CAPUT DARK ENERGY TOPICS, 2013

1. The Cosmological Constant

- The acceleration of the Universe may be ascribed to the cosmological constant. That is, just that, the cosmological constant as curvature term in the Einstein field equation and not a form of dark energy. Provide a critical discussion of this point of view, both including arguments pro and con.
- Literature suggestions:
 - Kolb R., Bianchi E., Rovelli C., 2010
 Is dark energy really a mystery ?
 Nature, 466, 321 (July 2010)
 - Padmanabhan T., 2003
 Cosmological Constant Weight of the Vacuum arXiv:hep-th/0212290
 - Peebles P.J.E., Ratra B., 2003
 The cosmological constant and dark energy Rev. Mod. Phys., 75, 559
 - Triay R., 2010 Dark energy: fiction or reality ? arXiv: 1004.0091

2. Supernovae, Dark Energy and the Accelerating Universe

- With the realization that Supernovae Ia can be used as standard candles, in 1998 supernovae surveys came to the startling conclusion that our Universe is accelerating. This is commonly seen as the strongest support for the existence of dark energy. Provide a review of supernova cosmology and the prospects for using future supernova surveys for measuring the equation of state, its possible evolution, and determining the nature of dark energy.
- Literature suggestions:
 - Leibundgut B., 2001
 Cosmological Implications from Observations of Type Ia Supernovae
 Annual Review Astronomy Astrophysics, 39, 67-98
 - Frieman J.A., Turner M.S., Huterer D., 2008
 Dark Energy and the Accelerating Universe Annual Review Astron. Astrophys., 46, 385
 - Astier P., 2012
 The expansion of the universe observed with supernovae Physics Reports
 - Perlmutter S., Schmidt B., Riess A., 2012,
 Nobel Lectures
 Reviews of Modern Physics, Vol. 84 (July-Sept. 2012)

3. Baryonic Acoustic Oscillations

- Baryonic oscillations in the power spectrum of galaxy clustering have recently been detected from the 2dFGRS and SDSS galaxy redshift survey. At the moment various deep redshift surveys are being designed to measure the baryonic oscillations and get a precise measurement of the equation of state of dark energy from this purely geometric and robust method. Give a discussion of the fundamental physics behind the oscillations, on the detection and measurement of the oscillations and on the prospects to study dark energy.
- Literature suggestions:
 - Eisenstein D.J. and C.L. Bennett *Cosmic sound waves rule* Physics Today, april 2008, 44
 - Colless M.
 The WiggleZ Dark Energy Survey
 KIAS workshop 2008, ppt (on request from RvdW).
 - Eisenstein D.J., Hu W., 1997
 Baryonic Features in the Matter Transfer Function
 ApJ 496, 605
 - Seo H-J, Eisenstein D.J., 2005
 Baryon Acoustic Oscillations in Simulated Galaxy Redshift Surveys
 ApJ 633, 575
 - Seo H-J, Eisenstein D.J., 2005
 Probing Dark Energy with Baryonic Acoustic Oscillations ASP, Vol. 399
 - White M., 2005 Baryon Oscillations Astroparticle Physics 24, 334
 - Bassett B.A., Hlozek R., 2009 Baryon Acoustic Oscillations arXiv:0910.5224

4. Gravitational Lensing and Dark Energy Measurement

- One of the main probes for the accurate measurement of dark energy parameters is gravitational weak lensing. Provide a thorough discussion of the theoretical basis of this, as well as the observational and practical challenges involved and the prospects in the light of upcoming surveys suchs as KIDS, DES, Euclid and LSST.
- Literature Suggestions
 - ESA Euclid, Mapping the geometry of the dark Universe ESA/SRE(2011)12, July 2011
 - Heavens A., 2011 Cosmology with gravitational lensing arXiv:1109.1121
 - Hoekstra H., Jain B., 2008
 Weak Gravitational Lensing and Its Cosmological Applications
 Ann. Rev. Nucl. Part. Physics, 58, 99
 - Huterer D., 2010
 Weak lensing, dark matter and dark energy arXiv:1001.1758
 - Laszlo I., Bean R., Kirk D., Bridle S. Disentangling dark energy and cosmic tests of gravity from weak lensing systematics
 - Mellier Y., 1999
 Probing the Universe with Weak Lensing
 Ann. Rev. Astron. Astrophys. 37, 127-189
 - van Waerbeke L., Mellier Y., 2003
 Gravitational Lensing by Large Scale Structures: A Review arXiv:astro-ph/0305089

5. Redshift Space Distortions, Growth Rate and Dark Energy

- E. Linder pointed out that it should be possible to determine the cosmic structure growth rate history by looking at the redshift space distortion in the clustering of galaxies. Deep galaxy redshift surveys may thus provide strong constraints on the nature of dark energy. Provide a discussion of the fundamentals of the method, and subsequently provide a critical assessement of the observational projects that have been carried out and will be carried out along these lines.
- Literature suggestions:
 - Guzzo L., et al., 2008
 A test of the nature of cosmic acceleration using galaxy redshift distortions Nature, 451, Jan. 2008
 - Huterer D., Kirkby D., 2013
 Growth of Cosmic Structure: Probing Dark Energy Beyond Expansion arXiv:1309.5385
 - Linder E.V., 2005
 Cosmic growth history and expansion history Phys. Rev. D., 72, 043529
 - Linder E., 2007
 Redshift Distortions as a Probe of Gravity arXiv:0709.1113
 - Majerotto E., Guzzo L., et al., 2012
 Probing deviations from General Relativity with the Euclid spectroscopic survey MNRAS, 2012, arXiv:1205.6215

6. Integrated Sachs-Wolfe effect and Dark Energy

- The effect of dark energy on the evolution of gravitational potential perturbations is probed by the integrated Sachs-Wolfe effect. To this end, one needs to combine the large scale cosmic microwave background perturbations with the measured mass perturbations along the line of sight. Describe the theoretical context of the ISW, and discuss the complications and results of attempts towards measuring the ISW.
- Literature suggestions:
 - Crittenden R., Turok N., 1996,
 Looking for a Cosmological Constant with the Rees-Sciama Effect Phys. Rev. Lett., 76, 575
 - Giannantonio T., et al., 2008, Combined analysis of the integrated Sachs-Wolfe effect and cosmological implications
 Phys. Rev. D., 77, 123520
 - Giannantonio T., Crittenden R., Nichol R., Ross A.J., 2012, The significance of the integrated Sachs-Wolfe effect revisited MNRAS, 2012, arXiv:1209.2125
 - Hu W., Dodelson S., 2002
 Cosmic Microwave Background Anisotropies
 ARAA, 40, 171-216
 - Hu W., Sugiyama N., Silk J., 1997, *The Physics of the Microwave Background Anisotropies* Nature, Vol. 386, p. 37-43
 - Planck, 2013,
 Planck 2013 Results XIX. The Integrated Sachs-Wolfe effect.
 Astron. Astrophys., arXiv:1303.5079
 - Samtleben D., Staggs S., Winstein B., 2007, The Cosmic Microwave Background for Pedestrians: A Review for Particle and Nuclear Physicists
 Ann. Rev. Nucl. Part. Sci., 57, 245-283
 - Granett B.R., Neyrinck M.C., Szapudi I., 2008
 An imprint of superstructures of the microwave background due to the ntegrated Sachs-Wolfe effect
 Astrophys. J., 683, L99-L102

7. Dark Energy and Voids

- A recent very interesting observation is that voids in the mass distribution are highly sensitive probes of dark energy. Investigate the dynamical origin of this sensitivity, review recent studies on this, and discuss the prospects of using this sensitivity for probing dark energy in upcoming large surveys.
- Literature suggestions:
 - Bos E.G.P., van de Weygaert R., Dolag K., Pettorino V., 2012 The darkness that shaped the void: dark energy and cosmic voids MNRAS, 426, 440
 - Lavaux G., Wandelt B.D., 2010 2
 Precision cosmology with voids: definition, methods, dynamics MNRAS, 403, 1392
 - Lavaux G., Wandelt B.D., 2012
 Precision Cosmography with Stacked Voids
 Astrophys.J., 754, 109
 - Lee J., Park D., 2009
 Constraining the Dark Energy Equation of State with Cosmic Voids Astrophys. J., 696, 10
 - van de Weygaert R., Platen E., 2011
 Cosmic Voids: Structure, Dynamics and Galaxies
 IJMPS, 1, 41, arXiv:0912.2997

8. Cosmic Backreaction

- According to some cosmologists, the observed acceleration of the Universe is the result of an artefact, resulting from the fact that in a general relativistic description inhomogeneities in the mass distribution cause complex non-linear couplings between the curvature perturbations that will reveal themselves as an artificial acceleration term when assuming a uniform (FRW) background. There has been ample discussion on whether the amplitude of this effect is noticeable or not. Provide a discussion on the theoretical underpinning of this argument, and subsequently a balanced discussion between the different arguments pro and con this explanation.
- Literature suggestions:
 - Buchert T., 2007
 Dark Energy from Structure A Status Report
 Gen. Rel. Grav., arXiv:0707.2153
 - Célérier M-N., 2012
 Effects of inhomogeneities on the expansion of the Universe: a challenge to dark energy arXiv:1203.2814
 - Clarkson C., Ellis G., Larena J., Umeh O., 2011
 Does the growth of structure affect our dynamical models of the universe arXiv:1109.2314
 - Collins H., 2010, The influence of inhomogeneities on the large-scale expansion of the universe arXiv:1011,2946
 - Kolb E.W., Marra V., Matarrese S., 2009
 Cosmological background solutions and cosmological backreactions
 Gen. Rel. Grav., arXiv: 0901.4566
 - Schwarz D.J., 2010 Cosmological Backreaction arXiv:1003.3026
 - Singh T.P., 2011 The Effect of Cosmic Inhomogeneities On The Average Cosmological Dynamics arXiv:1105.3450
 - Wiltshire D.L., 2013
 Cosmic structure, averaging and dark energy arXiv:1311.3787

9. Modified Gravity and Dark Energy

- As a fundamental alternative to dark energy, one can ask whether the acceleration of the Universe is caused by a modification of gravity on large scales, i.e. departure from General Relativity, rather than an exotic form of energy. This possibility has generated a significant amount of theoretical work over the past decade; it furthermore provides strong motivation to search for and constrain modifications to GR using cosmological observations. There have been a range of suggested options for modified gravity models, amongst which f(R) gravity is amongst the best studied. f(R) gravity is a type of modified gravity theory which generalizes Einstein's General Relativity. f(R) gravity is actually a family of theories, each one defined by a different function of the Ricci scalar. Provide a review of possible options for modified gravity, and relate this to the observational evidence for dark energy. Amongst all options, you may want to concentrate on f(R) gravity models.
- Literature suggestions:
 - Carroll S.M., Duvvuri V., Trodden M., Turner M., 2004 Is cosmic speed-up due to new gravitational physics ? Phys. Rev. D. 70, 043528
 - Faraoni V., 2008
 f(R) gravity: successes and challenges
 arXiv:0810.2602
 - Huterer D., Kirkby D., 2013
 Growth of Cosmic Structure: Probing Dark Energy Beyond Expansion arXiv:1309.5385
 - Laszlo I., Bean R., Kirk D., Bridle S. Disentangling dark energy and cosmic tests of gravity from weak lensing systematics
 - Sotiriou T.P., Fararoni V., 2010
 f(R) theories of gravity
 Rev. Mod. Physics, 82, 451, Jan-March 2010
 - Tsujikawa S., 2011 Modified gravity models of dark energy arXiv:1101.0191

10. Dark Energy and the Vacuum

- One of the most immediate suggestions for the nature of dark energy concerns vacuum energy. However, it involves a major problem. Estimates of vacuum energy are a bit off, by 120 orders of magnitude. Discuss this —it Cosmological Constant Problem, the various proposals and options for vacuum energy, and possible solutions to the problem
- Literature suggestions:
 - Weinberg S., 1989
 The Cosmological Constant Problem
 Rev. Mod. Phys., 61, 1-23
 - Maggiore M., 2010 Zero-point quantum fluctuations and the cosmological expansion arXiv:1004.1782
 - Zeldovich Y., 1967
 Cosmological Constant and Elementary Particles
 ZhETF Pis ma Redaktsiiu, Vol. 6, p.883
 - Ziaeepour H., 2012
 Issues about vacuum energy as the origin of dark energy arXiv:1205.3304

11. Quintessence

- One of the first candidates for a dynamically evolving form of dark energy is quintessence. Quintessence differs from the cosmological constant explanation of dark energy in that it is dynamic, changing over time, unlike the cosmological constant which always stays constant. The name comes from the classical elements in ancient Greece. The aether, a pure "fifth element", was thought to fill the Universe beyond Earth. Similarly, modern quintessence would be the fifth known contribution to the overall mass-energy content of the Universe. It is suggested that quintessence can be either attractive or repulsive depending on the ratio of its kinetic and potential energy. Prepare a review on the physics of quintessence, and its dynamical evolution.
- Literature suggestions:
 - Amendola L., 2000
 Coupled quintessence
 Phys. Rev. D., 62, 043511
 - Ostriker JP, Steinhardt P
 The Quintessential Universe
 Scientific American, 284, 46, (Jan. 2001)
 - Peebles P.J.E., Ratra B., 1988
 Cosmology with a time-variable cosmological 'constant' Astrophys. J. Lett., 325, 17
 - Peebles P.J.E., Ratra B., 2003 The cosmological constant and dark energy Rev. Mod. Phys., 75, 559
 - Pettorino V., Baccigalupi C., 2008
 Coupled and extended quintessence: Theoretical differences and structure formation
 - Phys. Rev. D., 77, 103003
 - Steinhardt P.J., Wang L., Zlatev I., 1999 Cosmological tracking solutions Phys. Rev. D., 59, 123504

12. Phantom Dark Energy

- Phantom dark energy is a hypothetical form of dark energy with an equation of state parameter w < -1, ie. it is even stronger than the cosmological constant at increasing the expansion of the universe. If it exists, it could cause the expansion of the universe to accelerate so quickly that a scenario known as the Big Rip would occur. As a result of its growing energy density and repulsive nature, it would bring about the gravitational unbinding of structures, starting with the largest superclusters and hypothetically down to microscopic scales of planets, rocks, etc. Discuss whether this is a physically viable scenario, or whether there are causality problems with such a scenario. In addition, discuss observational probes that seem as yet to allow the possibility for such a form of dark energy.
- Literature suggestions:
 - Cadlwell R.R., 2002

A phantom menace: cosmological consequences of a dark energy component with super-negative equation of state

Phys. Lett. B., 545, 23, arXiv:astro-ph/9908168

- Caldwell R.R., Kamionkowski M., Weinberg N.N., 2003
 Phantom Energy: Dark Energy with wj-1 Causes a Cosmic Doomsday
 Phys Rev Lett., 91, 071301-1, arXiv:astro-ph/0302506
- Hu W., 2005

Crossing the phantom divide: Dark energy internal degrees of freedom Phys. Rev. D., 71, 047301, arXiv:astro-ph/0410680

 Novosyadlyj, Sergijenko O., Durrer R., Pelykh V., 2012
 Do the cosmological observational data prefer phantom dark energy ? arXiv:1206.5194

13. Holographic Dark Energy

- The holographic principle of quantum gravity theory has been applied to the dark energy problem. So far, three holographic dark energy models have been proposed: the original holographic dark energy model, the agegraphic dark energy model, and the holographic Ricci dark energy model. Give a discussion of these various candidates for dark energy.
- Literature suggestions:
 - Gao C., Fengquan W., Xuelei C., Shen Y-G., 2009
 Holographic dark energy model from Ricci scalar curvature Phys. Rev. D., 79, 043511
 - Li M., Li X-D., Wang S., Zhang X., 2009
 Holographic dark energy models: A comparison from the latest observational data
 JCAP 06, 036, arXiv:0904.0928
 - Li M., Li X-D., Lin C-S., Wang Y., 2009
 Holographic Gas as Dark Energy
 Comm. Theor. Phys., 51, 181, arXiv:0811.3332

14. Future Dark Energy Surveys: Euclid and LSST

- In the coming decade, two huge systematic survey instruments will be focussed on the determination of the equation of state of dark energy. These are the ESA Euclid mission and the Large Synoptic Telescope. Discuss these instruments, and their promise for the determination of dark energy. This involves a detailed discussion of the improvement of the determination of dark energy parameters by various dark energy probes, such as gravitational weak lensing, supernova acceleration measurements, baryonic acoustic oscillations, etc.
- Literature suggestions:
 - Amendola L., et al., 2012 Cosmology and fundamental physics with the Euclid satellite arXiv:1206.1225
 - ESA

Euclid, Mapping the geometry of the dark Universe ESA/SRE(2011)12, July 2011

- Abate A., et al., 2012 Large Synoptic Survey Telescope Dark Energy Science Collaboration
 - arXiv:1211.0310