formalism

- Approximate Description S_R for Hierarchical Structure Formation Scenarios
- simplest form based on spherical model.
(in principle suited to invoke further complexities).



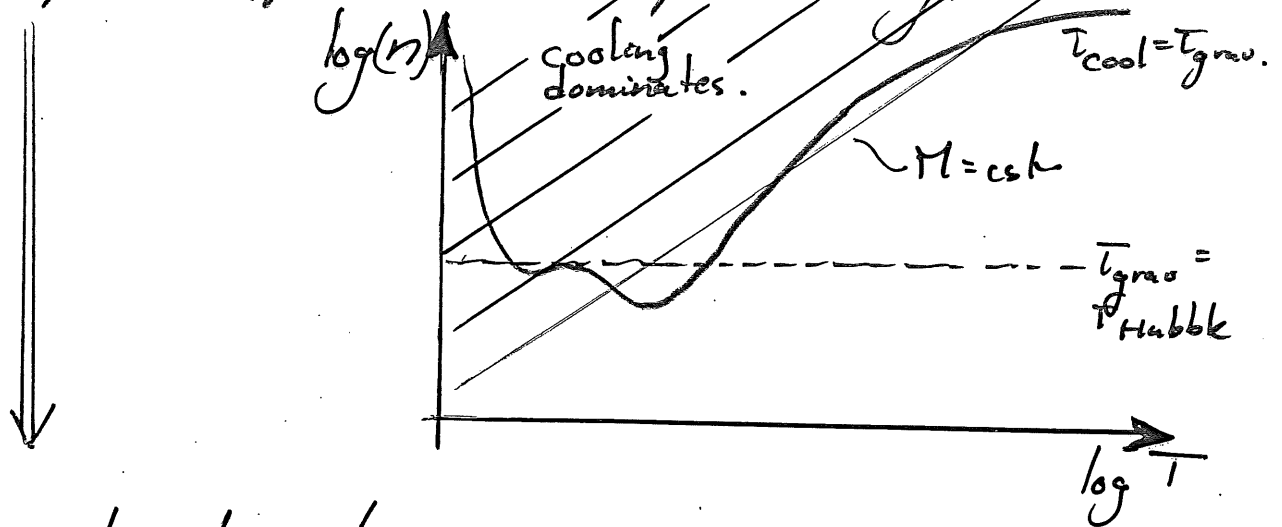
predictions influence cosmology on condensations nonlinear objects.
e.g. cluster mass spectrum

Stage 4

DISSIPATIONAL phase

Dissipational :

- as gas follows into nonlinear potential wells, dissipative effects kick in prominently.



energy loss by cooling processes \rightarrow
reduction of pressure \rightarrow
infall and further compression baryons.

Crucial for formation of galaxies & stars

highly complex interplay of various physical processes:
radiative processes
thermodynamic processes
star formation,
stellar feedback processes
etc.

! Still largely un-understood !