

## VIRTUAL OBSERVATIONS

Lecture Course for Master and last year Bachelor Students

Kapteyn Astronomical Institute  
University of Groningen  
27 Nov 06 – 2 Feb 07

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In collaboration with AstroGRID, CDS, LOFAR, Astro-Vilse, OmegaCEN  
Part of Master-Variant Instrumentation and Informatics in Physics, Astronomy and Space Research

Information: <http://www.astro.rug.nl/~valentyn/vo.html>

## This course

- A novum
- Programme
  - <http://www.astro.rug.nl/~valentyn/vo.html>
  - Basis
    - intro
    - concepts – systems
    - Astrometry – photometry
  - CDS – VO
    - EURO-VO tools
  - E – Lofar
  - AstroWise
  - AstroGRID
  - Students VO

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## E-science

- Beyond “workstation science” of the 80- 90’s
- Distributed services
- Distributed communities
- Distributed archives
- p2p networks – KAZAA- NAPSTAR
  - Share cpu
  - Share storage
  - Share info / meta data /knowledge

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## different views on

- Surveys
- Templates
- Pipelines
- Virtual Observatories

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## Surveys

- Defined area on sky
- Homogeneous
  - Survey limit
    - Flux (magnitude)
    - Size
    - Surface brightness
    - distance
- Quality control

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## Templates

Standards very important for VO

- Observing templates
  - Astronomical Observing Templates at ESA
  - Templates / Template signature files at ESO

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## ESO parse info via headers

```

Photometry
- TSF- OCAM_img_obs_star
  TSF- OCAM_img_cal_monit
  Stars observations with N = 1, fiber/composite
  Recipe- PhotCal_Monitoring
  TPL ID = OCAM_img_cal_monit
  TPL NSFP = 1
  DPR CATG = 'CALIB'
  DPR TYPE = 'STD, EXTINCTION'
  DPR TSCH = 'IMAGE, DIRECT'
  DPR STRLTO = 'FREQ'
- TSF- OCAM_img_cal_sp
  Stars observations with N = 1, fiber/key band
  Recipe- PhotCal_Extract_ZeroPoint
  TPL ID = OCAM_img_cal_sp
  TPL NSFP = 1
  DPR CATG = 'CALIB'
  DPR TYPE = 'STD, ZEROPOINT'
  DPR TSCH = 'IMAGE, DIRECT'
  DPR STRLTO = 'FREQ'
  INS FILT1 = 'u', 'g', 'r', 'i', 'z', 'composite'
    
```

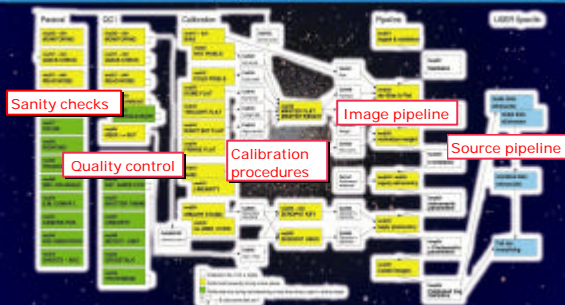
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## pipelines

- Workflow
- What triggers a pipeline?
  - Data items
  - Operators
  - users

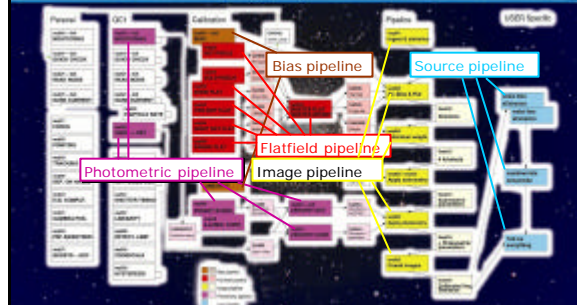
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## Data Model / flow



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## Astro-Wise Pipelines



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## Virtual observatories

• ????????????

[GigaPort seminar for astronomers](#)

link to VO

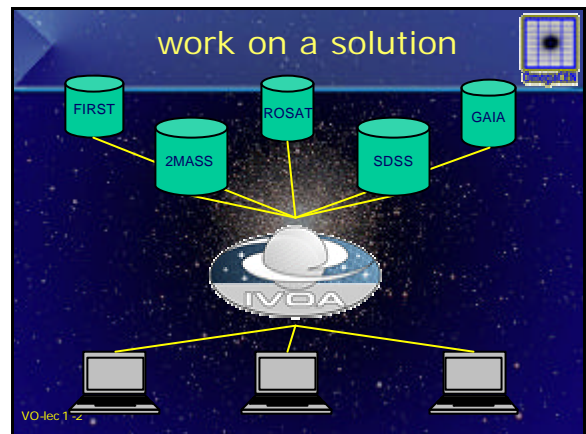
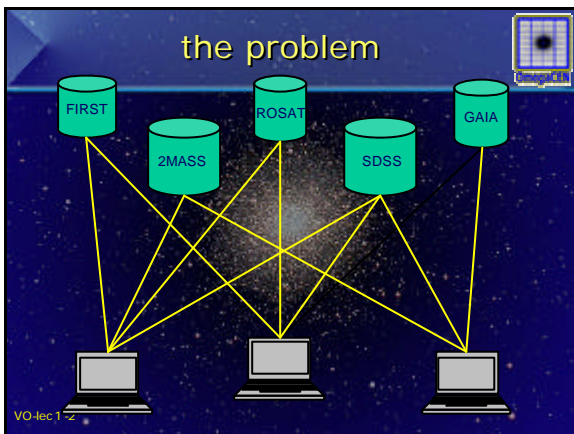
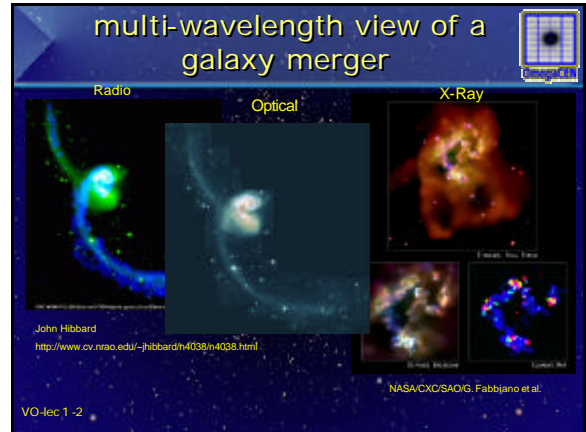
lecture by EV

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## Virtual observatories

- Broad VOs
  - IVOA
  - Euro VO
- EuroVO DataCenter Alliance
- Focussed VOs
  - AstroWise

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- ### AVO-> Euro-VO
- Standards
    - FITS
    - Universal Content descriptor UCD
    - VO table - VOT
  - Communities - workshops/training
  - Registry
  - Connecting archives
    - Cone search
    - X-match
    - All kinds of tools/ web services
  - Relatively static
- VO-lec 1 -2

### standards: Universal content Descriptor - UCD

- [http://www.ivoa.net/Documents/RFC/UCD/UCD\\_20050812.html#ToC8](http://www.ivoa.net/Documents/RFC/UCD/UCD_20050812.html#ToC8)

Examples of UCDs and how they are built:

- The maximum temperature of an instrument. This is a temperature, so the primary word will be `phys.temperature`. This temperature is that of an instrument, so we specify `inst`: `phys.temperature:inst`. And finally, we add a third word to indicate that this is the maximum value of a `phys.temperature:inst`, giving the final UCD: `phys.temperature:inst:ofmax`
- The name of a magnetic field in the F band. The quality is an uncertainty, so the primary word will be `star.name`. This uncertainty applies to a magnitude, so we will be `star.mag`: `star.mag`. Thus, we can specify the photometric band with another word, giving `star.mag:phot:band`: `star.mag:phot:band`

<http://www.ivoa.net/Documents/PR/UCD/UCDlist-20061006.html>

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# Standards: VO table VOT

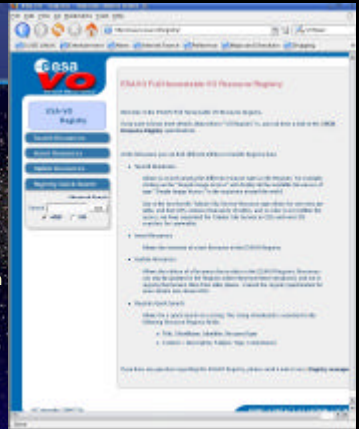
- [VOTable\\_Sourcelist.xml](#)
- <http://www.ivoa.net/twiki/bin/view/IVOA/IvoaVOTable>

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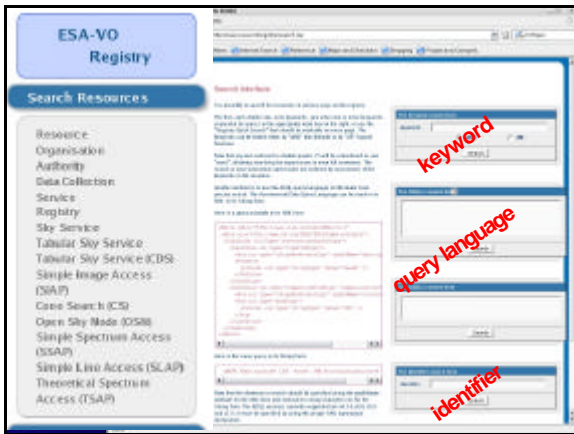
# Finding information in the VO

## Registries are here

- multiple interfaces
  - human readable
  - machine readable
- simple/advanced search

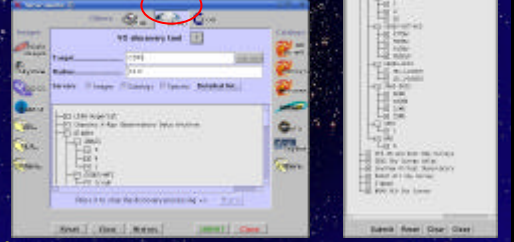


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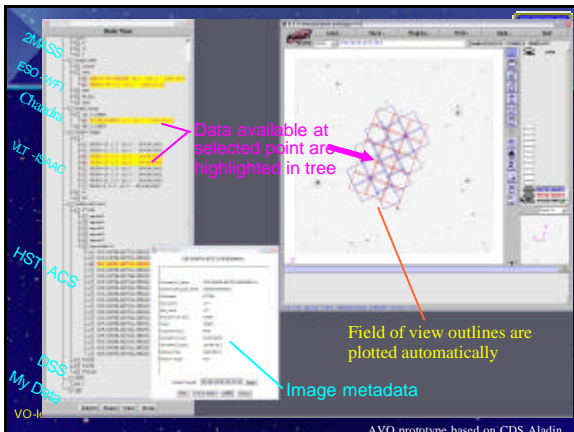


# Examples

CDS Aladin: combines registry discovery with other access methods



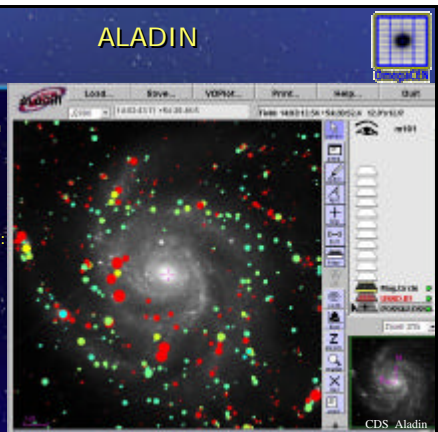
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AVO prototype based on CDS Aladin

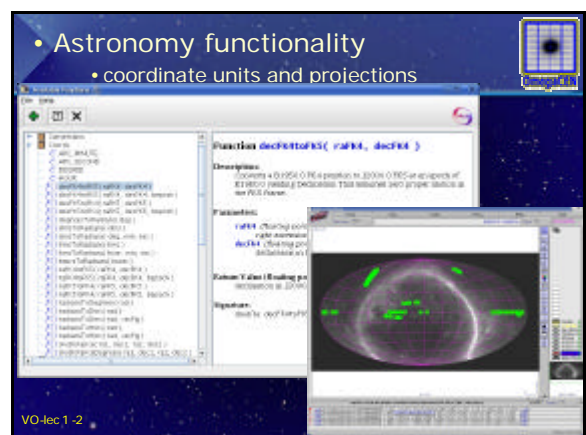
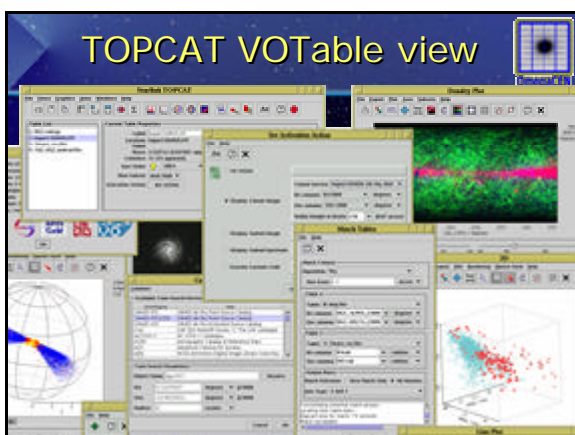
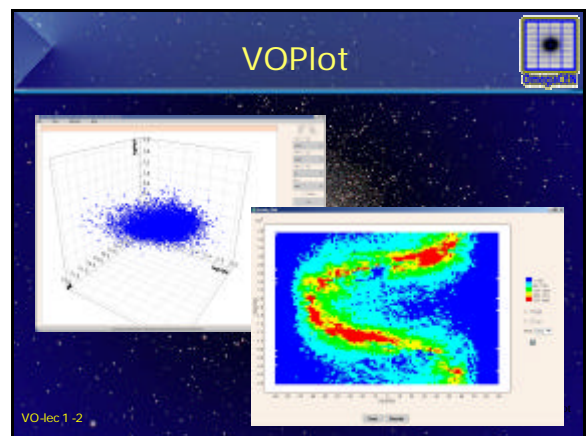
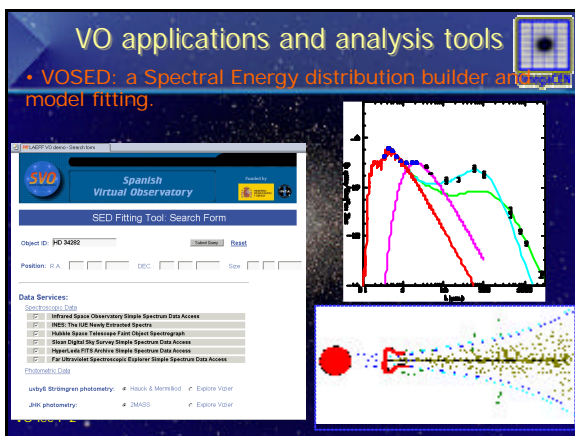
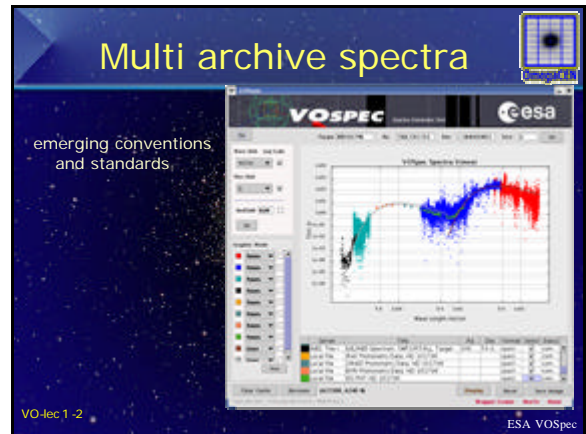
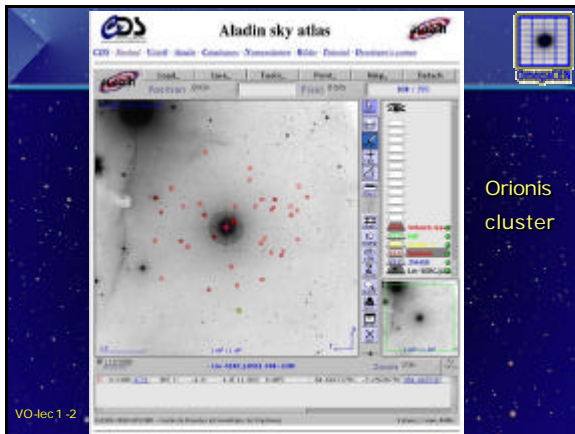
# ALADIN

- Manipulation
- X-match
- Visualization
- Direct links to:



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CDS Aladin



### Astronomical archives: take-up of VO standards.

The slide shows three screenshots: the top left is the INES (International Network of Extraordinary Sites) website; the top right is a GEM (Global Earthquake Model) interface; the bottom left is another GEM interface; and the bottom right is a 3D rendering of a large astronomical telescope structure.

### Theory in the VO

- Theoretical model web server + TSAP

The slide displays two screenshots: on the left, a web server interface with a search bar and a list of items; on the right, a VOSPEC (Virtual Observatory Spectroscopy) plot showing a spectral energy distribution with a peak in the infrared region.

### ASTRID -CM: VO Science

- From Class 0 to Class III: Search and classification

The slide shows several screenshots: on the left, a search results page for ASTRID; in the center, a 3D visualization of a protoplanetary disk; on the right, two spectral plots showing flux versus wavelength.

### SVO Thematic Network: VO Science

- Discovery of ultracool (T-type) brown dwarfs (IAC+LAEFF)

The slide displays several screenshots: on the left, a search results page for SVO; in the center, a 3D visualization of a protoplanetary disk; on the right, two spectral plots showing flux versus wavelength.

### applications

- Astro-GRID → Euro-VO
  - Workflow
    - Taverna 2 (Bioinformatics-life sciences) client side workflow editor and engine
  - Connecting webservices – grid
  - my space

### Euro-VO connect and mine archives

The diagram illustrates the Euro-VO architecture. At the top is the 'EURO VO Data centers take-up' box, which contains the 'DCA' (Data Center Architecture) component. Below this are three main components: 'VOTC' (Virtual Observatory Technology Center) on the left, 'DCA' in the center, and 'VOFc' (Virtual Observatory Facility) on the right. Arrows indicate the flow of information: 'STANDARDS, TOOLS, SYSTEMS' flow from VOTC to DCA and from DCA to VOFc; 'REQUIREMENTS' flow from VOFc to DCA and from DCA to VOTC.

## Data Center Alliance

- Consortium Agreement
  - F, I, G, UK, Sp, NL, ESO, ESA
- Now EU funded: FP6 1.5 M Euro
- COMMUNICATION NETWORK DEVELOPMENT COORDINATION ACTION
  - start 1/9/2006 2.5 year (startup project)
  - ~ 1 fte NL, + travel
- NOVA -NL

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## DCA tasks – WPs take-up VO technology

- WP3: support of take-up and implementation of VO technology
  - NL: AstroWise/OmegaCAM, Lofar
  - Census
- WP 4: inclusion theoretical data VObs
- WP 5: coord with computational GRIDS
  - NL: I-Grid- F.Pasian
- WP 6: support other EU countries

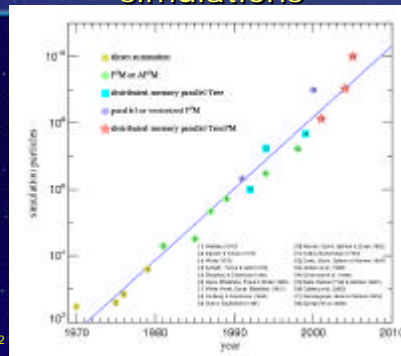
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## Theory in the VO: issues Gerard Lemson

- Simulations not so simple
  - complex observables
  - no standardisation (not even HDF5)
  - archiving ad hoc, for local use
- Moore's law makes useful lifetime relatively short: few years later can do better
- Current IVOA standards somewhat irrelevant
  - no common sky
  - no common objects
  - requires data models for content, physics, code

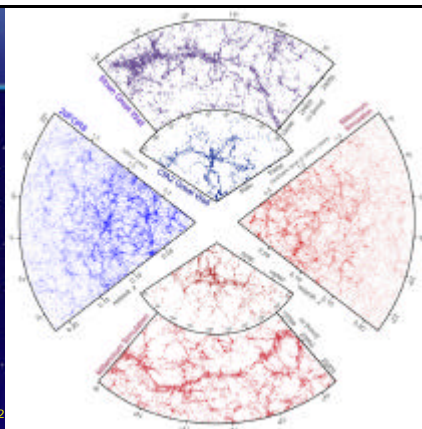
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## "Moore's law" for N-body simulations



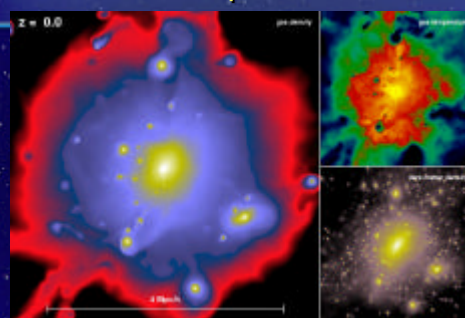
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Courtesy Simon White



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## Detailed predictions

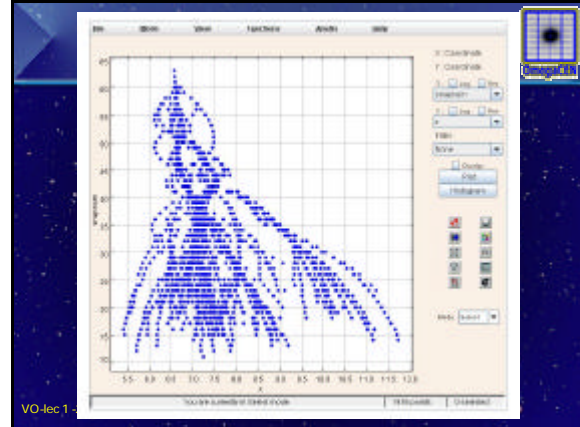


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## The Virgo consortium's Millennium simulation

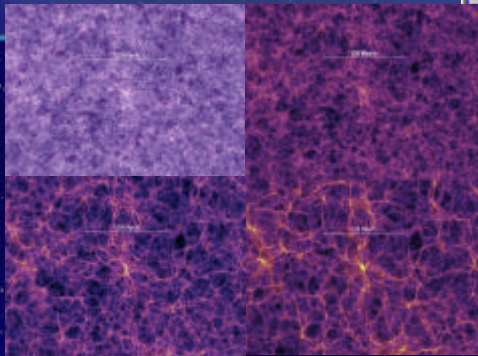
- Millennium simulation
  - 10 billion particles, dark matter only
  - 500 Mpc (~2Gly) periodic box
  - "concordance model" (as of 2004) initial conditions
  - 64 snapshots
  - 350000 CPU hours
  - O(30Tb) raw + post-processed data
- Postprocessing:
  - dark matter density fields smoothed at various scales (45 \* 256<sup>3</sup> grid cells)
  - dark matter cluster merger trees (~750 million)
  - galaxy merger trees (~1 billion/catalogue)
    - DeLucia & Balzot, 2006
    - Bower *et al*, 2006

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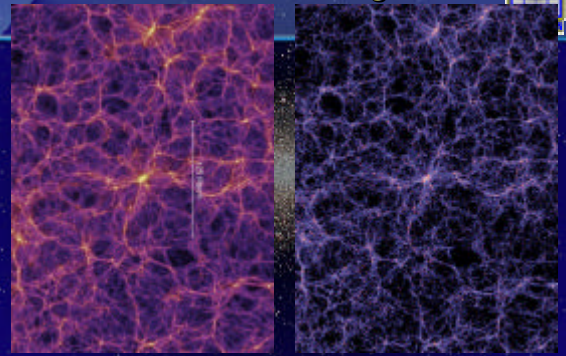
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## Evolution



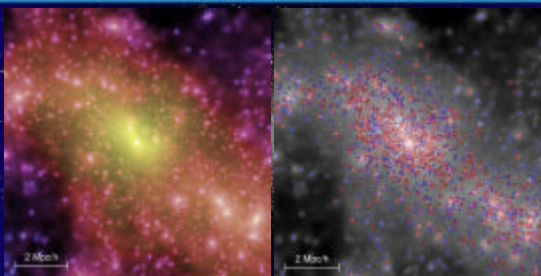
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## Dark matter and galaxies



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## Halos and galaxies



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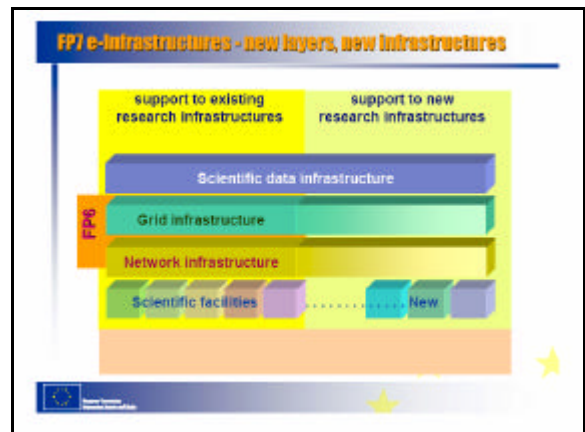
## the Millennium database + web server

- Post-processing results only
- SQLServerdatabase
- Web application (Java in Apache tomcat web server)
  - portal: <http://www.mpa-garching.mpg.de/millennium/>
  - public DB access: <http://www.g-vo.org/Millennium>
  - private access: <http://www.g-vo.org/MyMillennium>
  - MyDB
- Access methods
  - browser with plotting capabilities through VOPlot applet
  - wget + IDL, R
  - TOPCAT plugin

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## The observer >2006

- End-user, astronomer, research
  - Statistical approach
  - Individual result
- Overload of data:
  - Hubble ACS: 16 Mpix/image
  - Gaia: 10<sup>9</sup> stars
  - Ground based: OmegaCAM, MegaCAM 256Mpix/image → 10<sup>8</sup> galaxies/image → 100 Terabyte - Petabyte regime
- Where do we stop? How do we handle this?
- New approaches- paradigm- relation to system design

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## Lofar

IBM- Blue Gene/L



## from Design-> deliver

- Scientific requirements - SRD
- User requirements - URD
- Architectural design - ADD
- Detailed design - DDD
- Implementation
- Quantify
- Build
- Qualify – unit tests

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## New approaches new balances

Anarchy  $\leftrightarrow$  coordinated  
Freedom  $\leftrightarrow$  fixed system

Standard data products  $\leftrightarrow$  user tuned products  
Data releases  $\leftrightarrow$  user defined hunting

DESIGN

5 Essential STEPS:

## 1- calibration plan integrated up-link /down link

NOVA - Kapteyn Institute USM - OUPJ	<b>OmegaCAM</b> DFS	ID Issue Date Page	internal Version 2.11 7 Oct 2004 42
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**Selected Items from Odooc file system**

- Summary actions
- Issue
- Objective
- Fulfilling or fulfilled by
- When performed/frequency
- Source, obs...
- Inputs
- Outputs
- Required accuracy, constraints
- Estimated time needed
- Priority
- Template/Signature File
- Recipe
- Calibration Analysis spec's
- Needed Functionality
- CA implementation (pseudo code)
- Status of Item
- FLAG

www.astron.org.uk/~omegacam

VO-lec 1

NOVA - Kapteyn Institute USM - OUPJ	<b>OmegaCAM</b> DFS	ID Issue Date Page	internal Version 2.11 7 Oct 2004 42
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**A.2.1 Req - Cal F CCD read noise - cal**

**Objective:**  
Measure the CCD read noise (in ADU/e) as a standard health check. The read noise is measured from pairs of star exposures. The rms value of the difference between two exposures is computed and divided by  $\sqrt{2}$  to obtain read noise. This is the first and only daily health check.

**Fulfilling or fulfilled by:**  
Self-scheduling

**When performed/frequency:**  
Always. Contingency: during all emergency daily health checks.

**Required accuracy, constraints:**  
Readout noise less than 5e  
Validation or readout noise at 1 reference value less than 0.5e  
These are lab values. The corresponding limits in ADU can be calculated using the  $\sigma$  (ADU conversion factor from req.503)

**Estimated time needed:**  
Observation: 3 min. Reduction: 3 min (CPU)

**TSP:**  
Module: Stars (cal)  
(TSF - OCM\_img\_cal\_obs, N-1)  
TSF - OCM\_img\_cal\_readnoise

**Inputs:**  
3 star list frames

**Outputs:**  
CalFile-528: 2x double array in ADUs  
The CalFile corresponds to QC parameter read\_noise (a single number)

**Produce:**  
read\_noise (1) Read Noise [e rms] MAGCAM\_IMAGE[0] [e rms] MAGCAM\_IMAGE[0]

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NOVA - Kapteyn Institute	OmegaCAM	ID	internal
USM - GaPd	DFS	Issue	Version 2.11
		Date	7 Oct 2004
		Page	49

```

bias1, bias2      : the two raw bias images
MAXIMUM_ITERATIONS : maximum number of iterations for statistics
                    measurements (integer)
                    Range of allowed values: 2 - 10. Default: 5
REJECTION_THRESHOLD : rejection threshold for bad pixels in sigma (float).
                    Range of allowed values: 1.0 - 10.0. Default: 5.0

Before applying this recipe, use Recipe: Split—which is documented in seq-631—with the -t bias option to split the raw
multi-extension FITS input files.

CA:
Process (make):
1. Subtract the two bias frames to produce a difference image.
2. Iteratively reject outliers in the difference image.
3. Compute the mean, median and rms of the remaining pixels of the difference image.
4. The readout noise is the rms divided by sqrt(2).

CAP:
diff_image = eclipse.image_sub(bias1, bias2)
stats = eclipse.isor_stats(diff_image,
                           MAXIMUM_ITERATIONS,
                           REJECTION_THRESHOLD)
read_noise = stats.stdev / sqrt(2)
mean_diff = stats.avg_pix
median_diff = stats.median

QC Parameters:
read_noise : the readout noise in ADU
mean_diff  : mean difference between the two input raw bias frames
median_diff : median difference between the two input raw bias frames

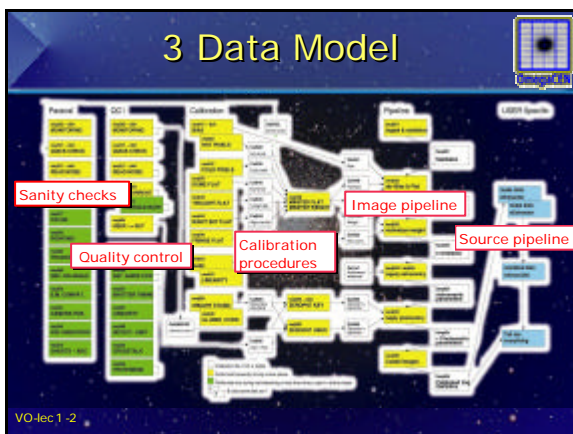
```

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## 2 -Procedurizing

- **Procedurizing**
  - Data taking at telescope for both science and calibration data - Templates
    - Observing Modes: ~~Stare~~ ~~Jitter~~ ~~Dither~~ ~~SSO~~
    - Observing Strategies: ~~Shan~~ ~~Deep~~ ~~Freq~~ ~~Mosaic~~
  - Full **integration** with data reduction
  - Design- **ADD**
  - **Data model** (classes) defined for data reduction and calibration
  - View pipeline as an **administrative problem**

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## 4 Integrated archive and Large Data Volume

- **Handling of the data is non-trivial**
  - Pipeline data reduction
  - Calibration with very limited resources
  - Things change in time:
    - Physical changes (atmosphere, various gains)
    - Code (new methods, bugs)
    - Human insight in changes
  - Working with source lists

**Science can only be archive based**

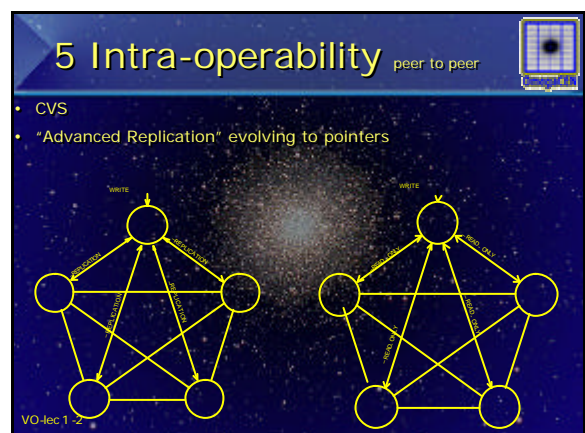
VO-lec 1 -2

## 4 archive and Virtual Survey System

- Environment that provides systematic and controlled
  - Access to all raw and calibration data
  - Execution and **modification** image/calibration pipelines
  - Execution of source extraction algorithms- **catalogues**
  - Archiving or regenerate **on-the-fly dynamically**
  - Paradigm: no static data releases but **dynamic on request data**
  - **federated** to link different data centers
- Dynamical archive continuously grows, can be used for
  - small or large science projects
  - generating and checking calibration data
  - exchanging methods, scripts and configuration

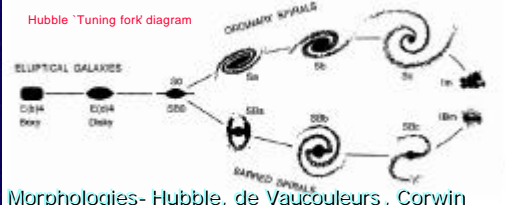
**raw pixel data ⇔ pipelines/cal files ⇔ catalogues**

VO-lec 1 -2



## Case: morphologies parameters vs theory -degeneracy

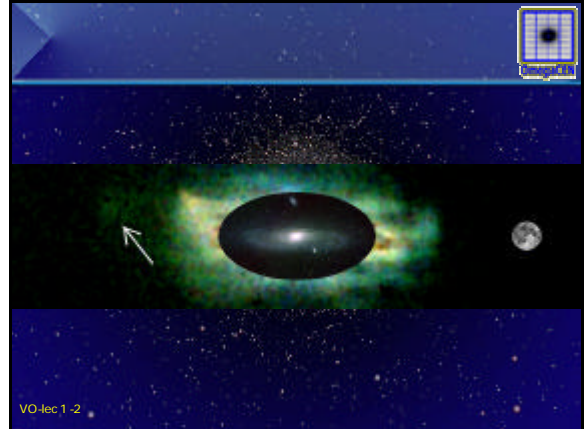
Hubble Tuning fork diagram



Morphologies- Hubble, de Vaucouleurs, Corwin

- U, B, V, R, I, Z, K
  - m\_tot, m\_25, m\_26, D\_26, D\_25, D\_90%, r\_eff, sb\_0, sb\_eff
- Structure/radial
  - a/b, pa, B/D, N, exp scale length, delta (B-R)
- colours
  - Central (B-R), (B-R)\_tot,,

VO-lec1-2



VO-lec1-2