

A visualization of the Cosmic Web, showing a complex network of orange and yellow filaments and nodes against a dark blue background. The filaments represent the large-scale structure of the universe, with nodes representing galaxy clusters and individual galaxies.

the Cosmic Web:

Lecture 5: Voids

Rien van de Weijgaert,
Cosmic Web, caput course, October 2017

Cosmic Web

Void Evolution

Void Formation

Void Evolution

an illustration

cosmology:

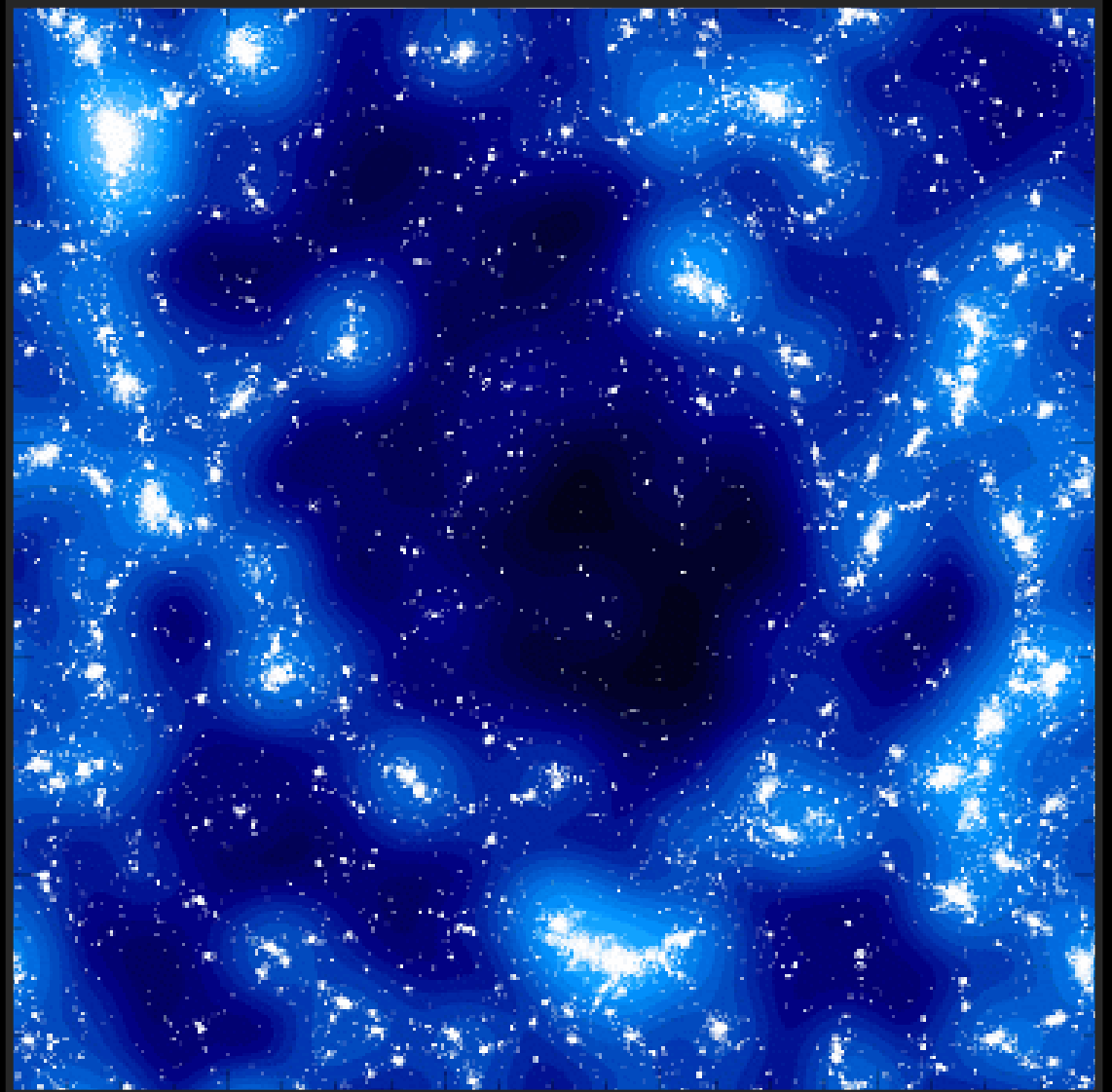
$$\Omega_m = 1.0; \quad H_0 = 70 \text{ km / s / Mpc}$$

initial conditions:

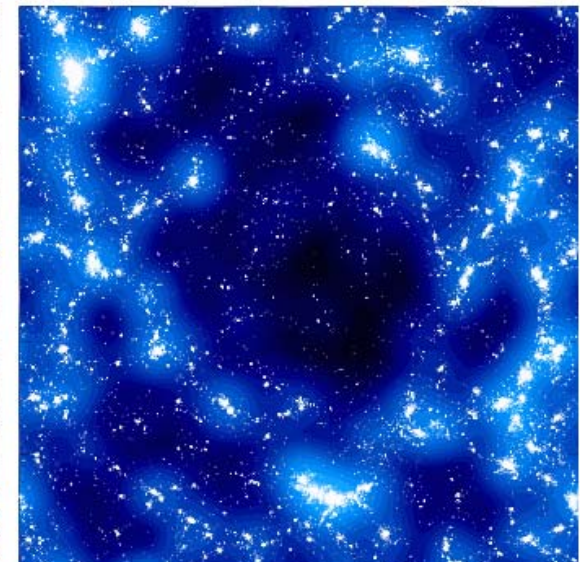
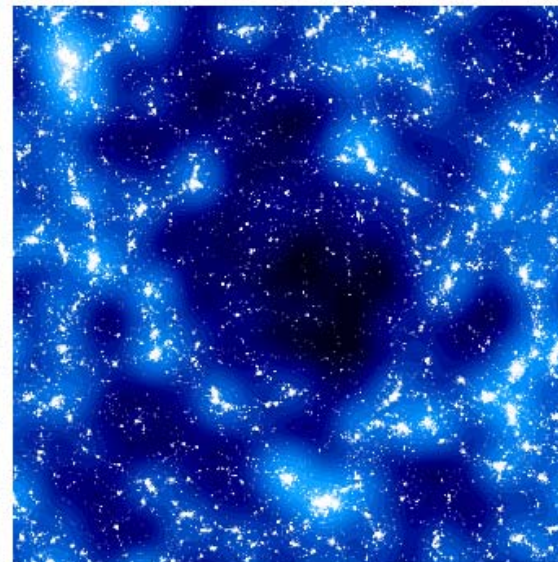
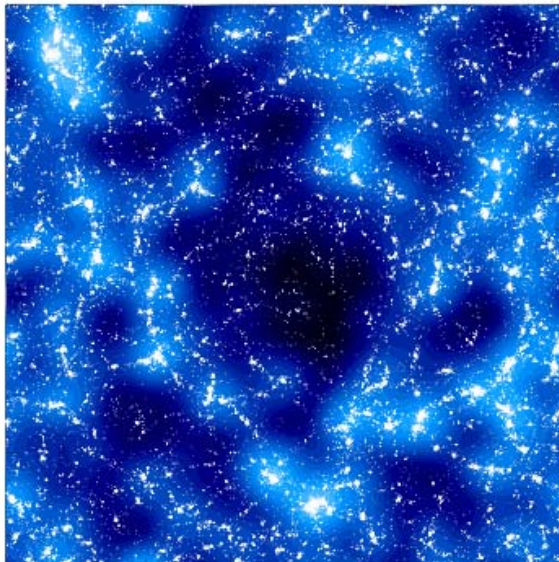
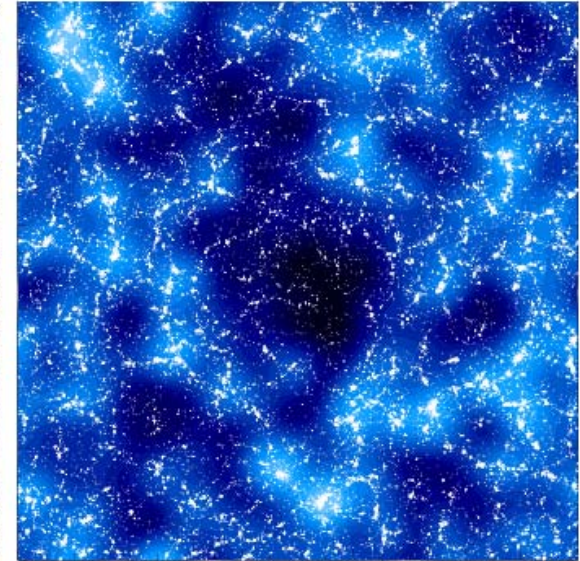
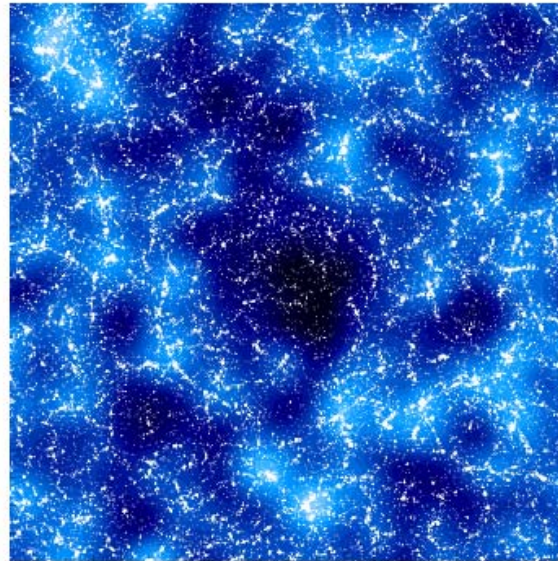
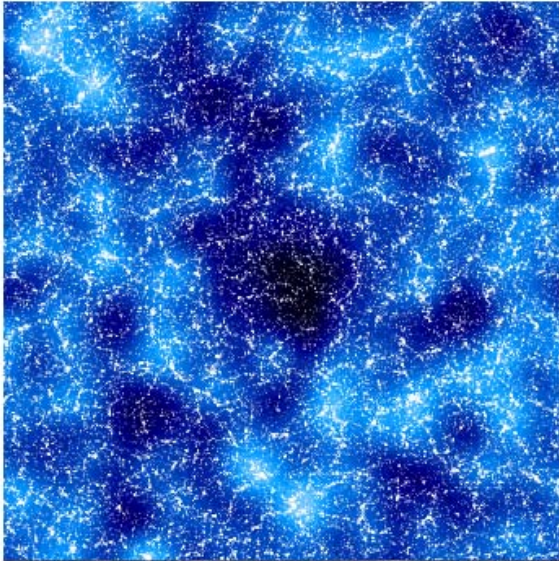
underdensity, Gaussian field

$$R_G \sim 4h^{-1} \text{ Mpc}$$

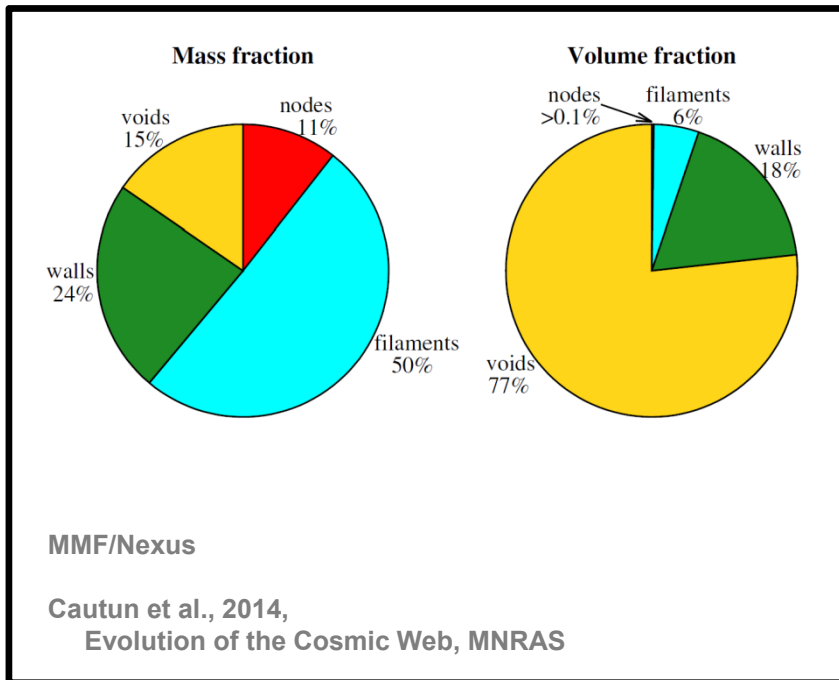
$$P(k) \propto k^{-0.5}$$



Void Formation



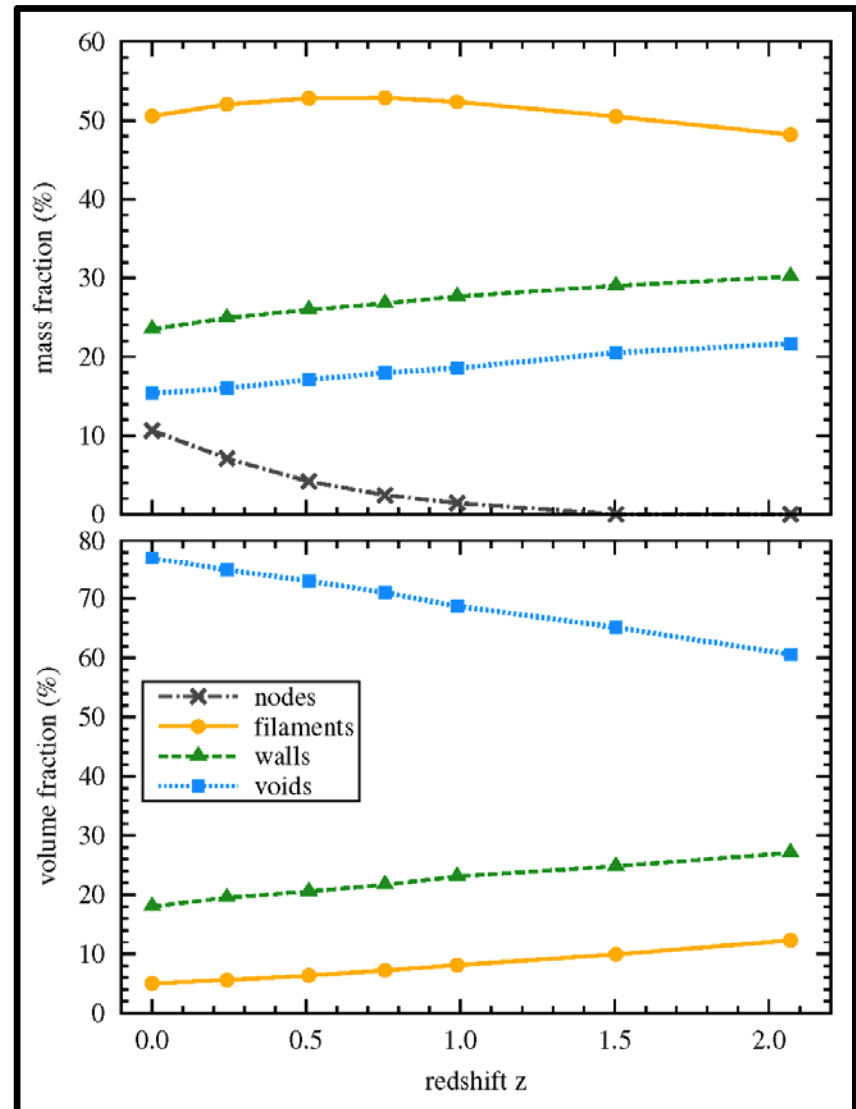
Cosmic Web & Voids



Voids: - occupy most of cosmic volume: **77%**
 - of mass, only: **15%**

Void evolution:

- volume fraction increases with time
(void expansion)
- mass fraction decreases with time
(void evacuation)

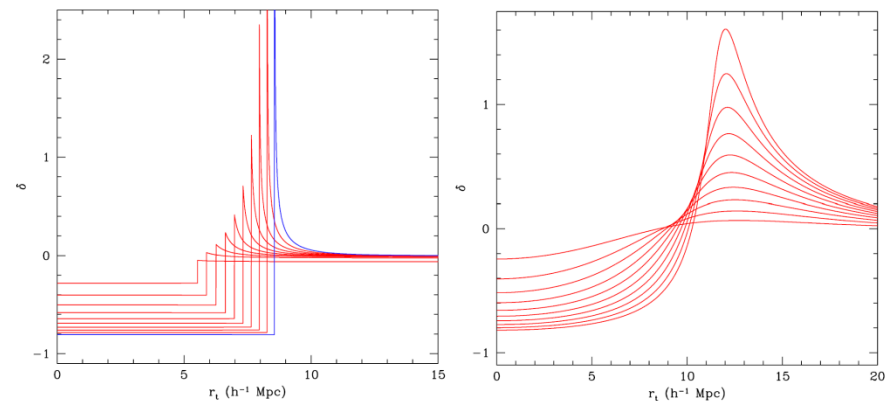


Voids:

Global Evolution

Void Evolution: The Perfect Sphere, Tophat as well

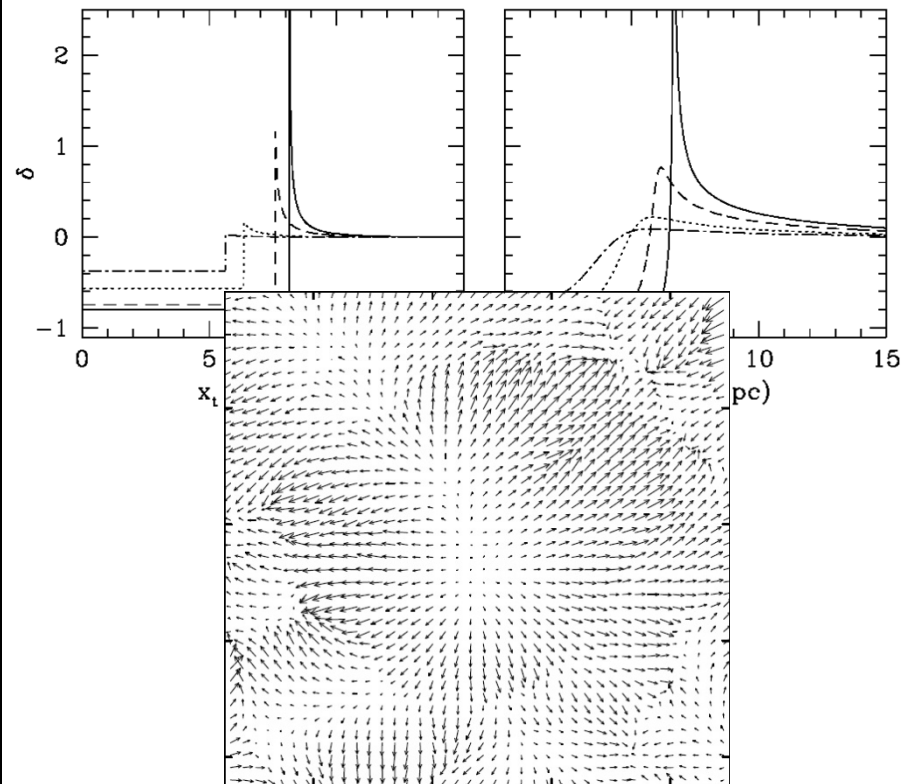
- **“Bubble Theorem”**
Voids become increasingly spherical, due to anisotropic outward directed force
- **Tophat Configuration**
Any initial configuration tends towards “bucket” shape
- **Density Ridge**
Except for gentlest initial density profiles, a ridge forms



Superhubble Expansion

- Superhubble Expansion tending towards “bucket” shape, the void outflow is one with uniform velocity divergence

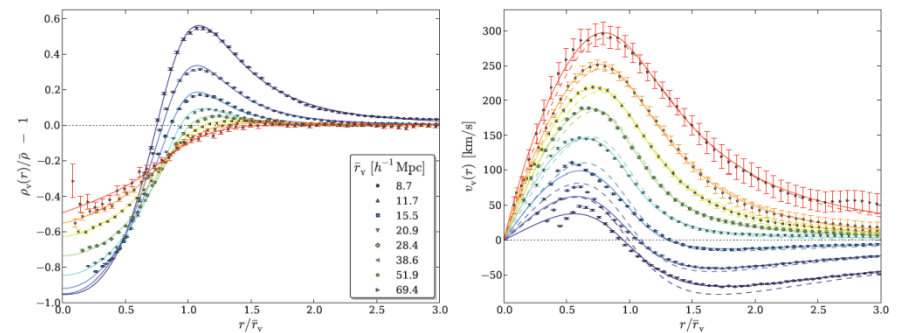
$$\theta = \frac{1}{H}(\nabla \cdot \vec{v}) \quad \Rightarrow \quad \theta_{\max} = 1.5\Omega^{0.6}$$



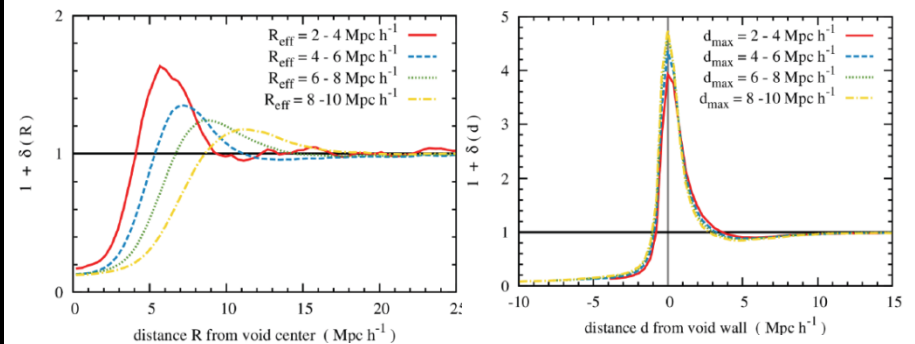
Void Density Profile

- **Tophat Configuration**
Any initial configuration tends towards “bucket” shape
- **Density Ridge**
Except for gentlest initial density profiles, a ridge forms

Is there a universal void density profile ?



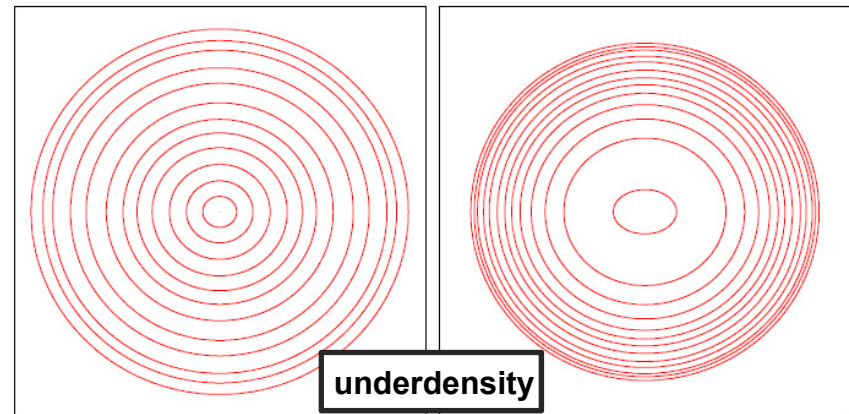
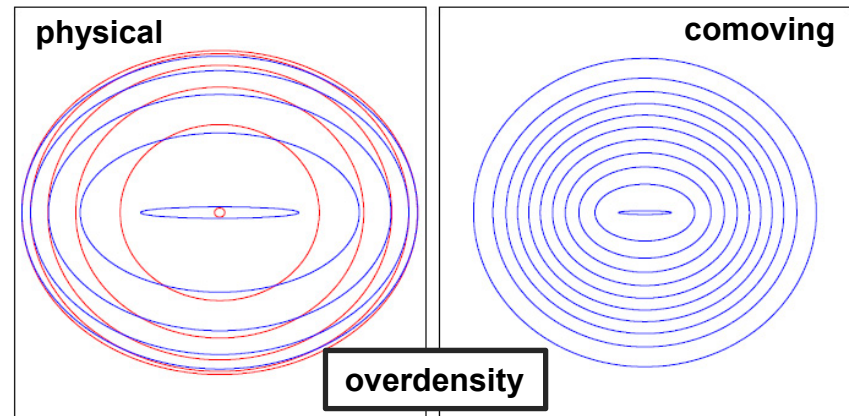
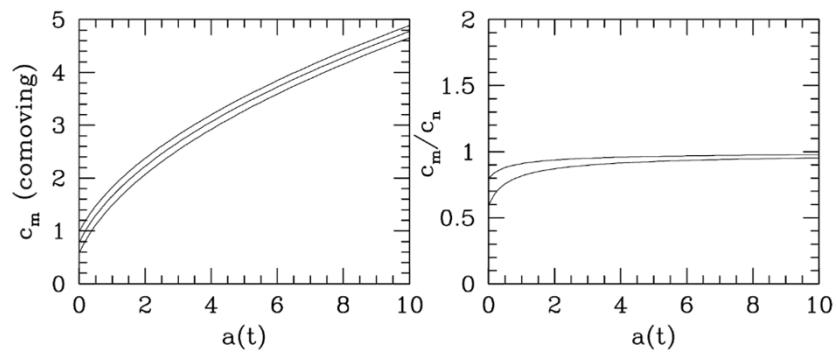
e.g. Hamaus et al. 2014



Cautun et al. 2015

Bubble Theorem

- **Bubble Theorem (Icke 1984)**
Isolated voids tend to become more spherical as they expand and evolve



Void Shapes & Environment

- **Bubble Theorem Revisited:**

voids will not be spherical:

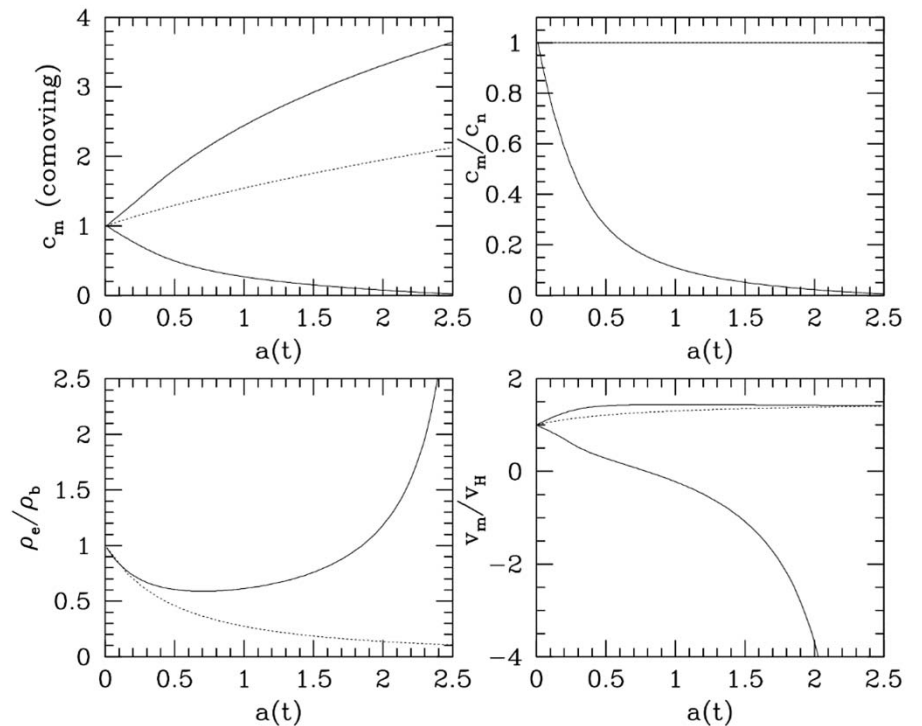
- ☐ Voids never isolated: run into neighbours

- ☐ Void evolution largely dominated by large scale (tidal) environment:

voids always represent restricted density fluctuation: $|\delta| < 1$

Evolution homogeneous ellipsoidal void in external tidal field T_{kl}

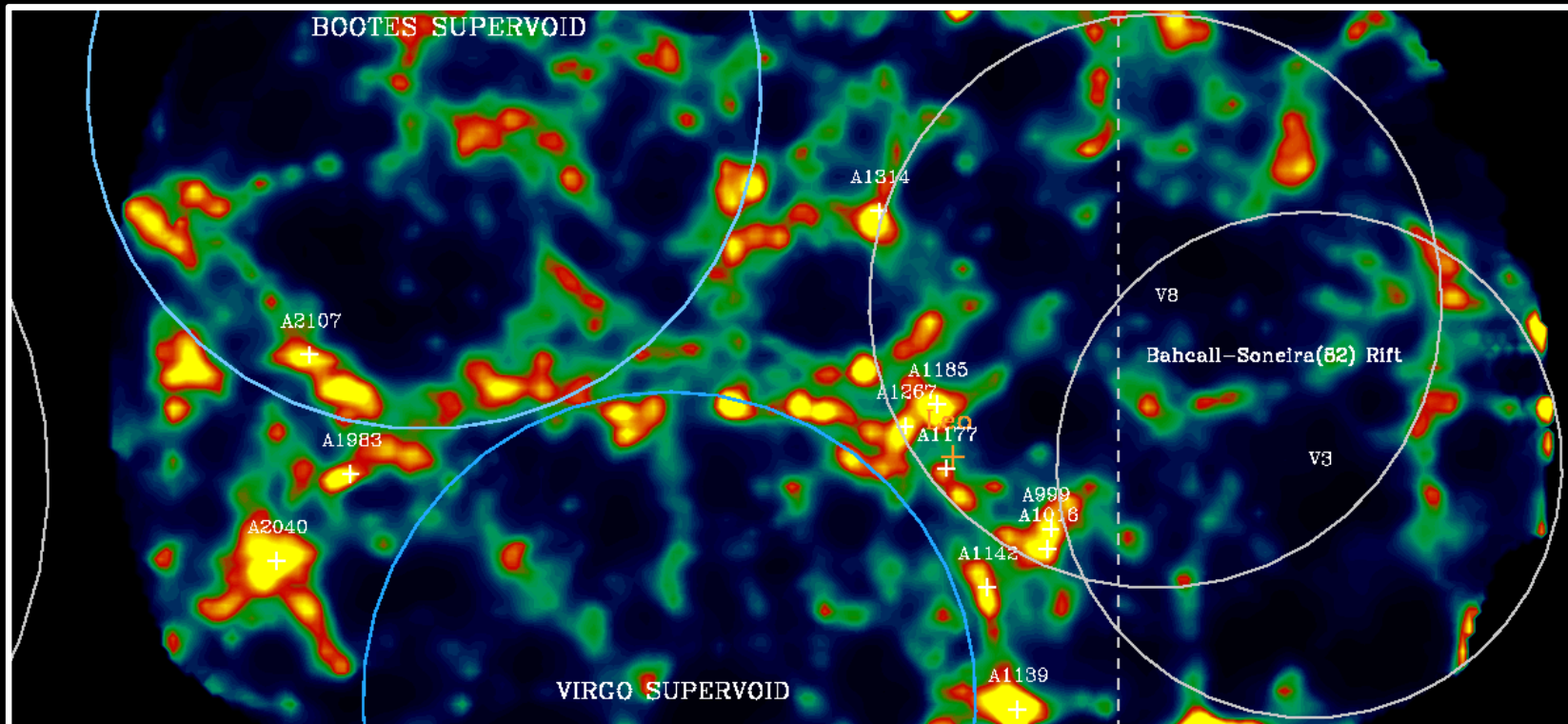
$$\frac{d^2 R_m}{dt^2} = -2\pi G \left[\alpha_m \rho_e + \left(\frac{2}{3} - \alpha_m \right) \rho_u \right] R_m - T_{mm} R_m$$



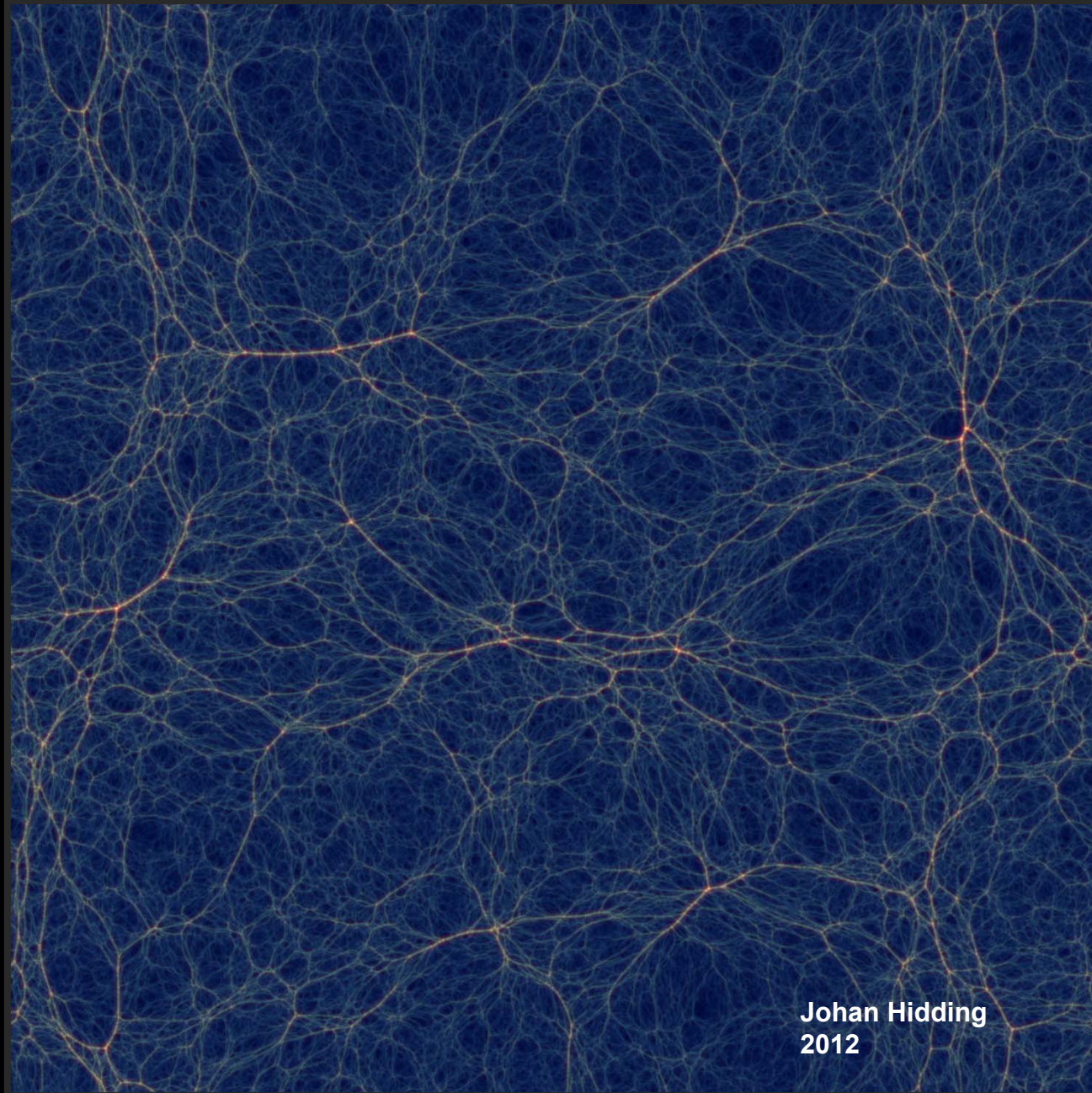
Voids:

Multiscale Infrastructure

Bootes Void: Substructure



Platen et al. 2009



Johan Hidding
2012

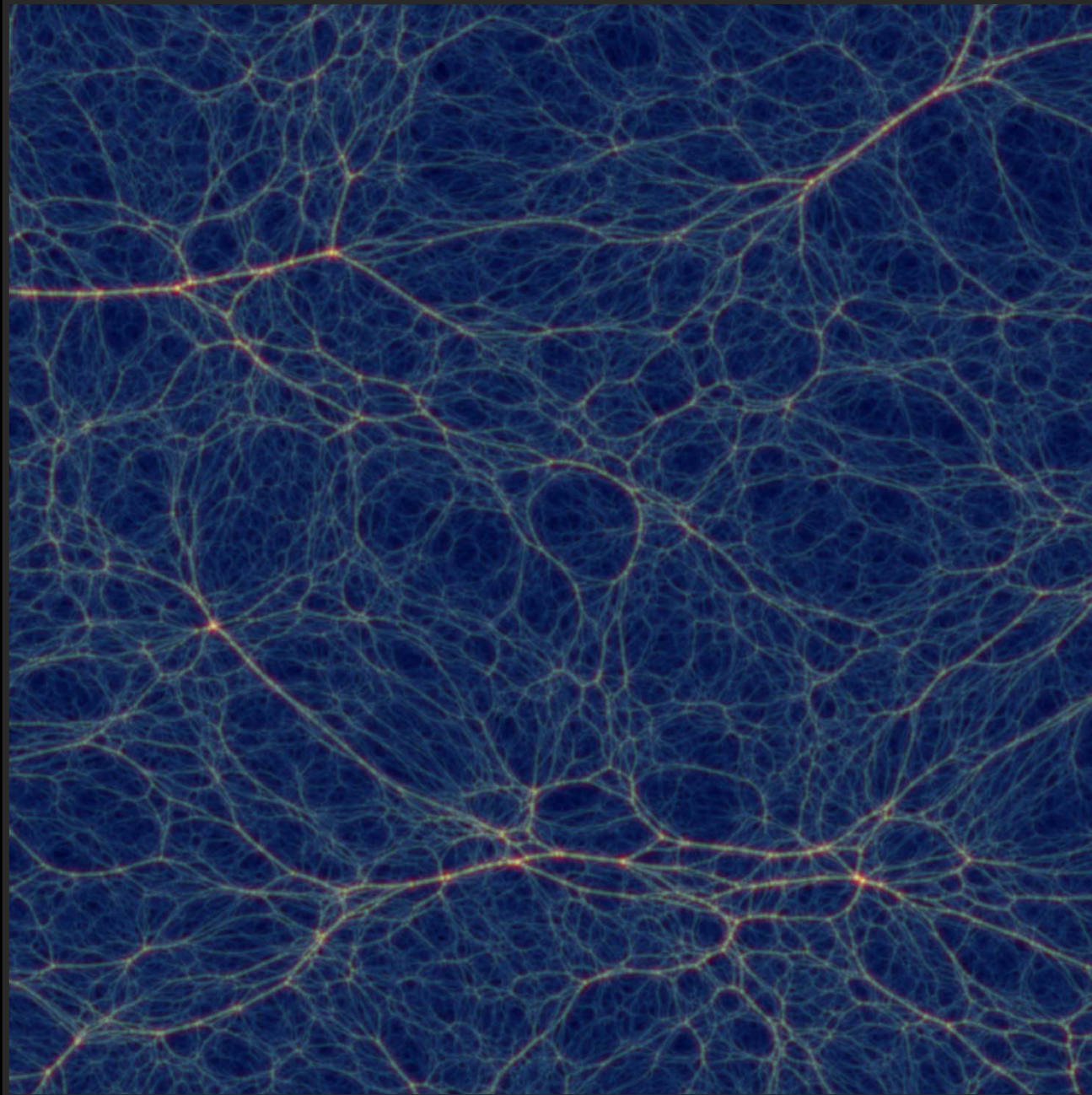
**Multiscale
Infrastructure**

Voids:



manifestation

**Hierarchical
Buildup of
Voids**



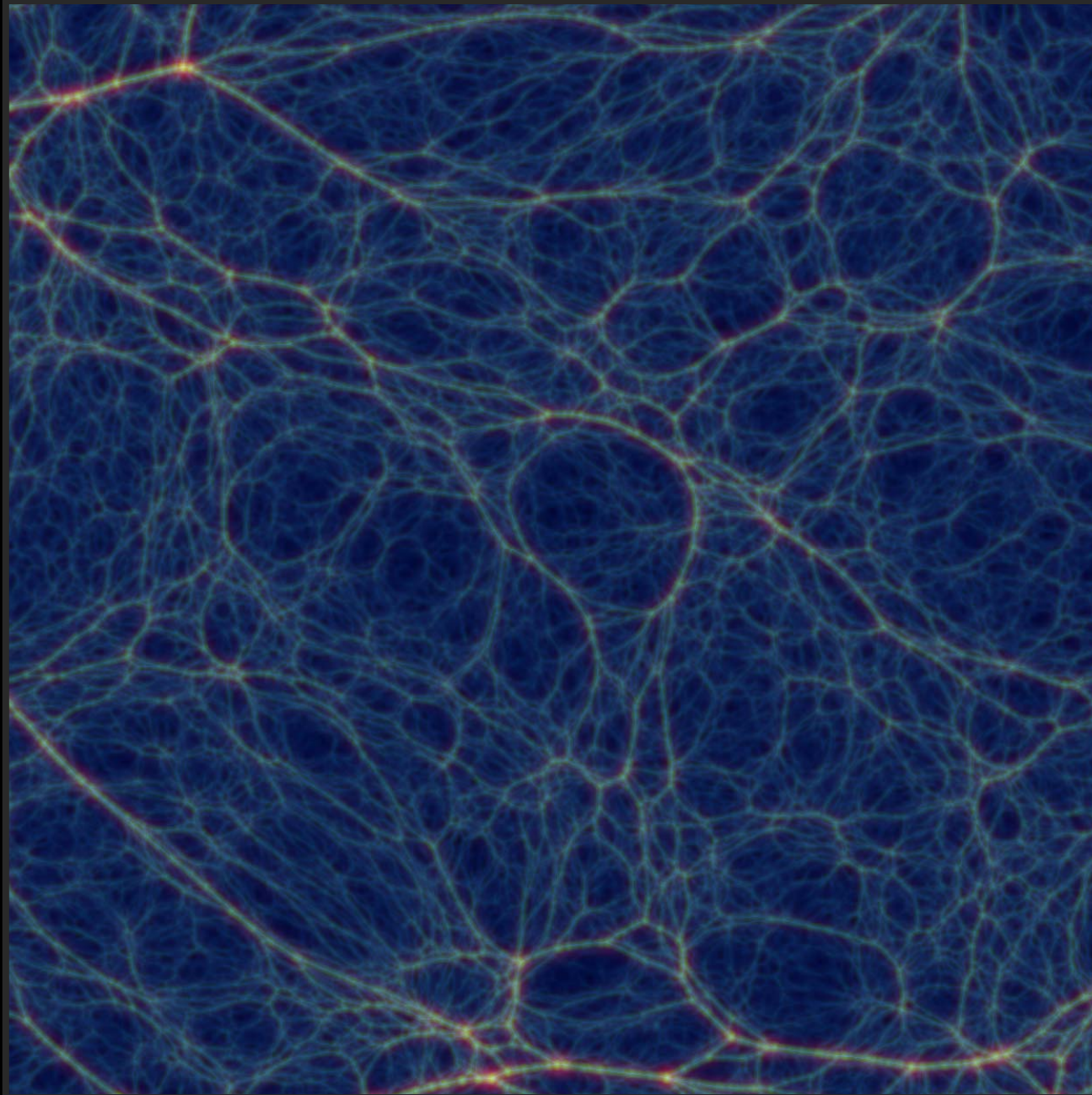
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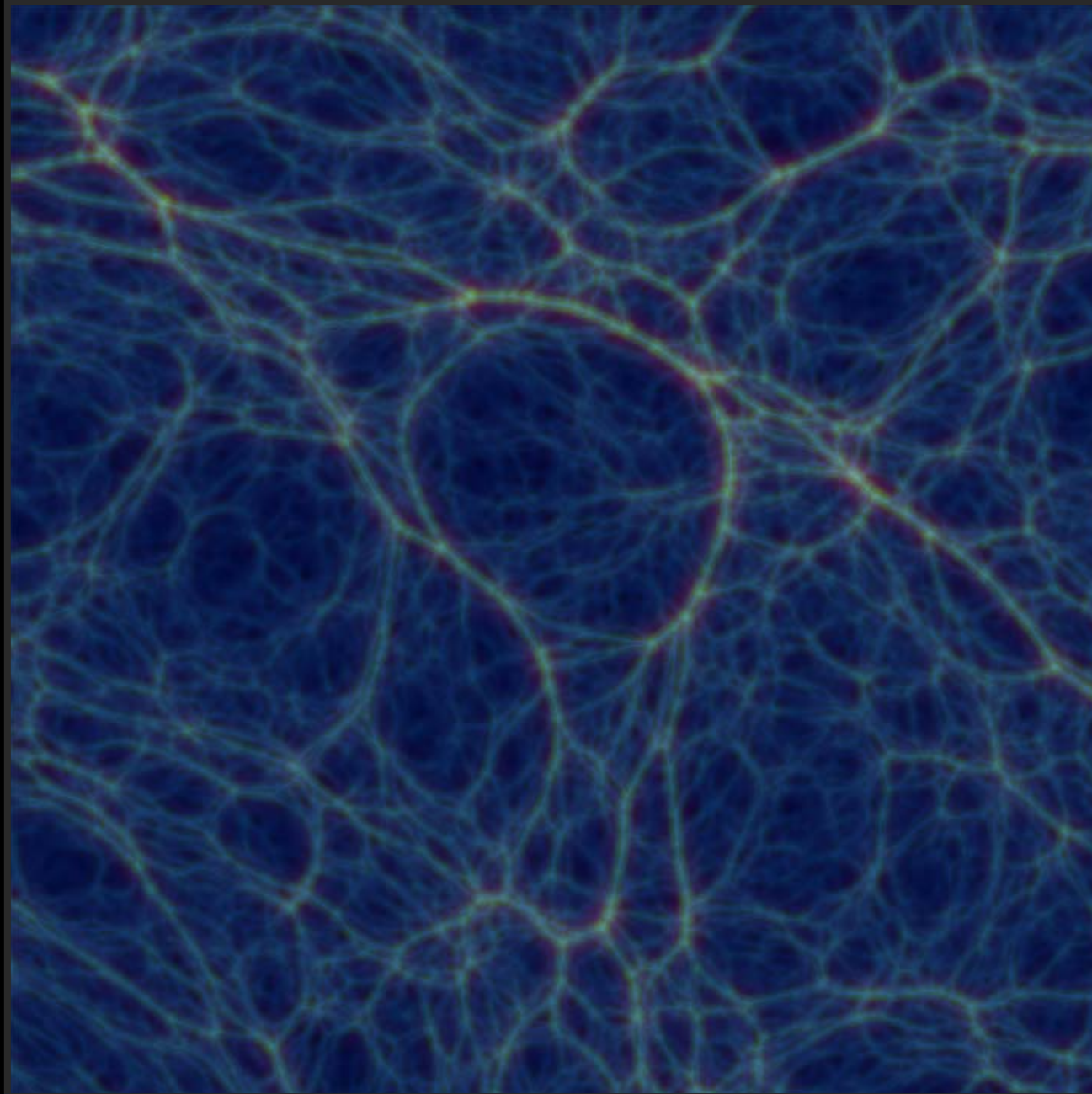
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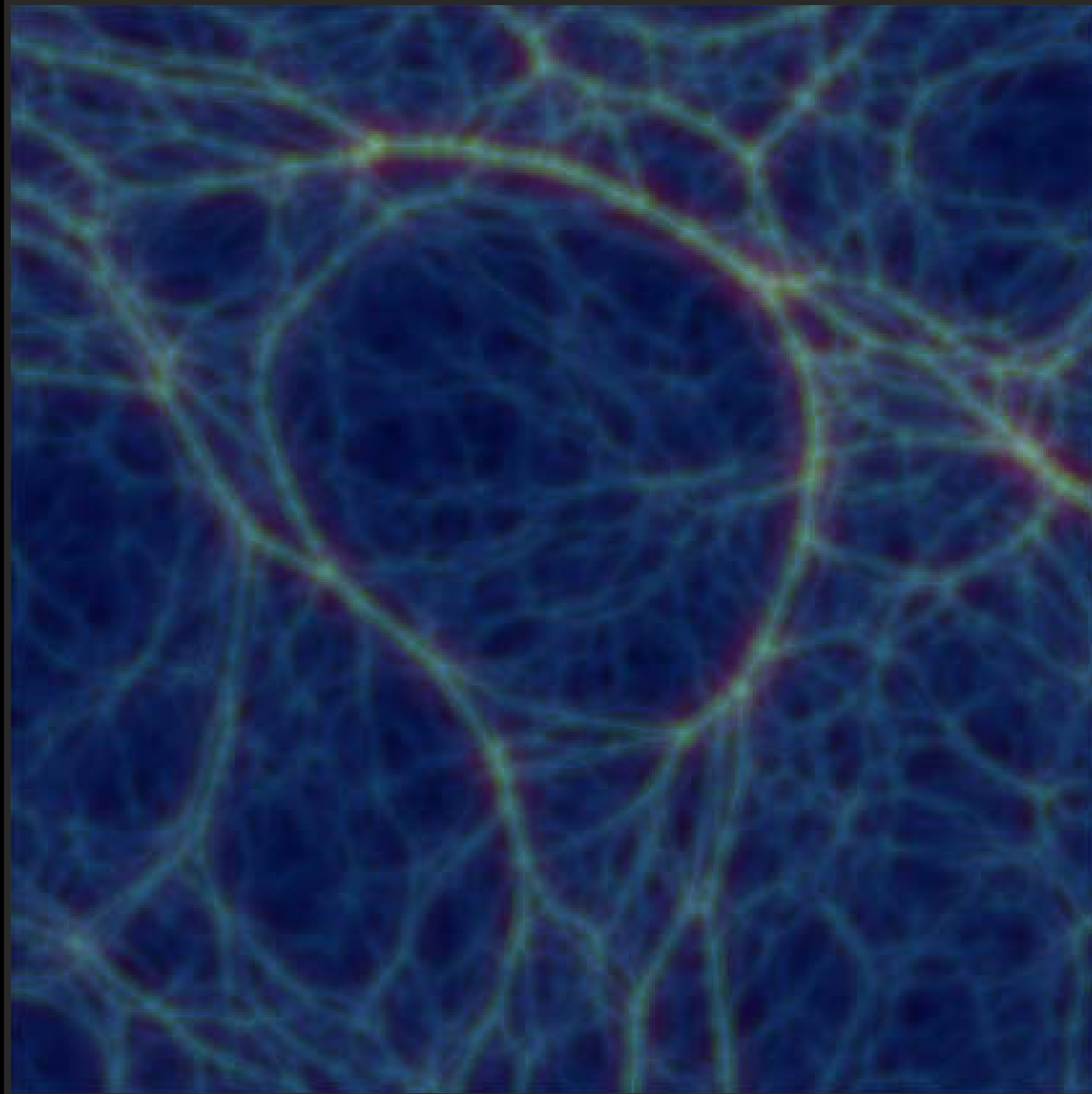
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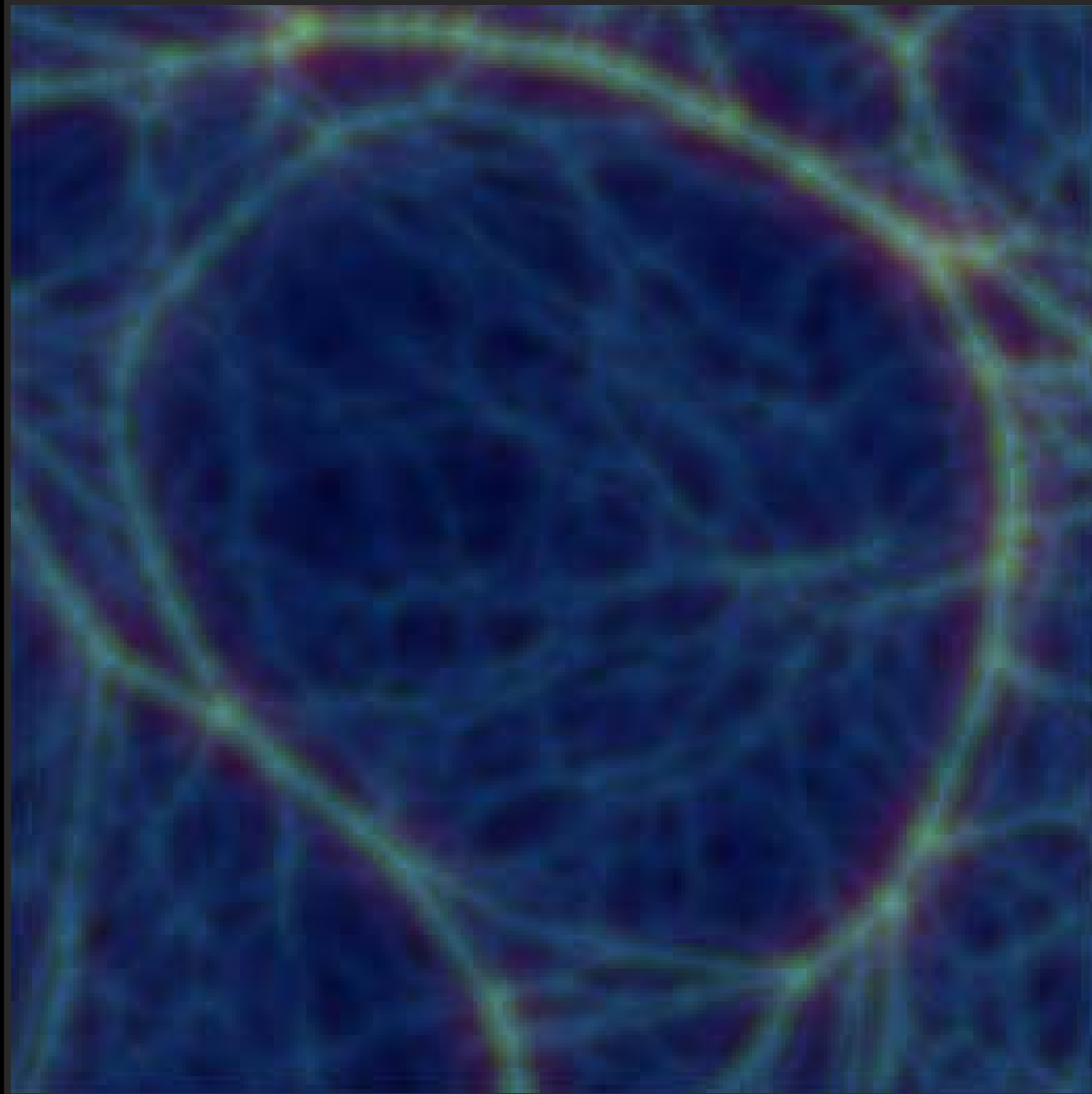
**Multiscale
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Voids:



manifestation

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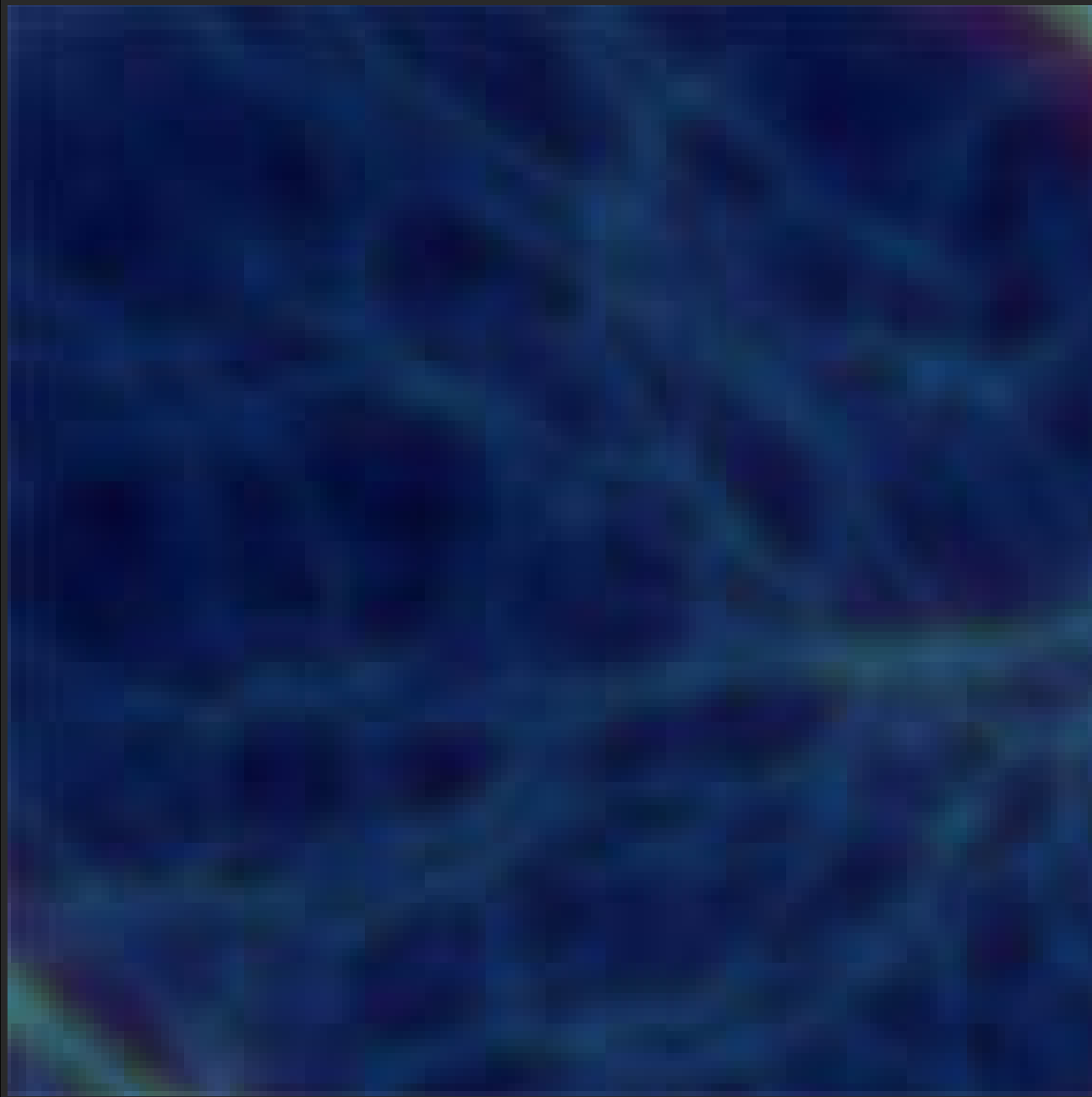
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Voids:



manifestation

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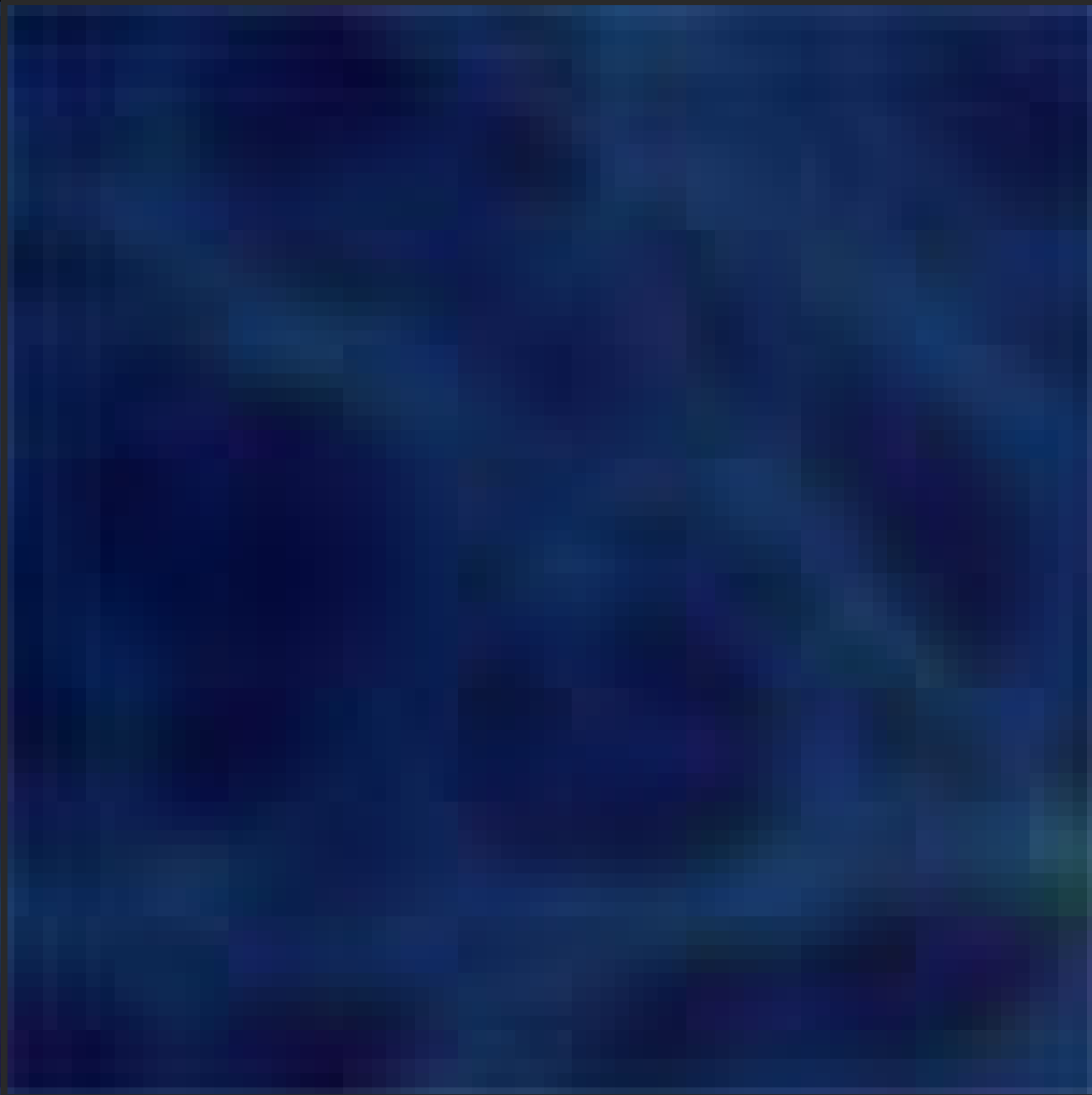
**Multiscale
Infrastructure**

Voids:



manifestation

**Hierarchical
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**Multiscale
Infrastructure**

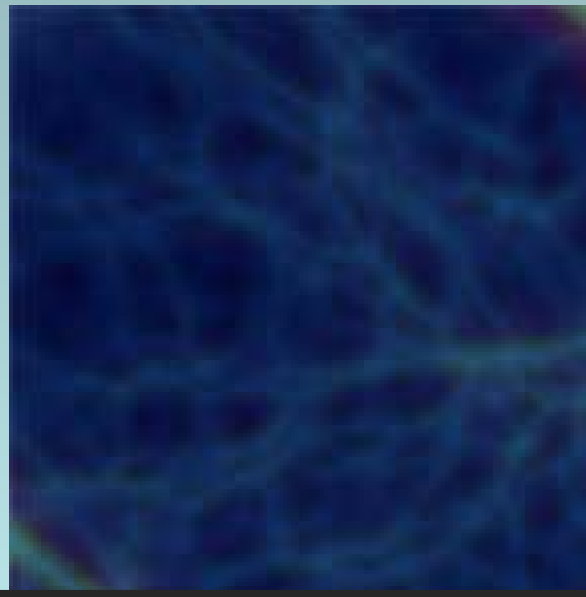
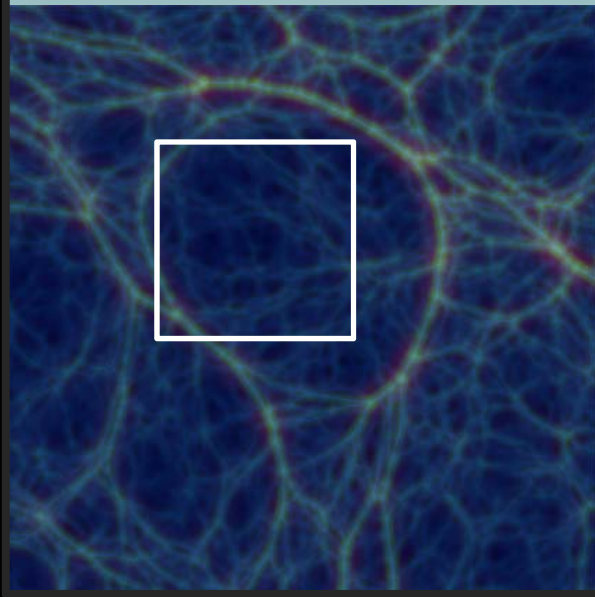
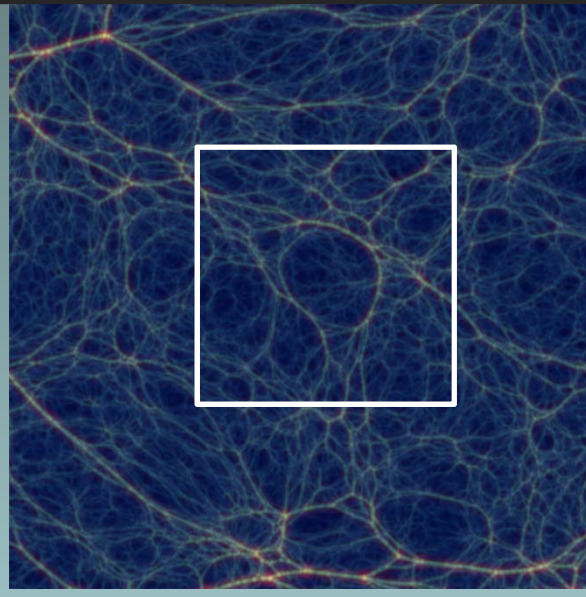
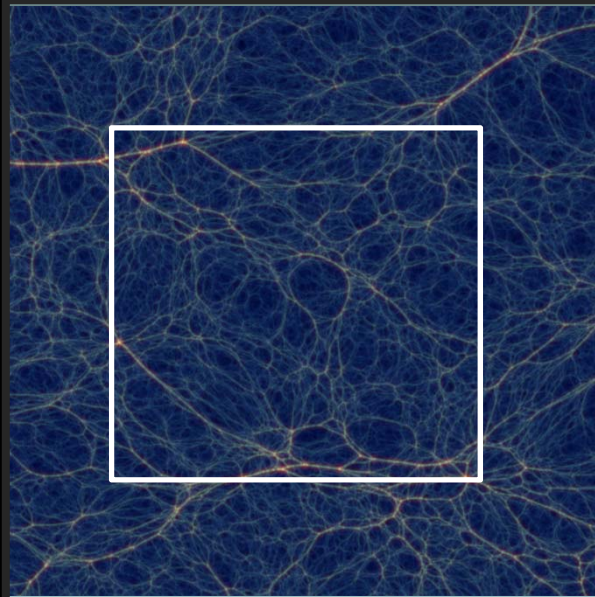
Voids:



manifestation

**Hierarchical
Buildup of
Voids**

Void Substructure



Zoom in: 3 levels

Substructure on
all scales:

amplitude
diminishing towards
smaller scales

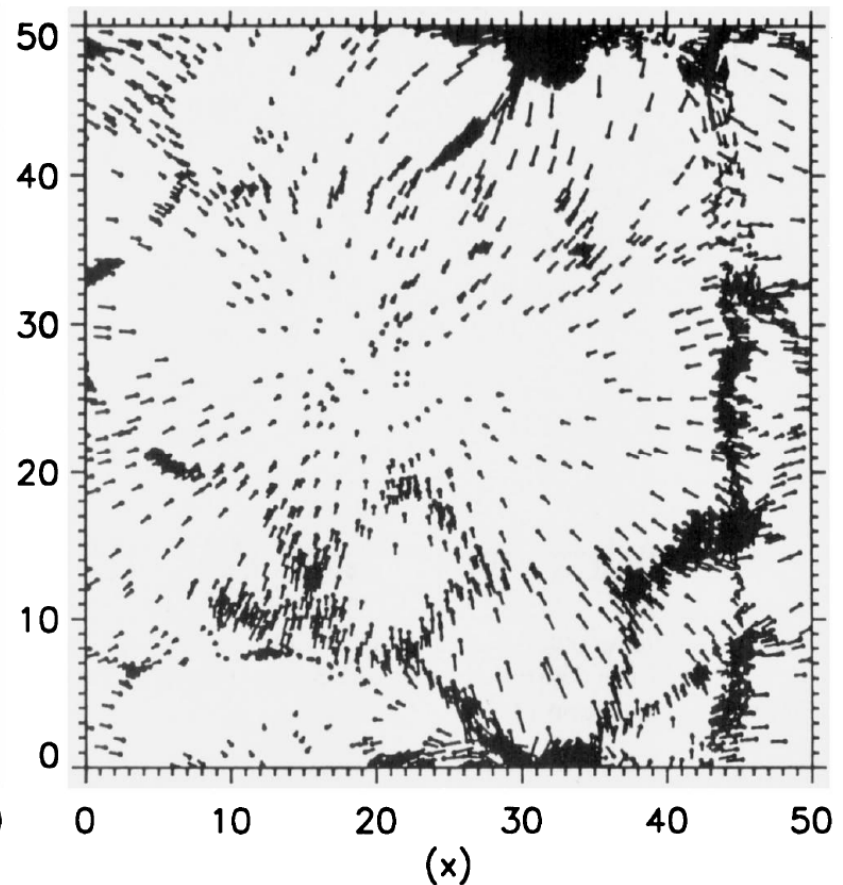
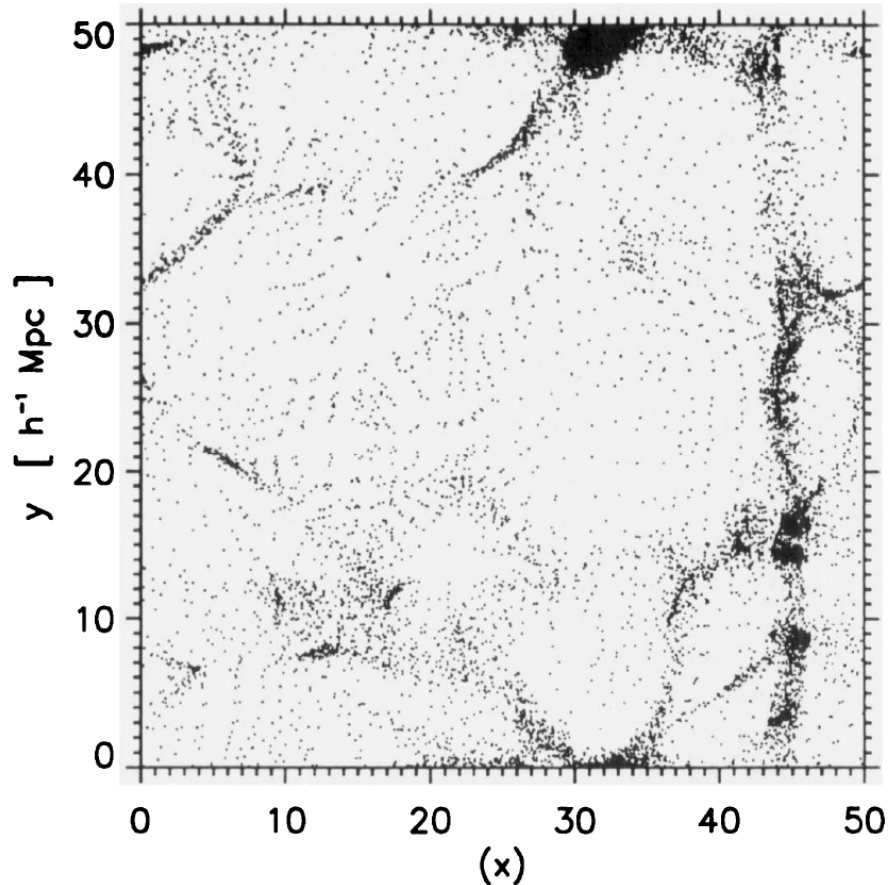
Voids:

Hierarchical Evolution

Void Hierarchy

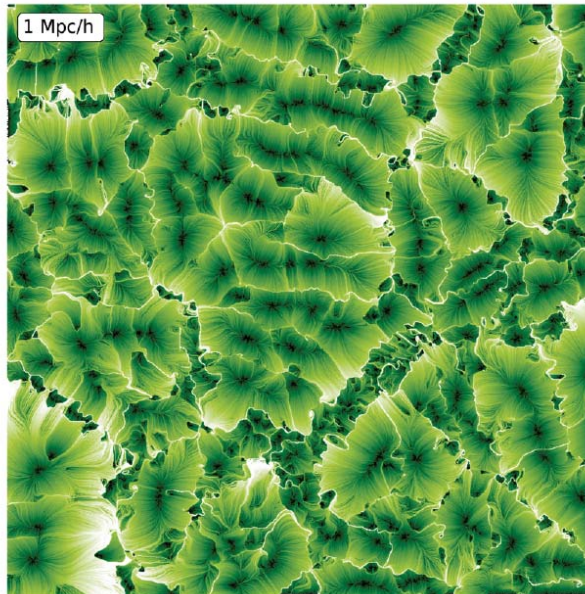
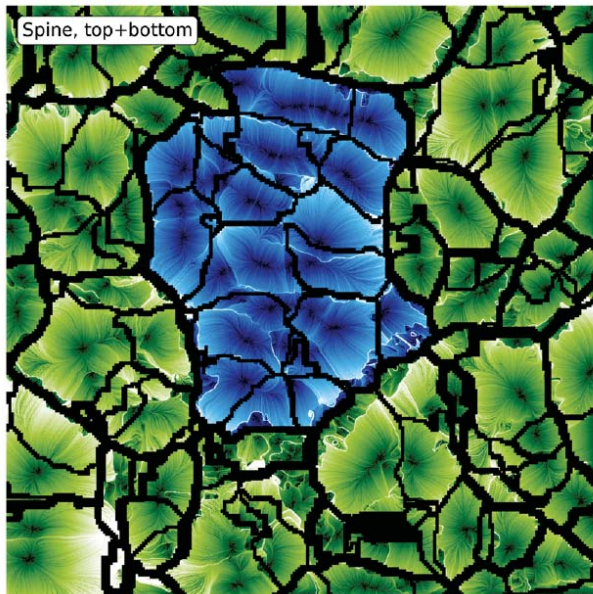
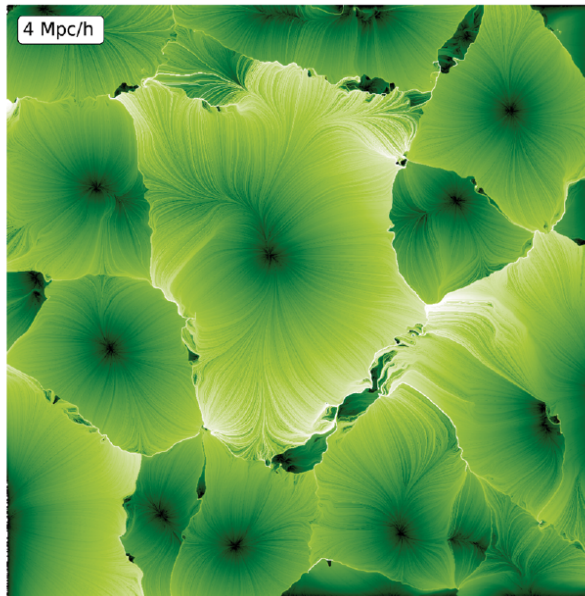
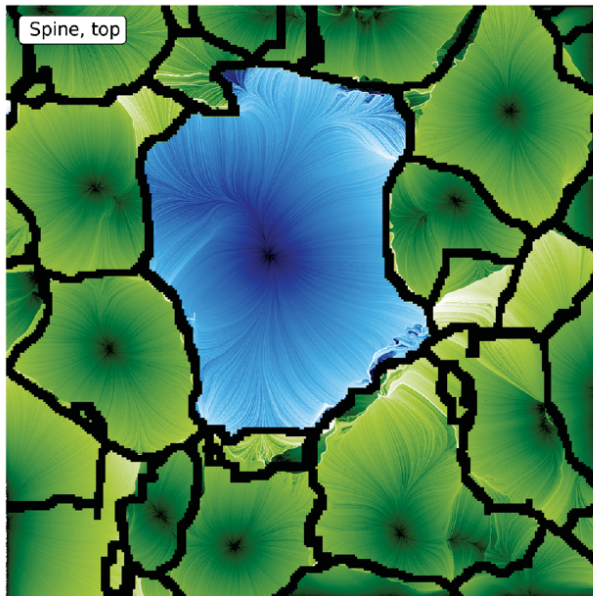
Void formation: - via gradual merging of voids
- demolition of small voids

Dubinski et al. 1993
Van de Weygaert & van Kampen 1993



Constrained Void Simulations:
Van de Weygaert 1991
vdW & van Kampen 1993

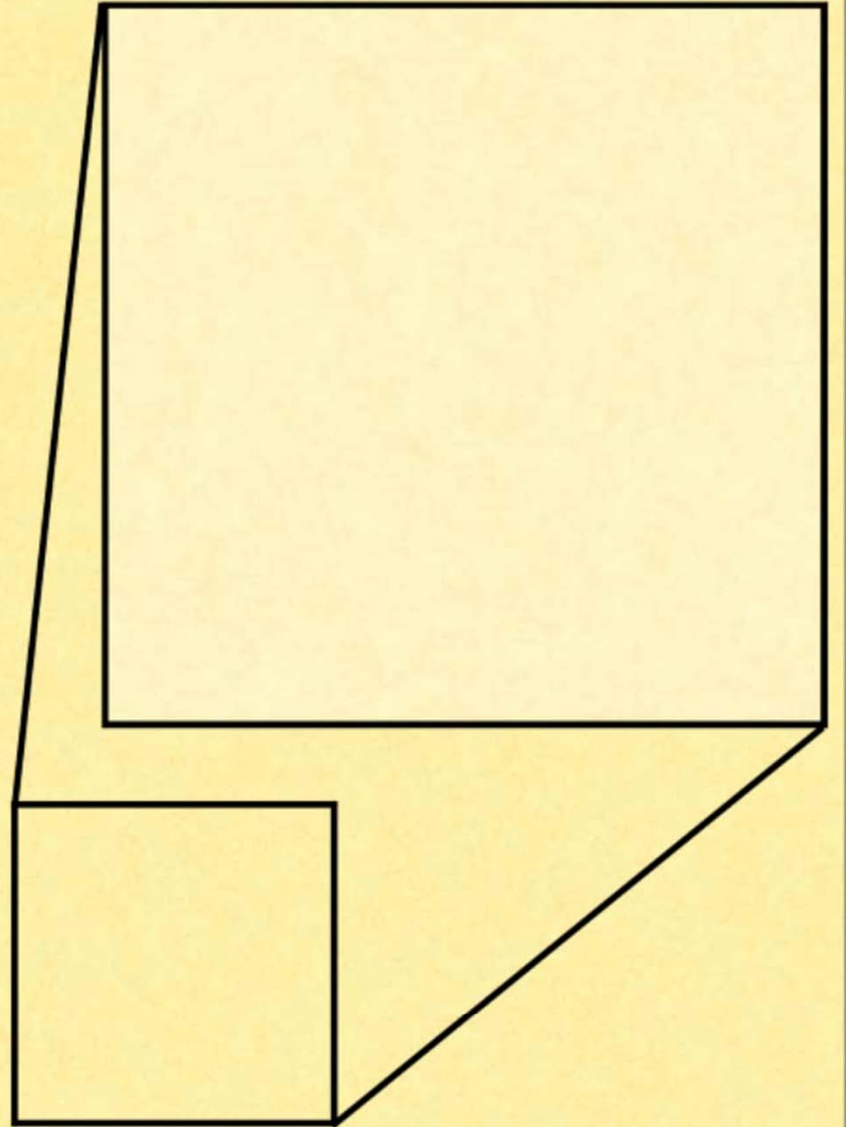
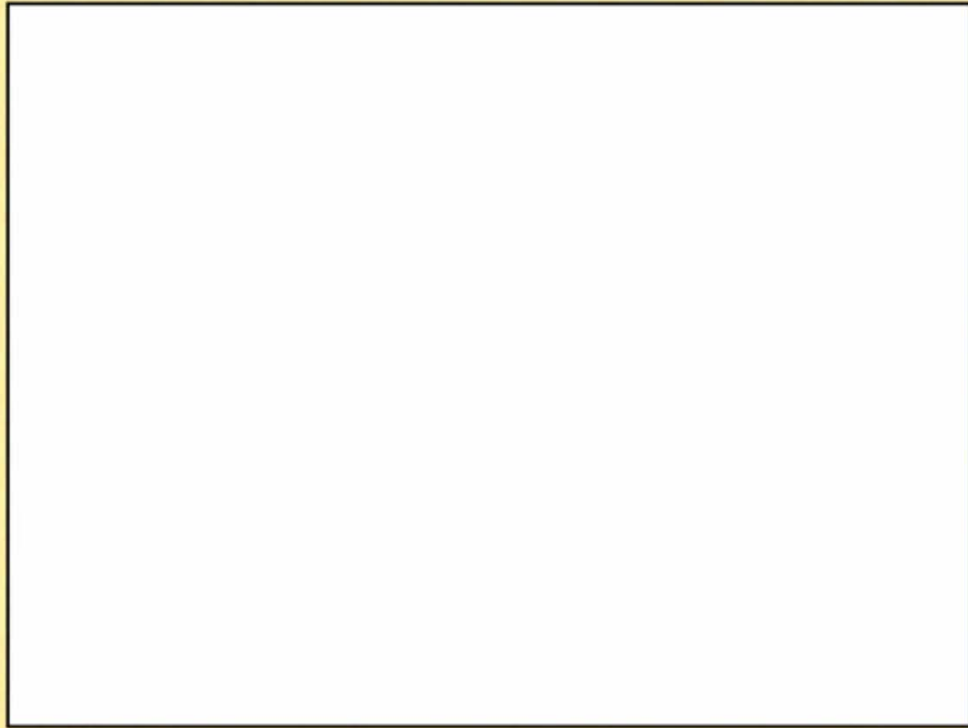
CRF formalism:
Bertschinger 1987
vdW & Bertschinger 1996



Hierarchical Web Evolution:

Void hierarchy
expressed in
multiscale structure
velocity outflow

Aragon-Calvo & Szalay 2012

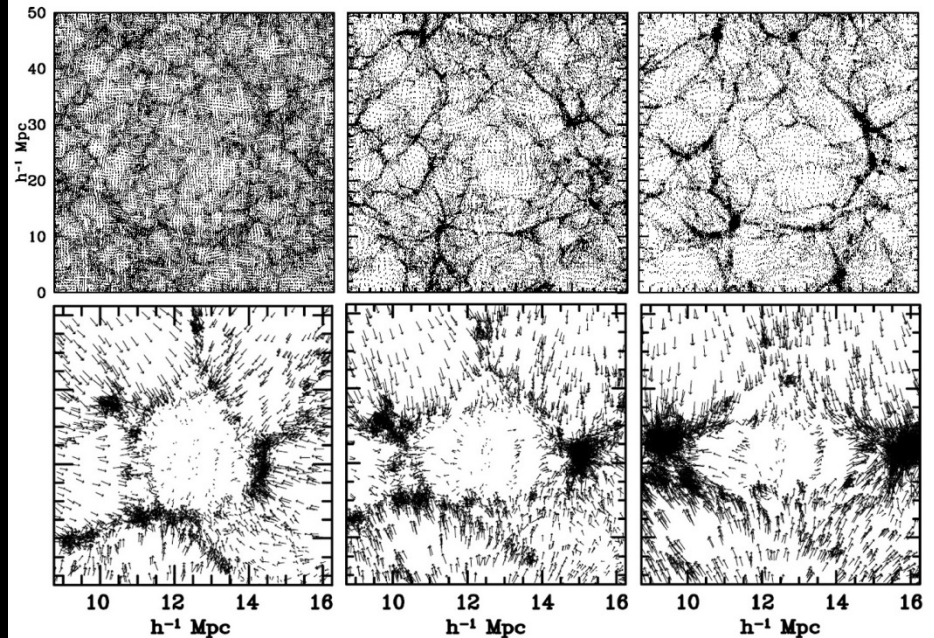


Void Dichotomy

- ❑ **Voids emerge out of**
primordial Gaussian density depressions
- ❑ **Primordial Gaussian Density Field:**
symmetry overdense \longleftrightarrow underdense
- ❑ **Why is void population:**
Not dominated by small voids ?
What happened to (PS) hierarchy ?

Void Evolution Processes

- **Void Merging**
as voids expand and meet their peers, they merge into ever larger voids ...
- **Void Collapse**
when embedded within an overdense or tidally sheared region (filaments ...), weak voids get squeezed out of existence...



Hierarchical

Web Evolution:

Adhesion simulation

buildup Cosmic Web

Johan Hidding
2012

Extended Press-Schechter

Barrier Excursions

- Spherical linear collapse overdensity:

$$\Delta_{lin}(r, S_m, t) > \delta_c$$

- Collapse time:

$$a_{coll}(r) = \delta_c / \Delta_{lin}(r, S_m)$$

- Initial density field:

prediction object formation time:

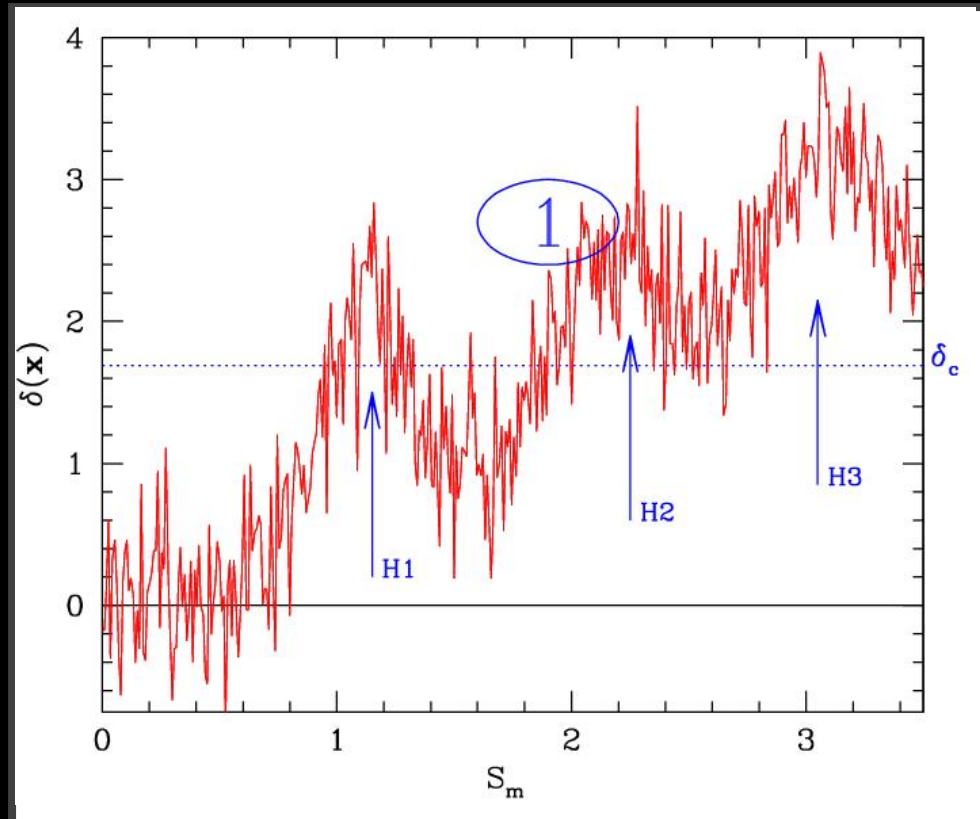
$$a_{coll}(r) \Leftrightarrow \Delta_{lin,0}(r, S_m)$$

dependent on:

Collapse Barrier δ_c

- Critical density value:

$$\text{EdS, } \Omega_0=1: \quad \delta_c \sim 1.69$$



cumulative random walk:

$$\delta_s(\vec{x}; \lambda_m) = \int_{|k| < k_m} \frac{d\vec{k}}{(2\pi)^3} \hat{\delta}(\vec{k}) e^{-i\vec{k} \cdot \vec{x}}$$

Void Evolution: the Two-Barrier Formalism

Two critical barriers:

- embedded in underdensity
VOID-IN-VOID
- embedded in overdensity
VOID-IN-CLOUD
- to be counted in as voids, take only the ones who pass the void barrier, but do not pass through the cloud barrier at larger scale

Void-in-Cloud:

- key towards explaining absence small voids
- small voids get demolished in overdensities
- population of cosmic “halos” dominated by small clumps (divergent)



void ranks dominated by voids ~ specific void size

Extended Press-Schechter

Barrier Excursions

- Spherical linear collapse overdensity:

$$\Delta_{lin}(r, S_m, t) > \delta_c$$

- Collapse time:

$$a_{coll}(r) = \delta_c / \Delta_{lin}(r, S_m)$$

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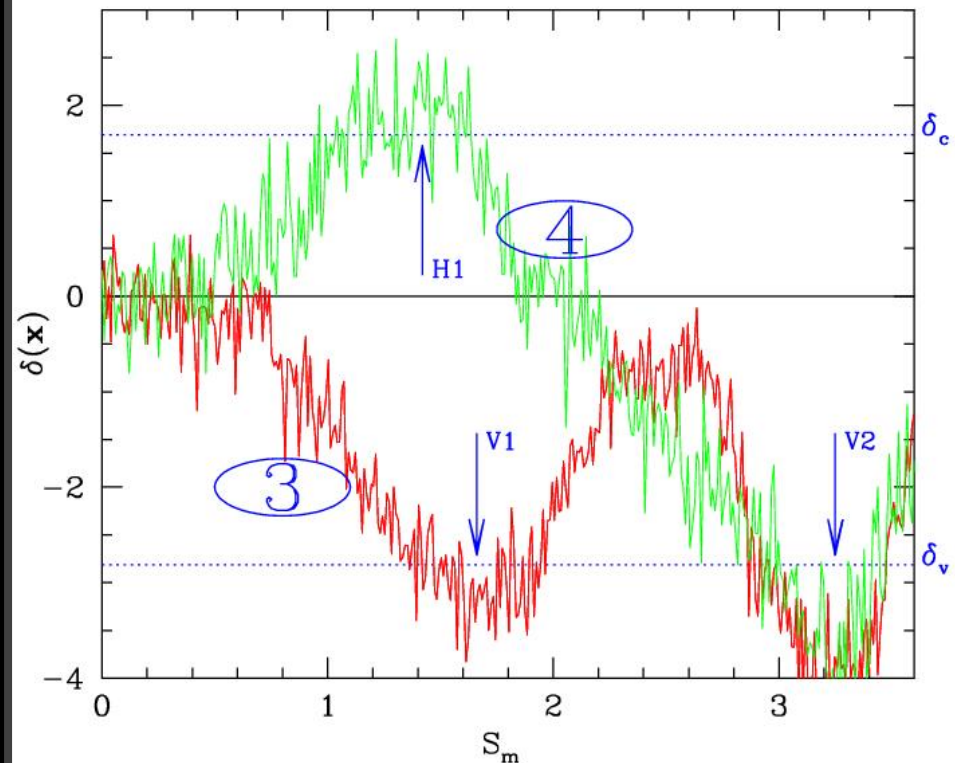
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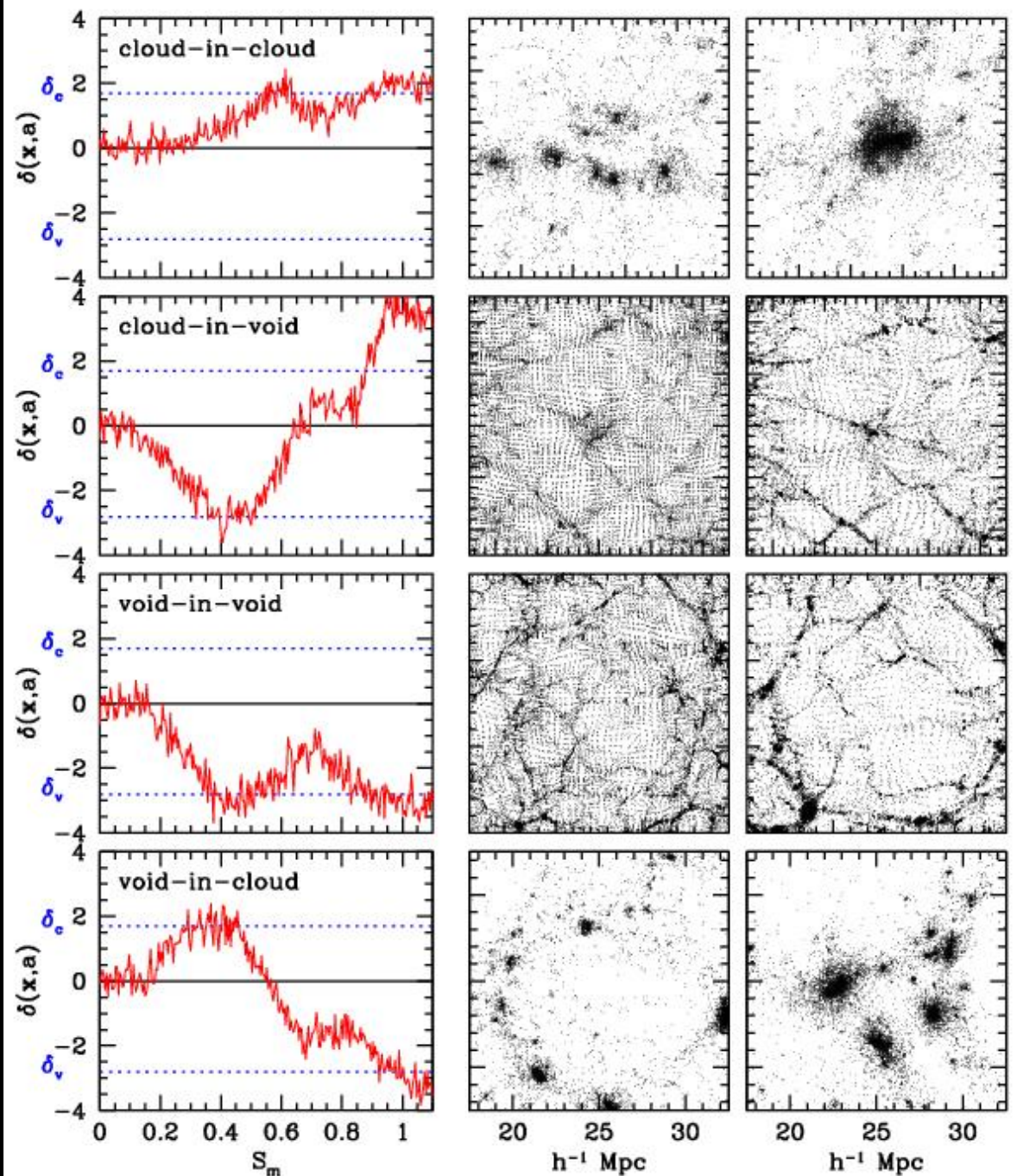
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Hierarchical Cosmic Structure Formation

Excursion Modes:

- Cloud-in-Cloud
- Cloud-in-Void
- Void-in-Void
- Void-in-Cloud



Void Volume Distribution

- **Small Void tail suppressed**
- **Peaked Void Size Distribution**
Characteristic Void Size
- **Self-Similar Evolution:**
increasing
characteristic Void Size
- **Volume-filling:**
at any cosmic epoch,
for power-law P(k),
approximately void-filling
- **Excess Void Expansion:**
“Super-Hubble” expansion

$$v_v(M) = \frac{|\delta_v|}{\sigma(M)} \quad \mathcal{D} \equiv \frac{|\delta_v|}{(\delta_c + |\delta_v|)}$$

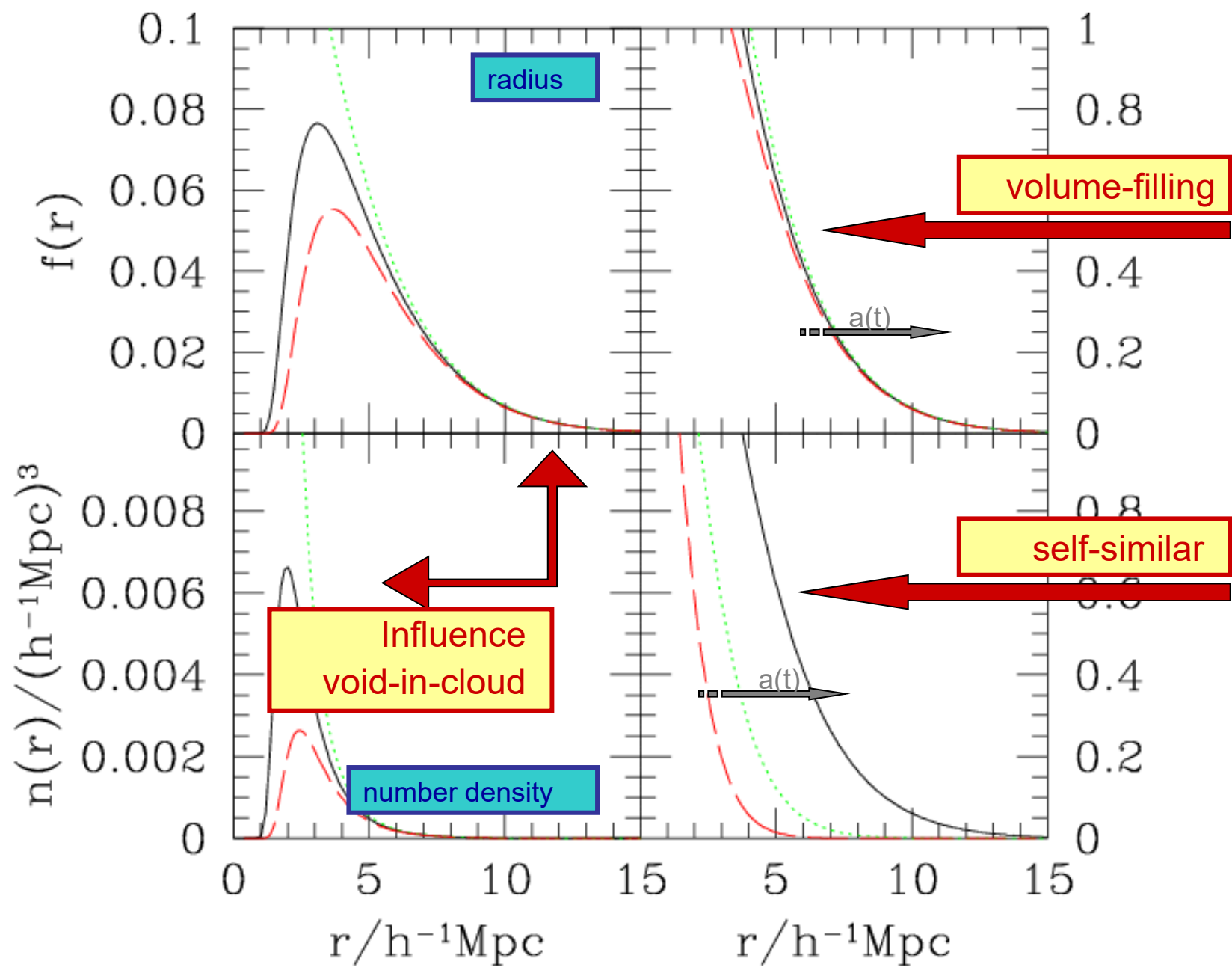
void mass distribution function

$$n_v(M) dM =$$

$$\sqrt{\frac{2}{\pi}} \frac{\rho_u}{M^2} v_v(M) \exp\left(-\frac{v_v(M)^2}{2}\right) \left| \frac{d \ln \sigma(M)}{d \ln M} \right| \exp\left\{-\frac{|\delta_v|}{\delta_c} \frac{\mathcal{D}^2}{4v_v^2} - 2 \frac{\mathcal{D}^4}{v_v^4}\right\}$$

for power-law power spectrum:

$$n_v(M) dM \approx \sqrt{\frac{1}{2\pi}} \left(1 + \frac{n}{3}\right) \frac{\rho_u}{M^2} \left(\frac{M}{M_{v,*}}\right)^{(n+3)/6} \exp\left(-\left(\frac{M}{M_{v,*}}\right)^{(n+3)/3}\right) \exp\left\{-\frac{\mathcal{D}^2}{2} \left(\frac{|\delta_v|}{4\delta_c} + \mathcal{D}^2 \left(\frac{M}{M_{v,*}}\right)^{-(n+3)/3}\right) \left(\frac{M}{M_{v,*}}\right)^{-(n+3)/3}\right\}$$



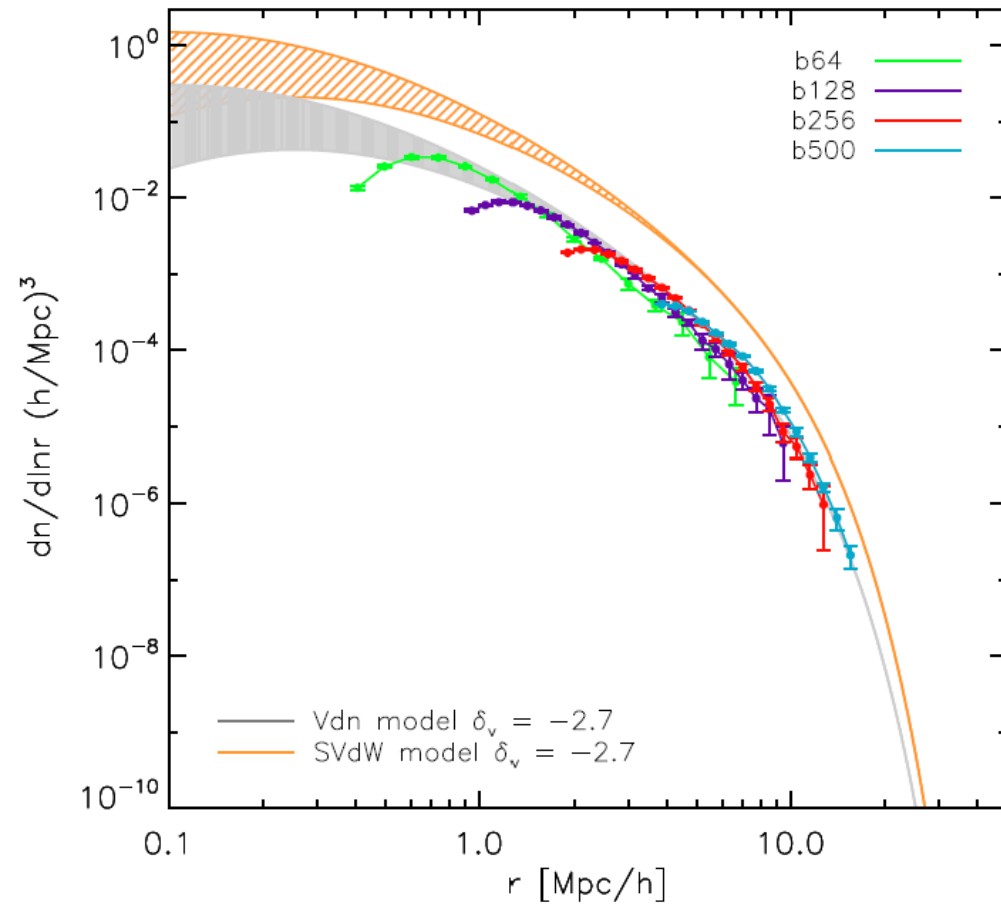
Void Population & Excursions

Jennings et al. 2013

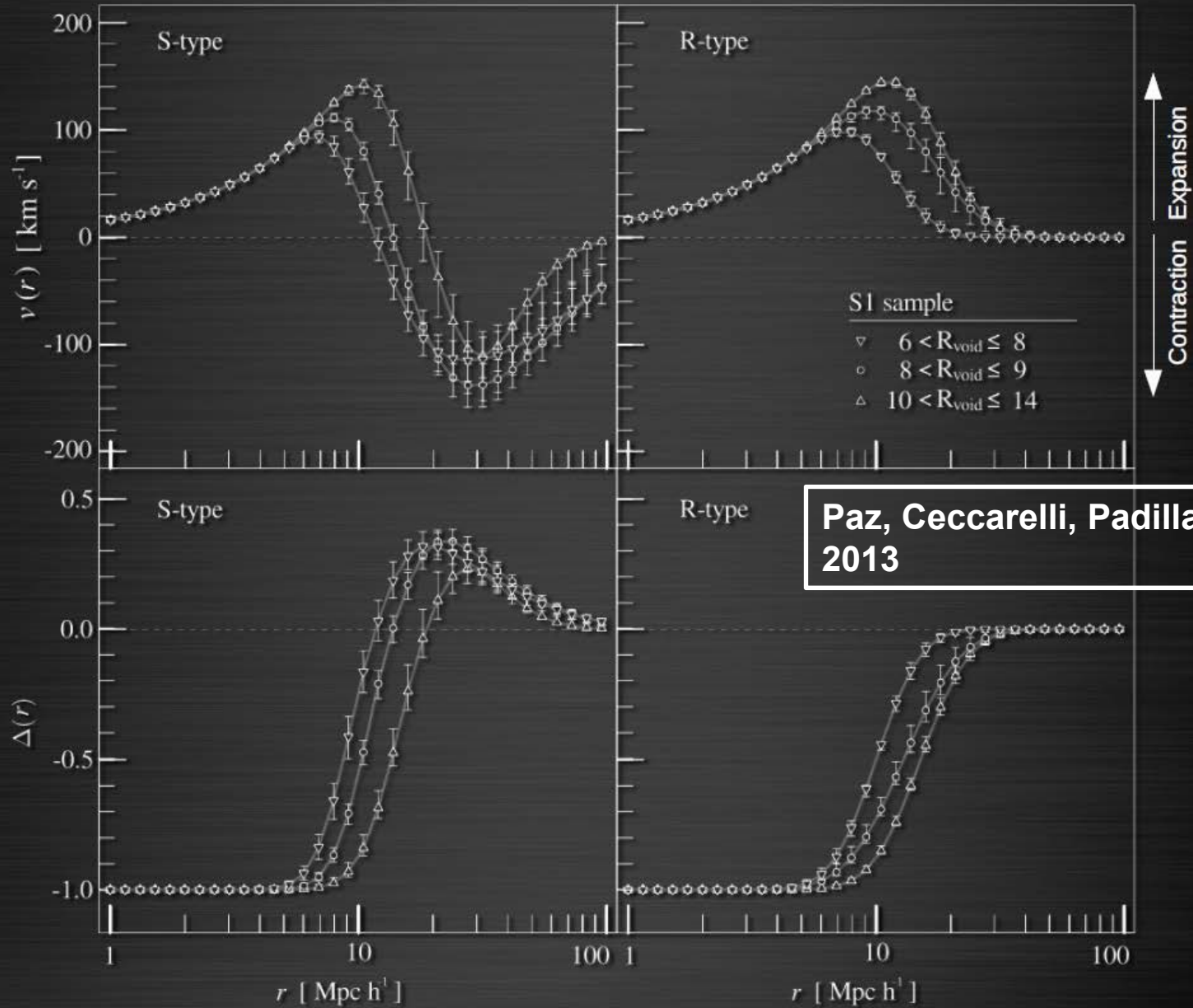
Two-barrier SvDW
void excursion set formalism:

correcting simple assumption
 δ_v criterion:

Jennings et al. 2013:
void volume occupation constraints



SDSS voids recovered velocity and density profiles

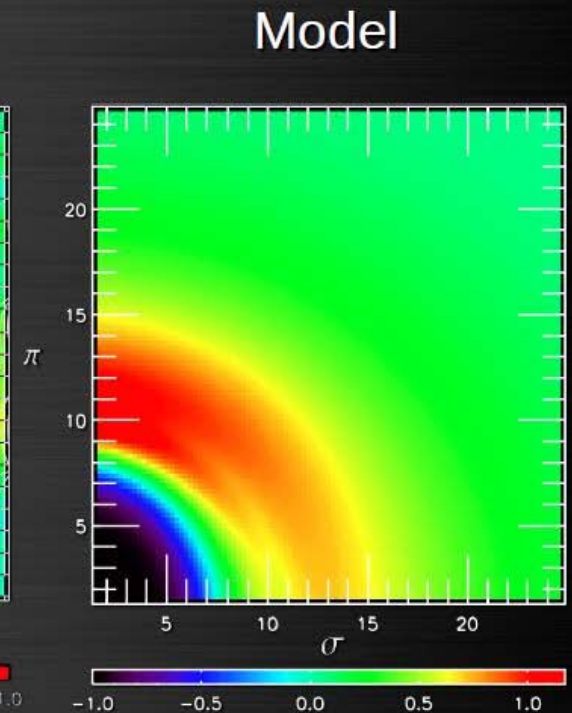
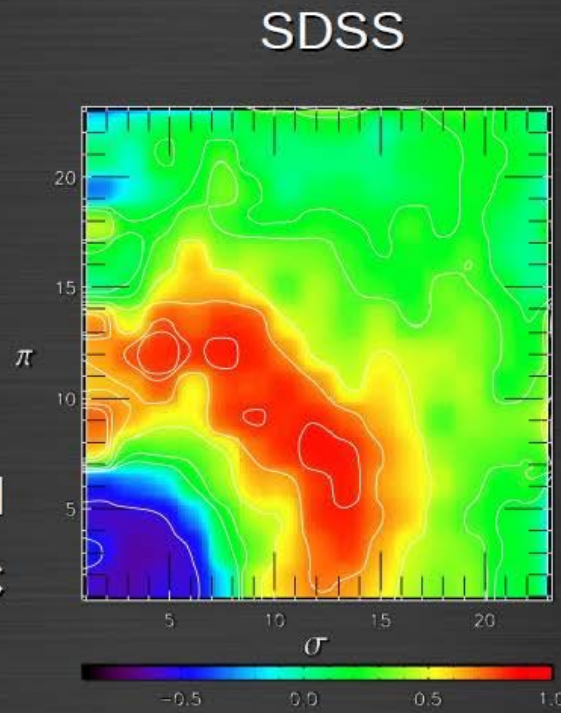


Paz, Ceccarelli, Padilla et al.
2013

SDSS void-galaxy correlations $\xi(\sigma, \pi)$

Paz, Ceccarelli, Padilla et al.
2013

Small
V-in-C



Volume Limited Sample
 $M_r < -20.3$
330 voids $Z < 0.12$

Large
V-in-V

