# Astrophysical Hydrodynamics - Assignment 7 

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## 1 Incompressible viscous flow

Consider fluid between two infinite flat plates separated by a distance $L$ (direction of $y$ ). The lower plate is stationary while the upper one is moving with a velocity $U$ parallel to itself (direction of $x$ ). The plates are also maintained at different temperatures. Neglect gravity and assume a steady flow in the x-direction, with $u=u(y), T=T(y)$. The governing equations are,

$$
\begin{gather*}
\frac{d p}{d x}-\mu \frac{d^{2} u}{d y^{2}}=0  \tag{1}\\
k \frac{d^{2} T}{d y^{2}}+\mu\left(\frac{d u}{d y}\right)^{2}=0 \tag{2}
\end{gather*}
$$

## 1. Velocity profile

If we assume that pressure gradient is imposed externally so that $d p / d x$ is a known constant, show that the velocity profile is,

$$
\begin{equation*}
u(y)=\frac{y}{L} U-\frac{L^{2}}{2 \mu} \frac{d p}{d x} \frac{y}{L}\left(1-\frac{y}{L}\right) \tag{3}
\end{equation*}
$$

[Hint: you have to use the boundary conditions: $u=0$ at $y=0$ and $u=U$ at $y=L$.]

## 2. Couette flow

If no pressure gradient is imposed on the flow then, $\frac{d p}{d x}=0$. Show that temperature profile $T(y)$ in this case turns out to be:

$$
\begin{equation*}
\frac{\left(T-T_{0}\right)}{\left(T_{1}-T_{0}\right)}=\frac{y}{L}+\frac{\mu}{2 k} \frac{U^{2}}{\left(T_{1}-T_{0}\right)} \frac{y}{L}\left(1-\frac{y}{L}\right) \tag{4}
\end{equation*}
$$

and maximum temperature occurs at

$$
\begin{equation*}
y_{m}=\frac{L}{2}+\frac{k\left(T_{1}-T_{0}\right)}{\mu U^{2}} L \tag{5}
\end{equation*}
$$

where $T=T_{0}$ at $y=0$ and $T=T_{1}$ at $y=L$.

## 3. Poiseuille flow

If the velocity of the upper plate is zero $(U=0)$ then the flow is driven solely by the pressure gradient. Show that maximum velocity in this case occurs at the center, $y=L / 2$ and it is

$$
\begin{equation*}
u_{m}=\frac{L^{2}}{8 \mu} \frac{d p}{d x} \tag{6}
\end{equation*}
$$

If mean velocity is defined by volume flow per unit area, show that in this case it is given by,

$$
\begin{equation*}
\bar{U}=\frac{1}{L} \int_{0}^{L} u(y) d y=\frac{2}{3} u_{m} \tag{7}
\end{equation*}
$$

## 2 Taylor-Couette Flow

In this exercise we will calculate the (steady state) flow of a fluid between two rotating (hollow) cylinders. The system is cylindrically symmetric and can be approximated as infinite in length. Therefore the best coordinate system is cylindrical, $r, \phi, z$. Assume the two cylinders have radii $R_{1}$ and $R_{2}$ with $R_{2}>R_{1}$ and angular velocities $\Omega_{1}$ and $\Omega_{2}$. Symmetry arguments lead us conclude that

$$
\begin{align*}
v_{z} & =0  \tag{8}\\
v_{r} & =0  \tag{9}\\
v_{\phi} & =v_{\phi}(r) \equiv \Omega r  \tag{10}\\
P & =P(r) \tag{11}
\end{align*}
$$

The cylindrical Navier-Stokes equations are written as (Landau \& Lifshitz p. 48)

$$
\begin{align*}
\frac{\partial v_{r}}{\partial t}+(\mathbf{v} \cdot \boldsymbol{\nabla}) v_{r}-\frac{v_{\phi}^{2}}{r} & =-\frac{1}{\rho} \frac{\partial P}{\partial r}+\nu\left(\Delta v_{r}-\frac{2}{r^{2}} \frac{\partial v_{\phi}}{\partial \phi}-\frac{v_{r}}{r^{2}}\right)  \tag{12}\\
\frac{\partial v_{\phi}}{\partial t}+(\mathbf{v} \cdot \boldsymbol{\nabla}) v_{\phi}+\frac{v_{\phi} v_{r}}{r} & =-\frac{1}{\rho r} \frac{\partial P}{\partial \phi}+\nu\left(\Delta v_{\phi}-\frac{2}{r^{2}} \frac{\partial v_{r}}{\partial \phi}-\frac{v_{\phi}}{r^{2}}\right)  \tag{13}\\
\frac{\partial v_{z}}{\partial t}+(\mathbf{v} \cdot \boldsymbol{\nabla}) v_{z} & =-\frac{1}{\rho} \frac{\partial P}{\partial z}+\nu \Delta v_{z} \tag{14}
\end{align*}
$$

with $\nu$ the kinematic viscosity $(\nu=\eta / \rho)$.

1. Show that under the above cylindrical symmetry assumptions the Navier-Stokes equations reduce to

$$
\begin{align*}
\frac{\partial P}{\partial r} & =\frac{\rho v_{\phi}^{2}}{r}  \tag{15}\\
\frac{\partial^{2} v_{\phi}}{\partial r^{2}}+\frac{1}{r} \frac{\partial v_{\phi}}{\partial r}-\frac{v_{\phi}}{r^{2}} & =0 \tag{16}
\end{align*}
$$

2. Solve the latter equation by assuming a power law $\left(v_{\phi}=C r^{n}\right)$ and show that the general solution for the velocity obeys

$$
\begin{equation*}
v_{\phi}=\Omega r=A r+\frac{B}{r} \tag{18}
\end{equation*}
$$

$A$ and $B$ are given by boundary conditions, what are the boundary conditions? Derive expressions for $A$ and $B$.

Such flow becomes unstable to the formation of Taylor vortices if $\Omega_{1} \gg \Omega_{2}$.
For low angular velocities the flow is steady and purely azimuthal. This basic state is known as circular Couette flow, after Maurice Marie Alfred Couette who used this experimental device as a means to measure viscosity. Sir Geoffrey Ingram Taylor investigated the stability of the Couette flow in a ground-breaking paper which has been a cornerstone in the development of hydrodynamic stability theory.
Taylor showed that when the angular velocity of the inner cylinder is increased above a certain threshold, Couette flow becomes unstable and a secondary steady state characterized by axisymmetric toroidal vortices, known as Taylor vortex flow, emerges. Subsequently increasing the angular speed of the cylinder the system undergoes a progression of instabilities which lead to states with greater spatio-temporal complexity, with the next state being called as wavy vortex flow. If the two cylinders rotate in opposite sense then spiral vortex flow arises. Beyond a certain Reynolds number there is the onset of turbulence.

We will examine a few cases of (in)stability in this flow in what follows.
3. The centrifugal force $F_{C}$ acting on the fluid particles is balanced by the pressure force $F_{P}$. Calculate the angular momentum $\mu$ of a fluid particle at position $r$ and show that the centrifugal force $F_{C}=\Omega^{2} r$ can be written as

$$
\begin{equation*}
F_{C}(r)=\frac{\mu(r)^{2}}{m r^{3}} . \tag{19}
\end{equation*}
$$

Hint: $\mu(r)=|\mathbf{p} \times \mathbf{r}|$.
4. Consider a particle initially at radial postion $r=r_{0}$ is displaced by a very small fraction to $r^{\prime}=$ $r_{0}+\delta r>r_{0}$. If we assume angular momentum conservation, the centrifugal force of the displaced particle is

$$
\begin{equation*}
F_{C}=\frac{\mu\left(r_{0}\right)^{2}}{m r^{\prime 3}} \equiv \frac{\mu_{0}^{2}}{m r^{\prime 3}}, \tag{20}
\end{equation*}
$$

but the pressure force is the same as for the other particles at that distance given by

$$
\begin{equation*}
F_{P}=\frac{\mu\left(r^{\prime}\right)^{2}}{m r^{\prime 3}} \equiv \frac{\mu^{\prime 2}}{m r^{\prime 3}}, \tag{21}
\end{equation*}
$$

Stability occurs when the forces on a the displaced particle will push it back, i.e. $F_{P}>F_{C}$. Therefore stability occurs when $\mu^{2}>\mu_{0}^{2}$.
Use a Taylor expansion

$$
\begin{equation*}
f(x)=\sum_{n=0}^{\infty} \frac{1}{n!} f^{(n)}(a)(x-a)^{n} \tag{22}
\end{equation*}
$$

of $\mu\left(r^{\prime}\right)$ around $r_{0}$ to show that the criterion can be written as

$$
\begin{equation*}
\mu \frac{\partial \mu}{\partial r}>0 \tag{23}
\end{equation*}
$$

Note that $\left(r^{\prime}-r_{0}\right)$ is a positive difference.
5. Evaluate this expression to

$$
\begin{equation*}
\left(\Omega_{2} R_{2}^{2}-\Omega_{1} R_{1}^{2}\right) \Omega>0 \tag{24}
\end{equation*}
$$

( $\Omega=v_{\phi} / r$ ).
Hint 1: evaluate $\partial_{r} \Omega$ and $\partial_{r} \mu$.
Hint 2: always-positive terms are always positive.
6. Argue for each of the following cases whether the flow is stable or instable.

1. The cylinders rotate in a different direction
2. The cylinders rotate in the same direction
(a) The center cylinder is stationary
(b) The outer cylinder is stationary
