

# Degeneracy pressure: an alternate derivation

An addendum to “Stellar Structure and Evolution,” part  
of “Sterrenstelsels en Kosmos”

# Degeneracy pressure for non-degenerate electrons

- To begin, we need to discuss the concept of “dynamical phase space”
- In dynamics, our interest is in the *positions* and *momenta* of particles
- We can imagine a 6-dimensional “hyperspace” that tells us the position and momentum of each particle we’re following:  $(x, y, z, p_x, p_y, p_z)$

- Now, because of the Pauli exclusion principle, only *two* electrons can occupy some volume of phase space
- two because each electron has the possibility of having spin= $1/2$  or spin= $-1/2$
- The size of this volume in phase space is given by the Heisenberg uncertainty principle:  $\Delta x \Delta p_x \geq h/2\pi$
- so the minimum volume per particle must be  $\Delta x \Delta y \Delta z \Delta p_x \Delta p_y \Delta p_z \sim h^3$

- Now, the *number density of electrons* in real space is  $n_e = 1/(\Delta x \Delta y \Delta z)$
- If all the electrons have some *characteristic momentum*, which we'll call  $p_0$ , then we can say that there are two electrons per allowable volume of *phase space*,  

$$p_0^3/n_e \sim h^3/2$$
- or  

$$n_e/p_0^3 \sim 2/h^3$$

- Of course, the number density of electrons is the total density divided by the mass of available electrons,  $n_e = \rho / (\mu_e m_p)$
- here  $\mu_e$  is the number of electrons per nucleon and  $m_p$  is the proton mass
- so the momentum can be found from (ignoring factors of order unity)
 
$$p_0^3 = h^3 \rho / (\mu_e m_p)$$
- and therefore
 
$$p_0 = h [\rho / (\mu_e m_p)]^{1/3}$$

- Now, the *pressure* is the just the energy density (again ignoring term of order unity)

$$P = \rho v^2$$

- but remember that the velocity of the electrons is just  $v = p_0/m_e$

- On substituting, we find  $P \propto \rho p_0^2$

- and therefore, substituting in the momentum,  $P \propto \rho \rho^{2/3}$

- So finally, we have the desired relation:

$$P \propto \rho^{5/3}$$