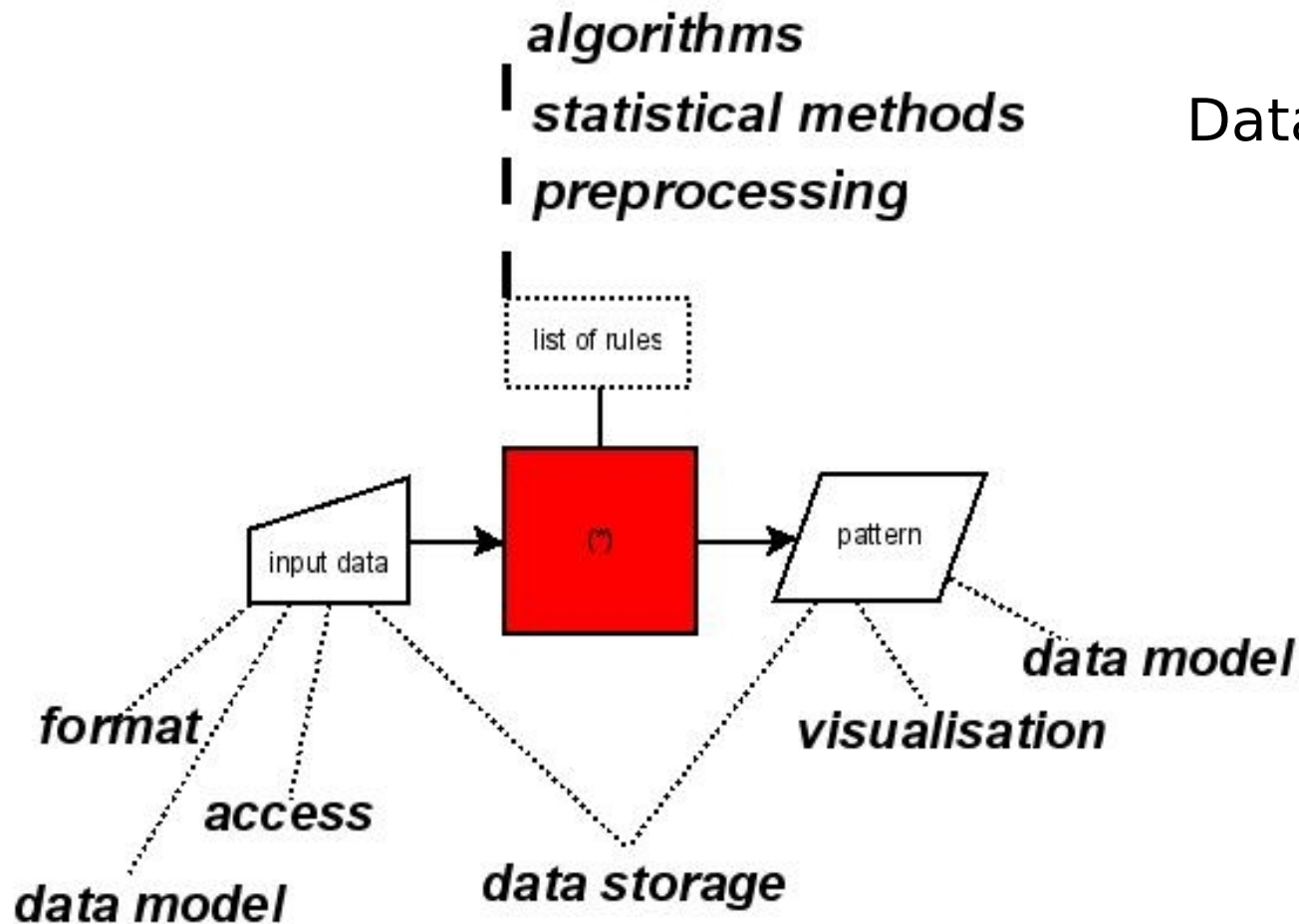


Scientific Paradigm

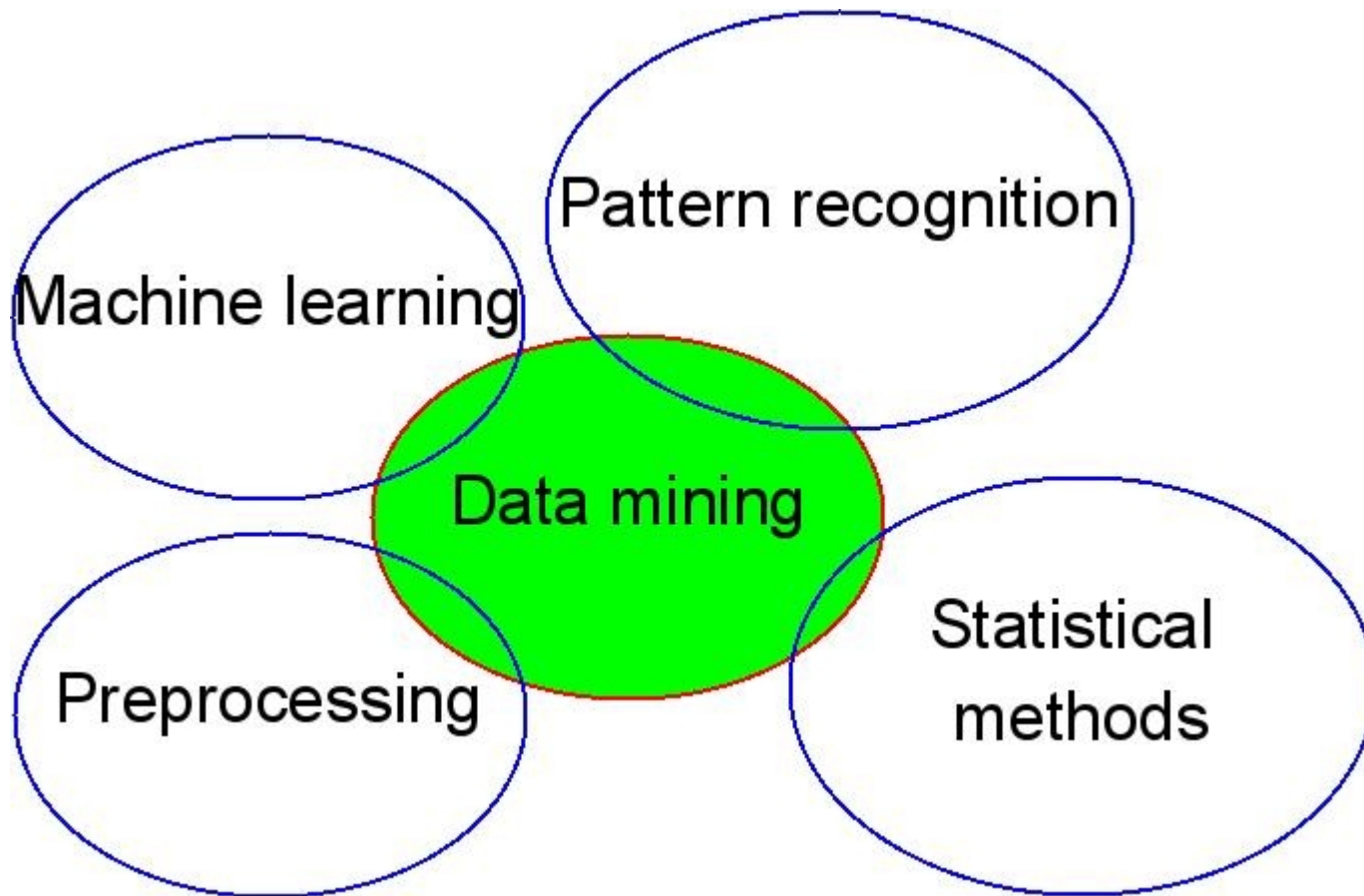
- First : Observations
- Second : Theory
- Third: Computer simulations
- Fourth: Data mining

Data Mining

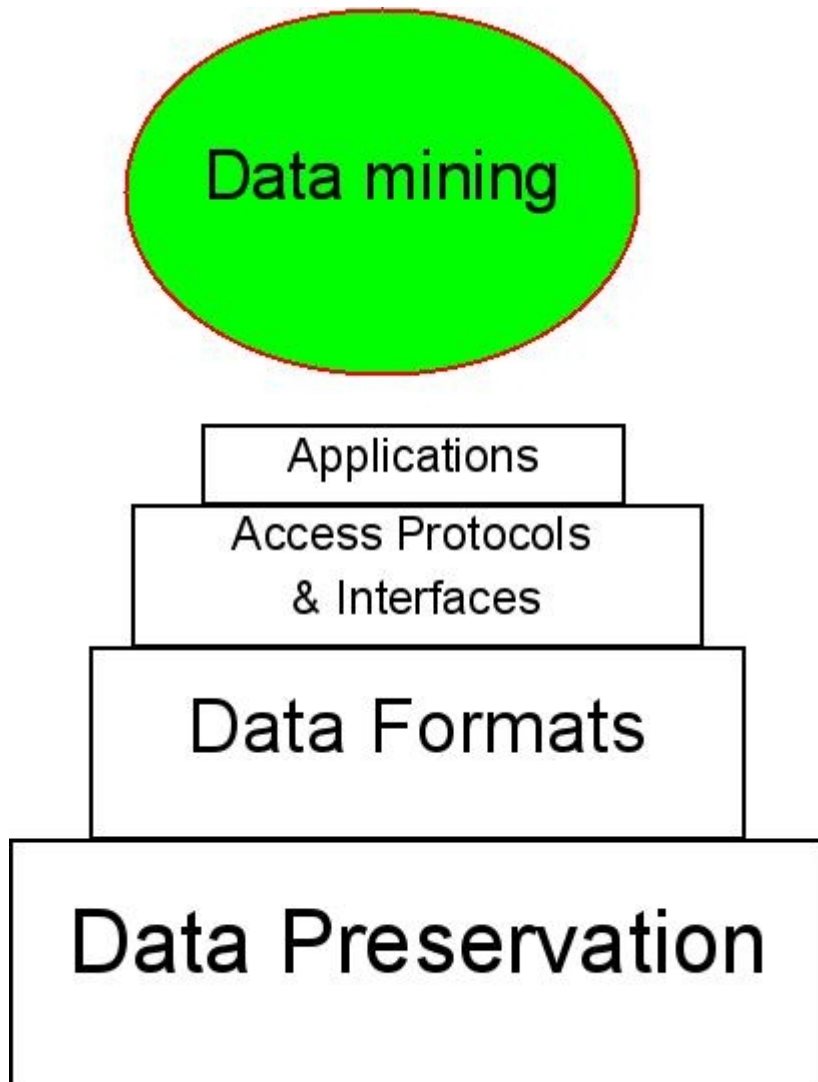


Database: Data format
Access
Data model
Preprocessing
Processing

Data Mining: Connection

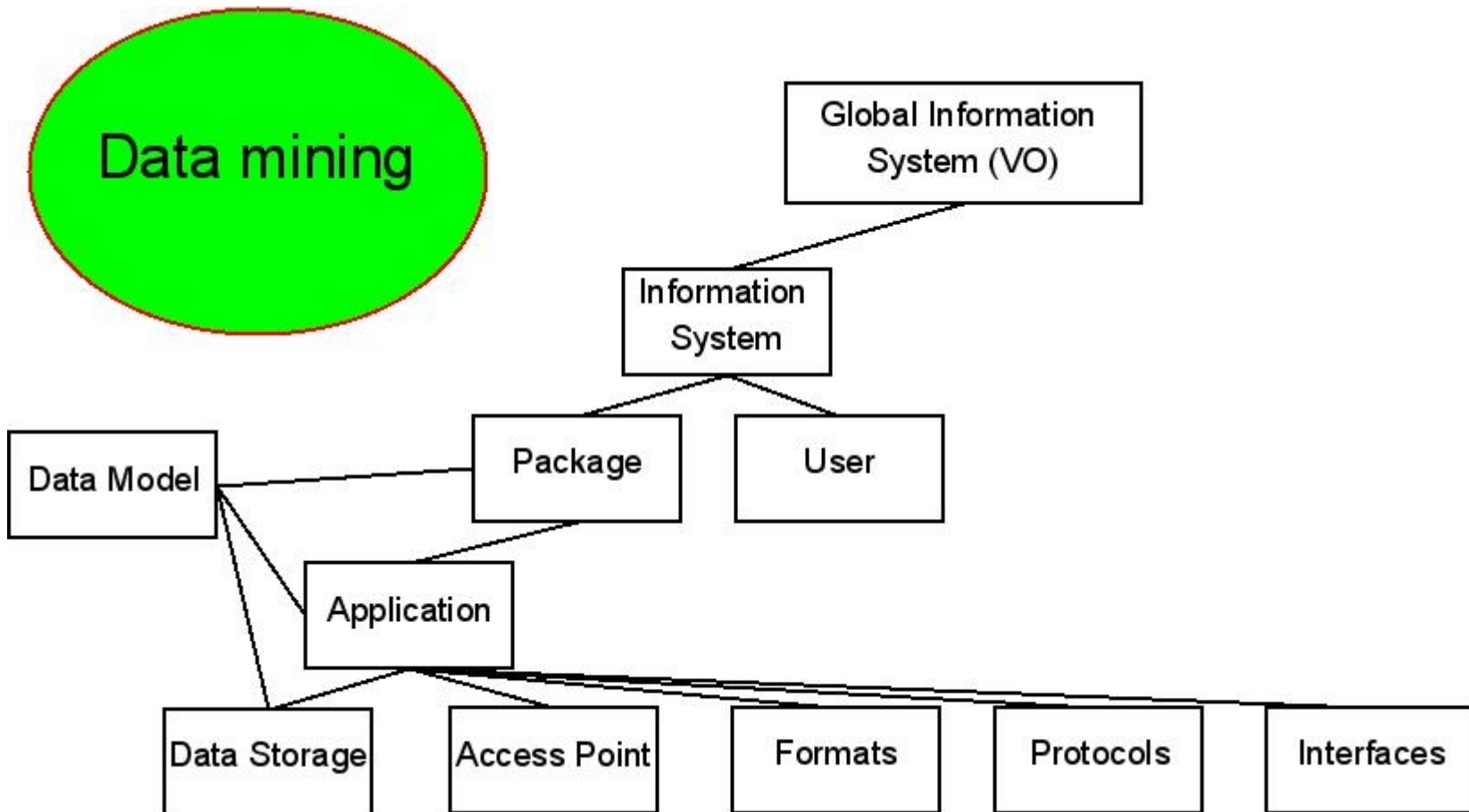


Data Mining: Concept



- Data Preservation: OAIS
- Data Formats: FITS + metadata
- Access & Interfaces: http, web interfaces, XML
- Applications

Data Mining: Implementation



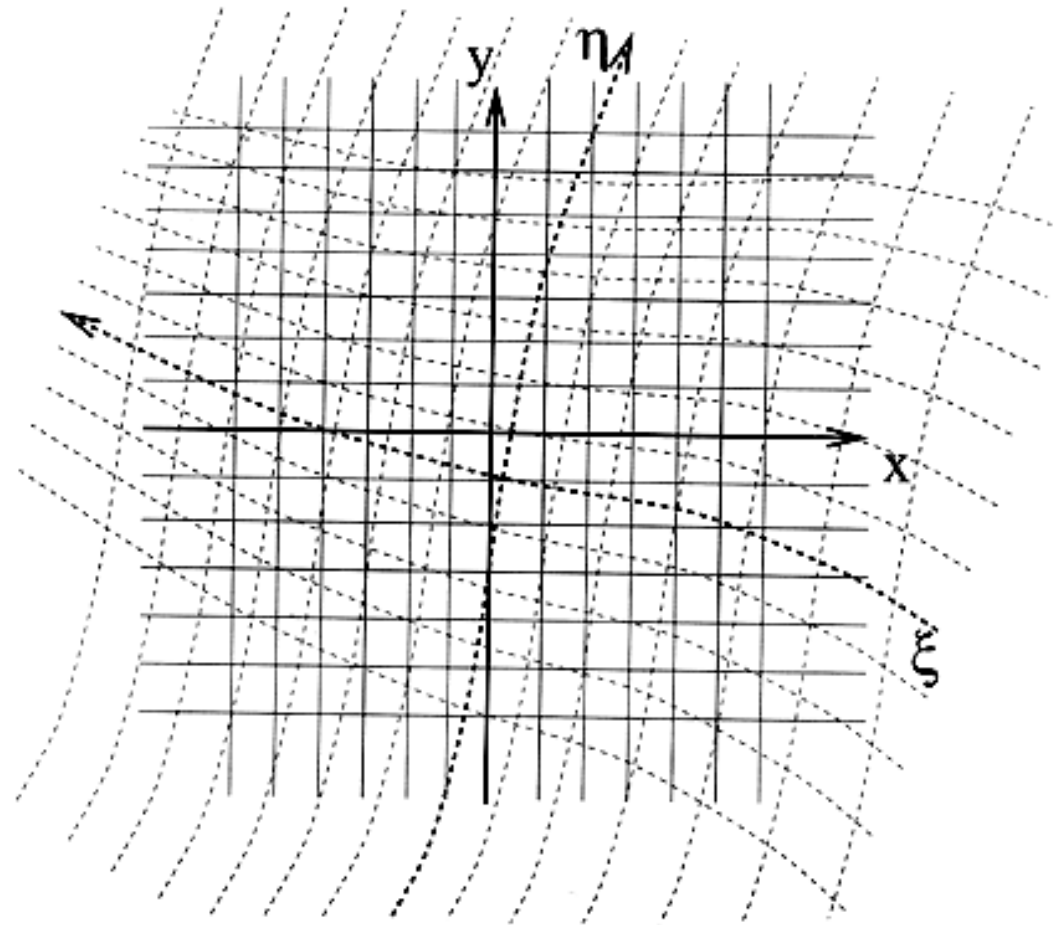
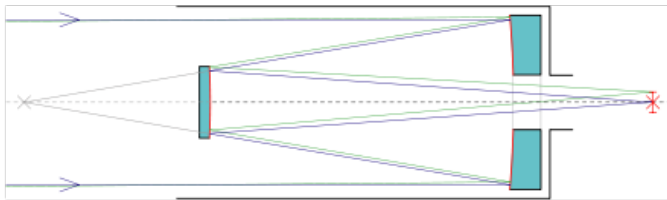
Data Mining in Astronomy

Use cases

- Cross-identification
- Parametric search
- Pattern recognition
- Coordinate transformation
- Source extraction

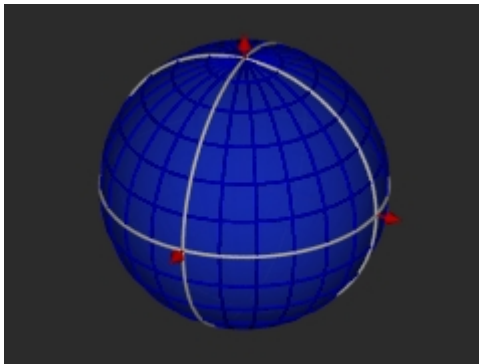
Cross-identification

- Different coordinates
- Different epochs
- Reference system

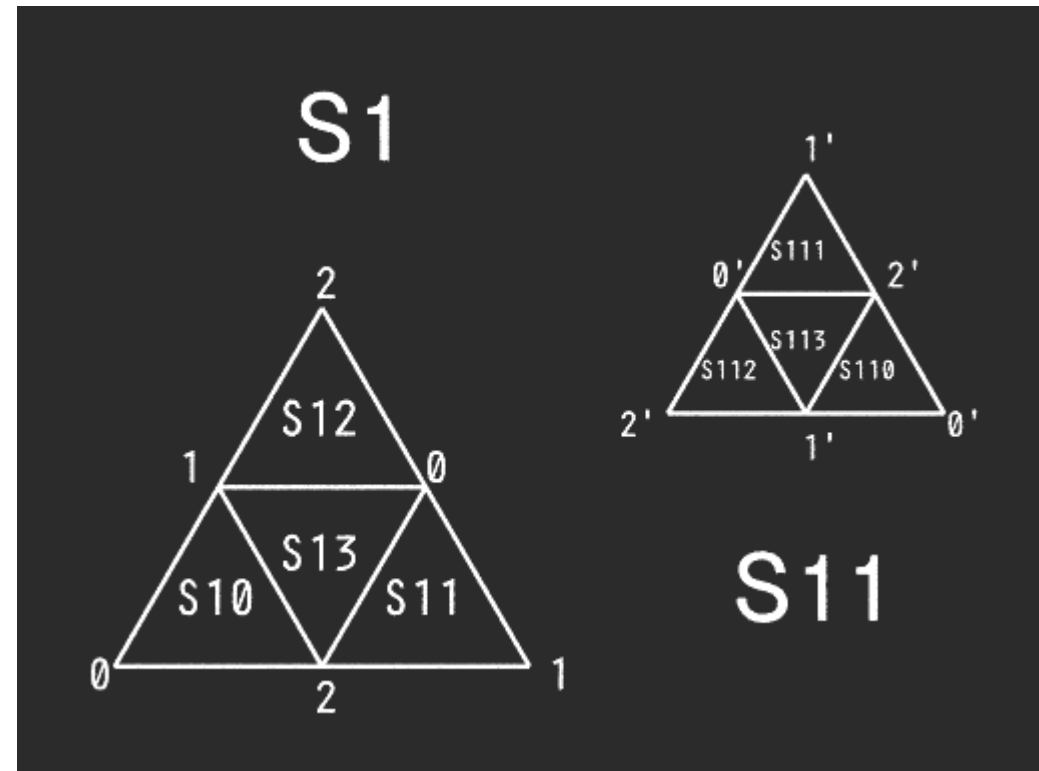


Cross-identification: HTM

<http://skyserver.org/htm/>



- Time consuming
- Distance



Parametric search

- Photometric and coordinate search
- Object parameters (shapes, classifiers)
- SDSS: benchmark of 20 “typical” queries

Search for Cataclysmic Variables and pre-CVs with White Dwarfs and very late secondaries:

$u-g < 0.4$

$g-r < 0.7$

$r-i > 0.4$

$i-z > 0.4$

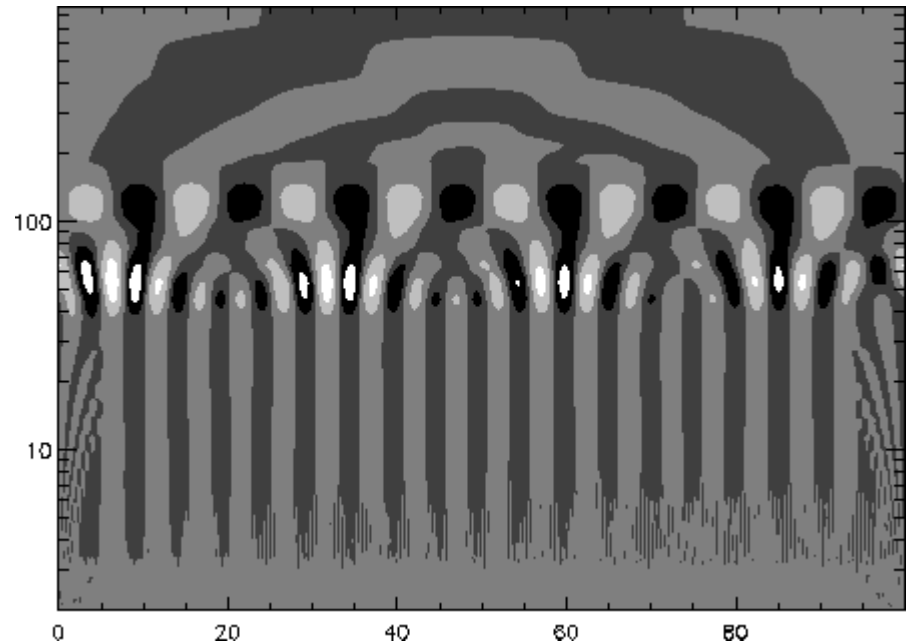
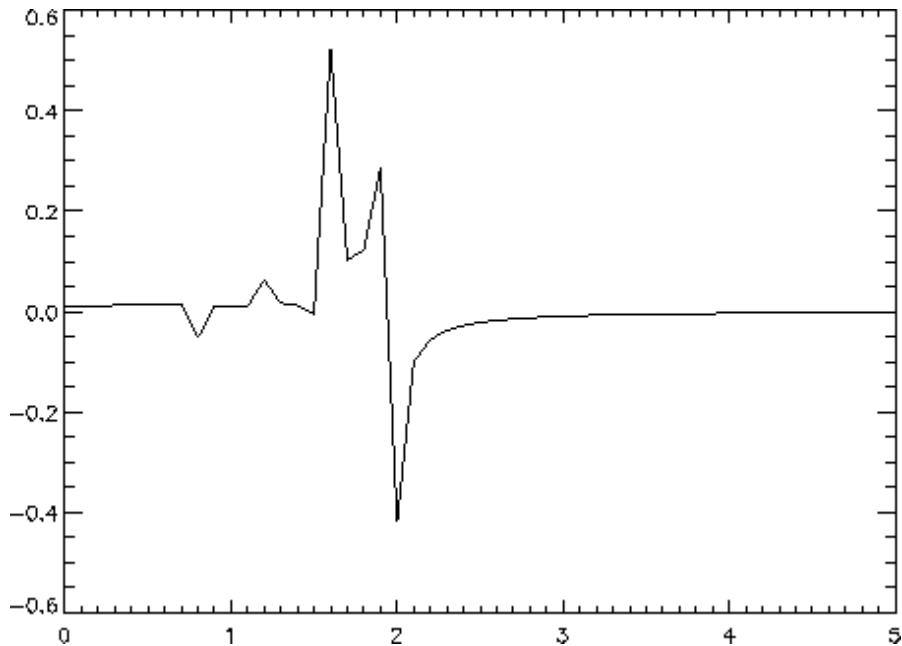
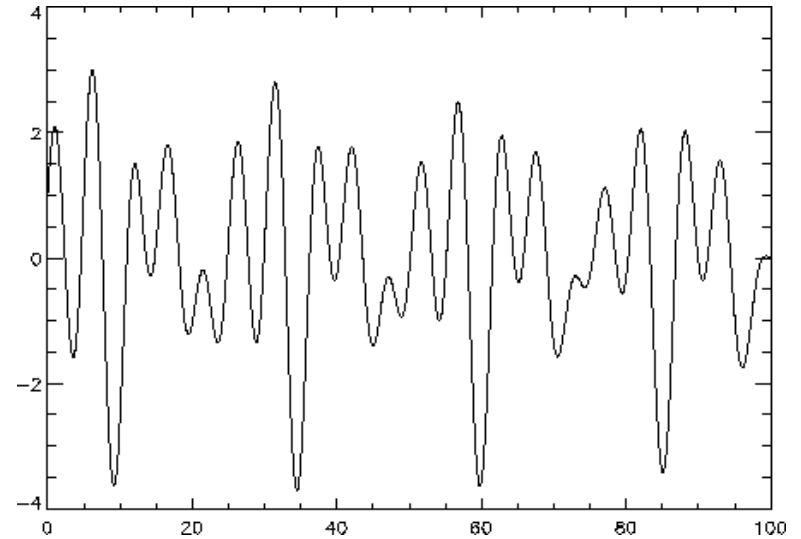

```
SELECT run, camCol, rerun, field, objID,  
u,g,r,i,z,  
ra, dec  
INTO ##results  
FROM PhotoPrimary  
WHERE (u - g) < 0.4 and  
(g - r) < 0.7 and  
(r - i) > 0.4 and  
(i - z) > 0.4
```

Pattern recognition

- Parametric
- Non-parametric

Coordinate transformation

- FFT
- Wavelets



Source extraction

- Sextractor

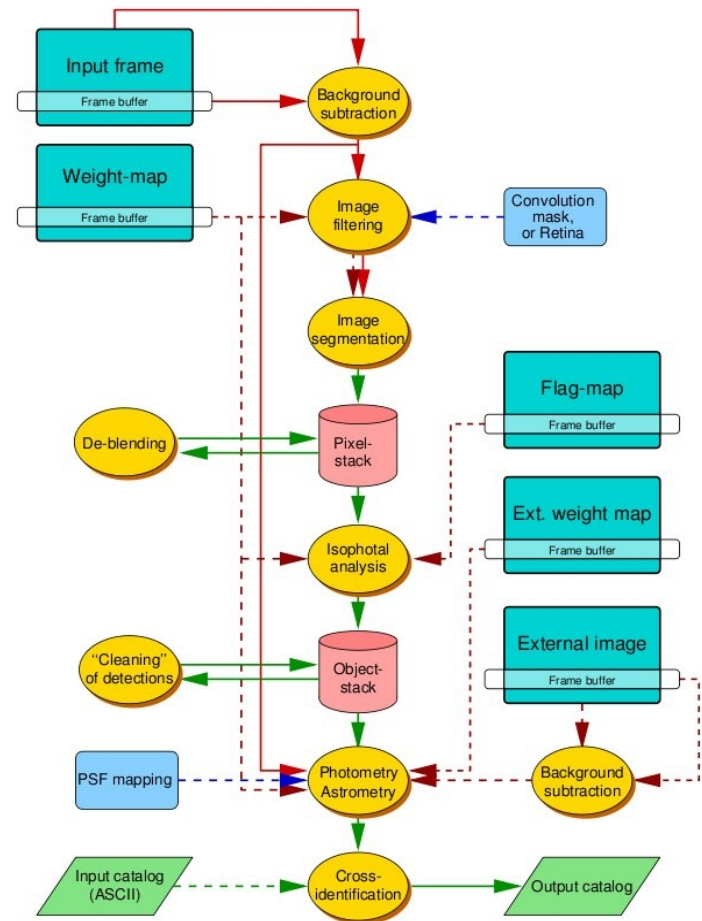
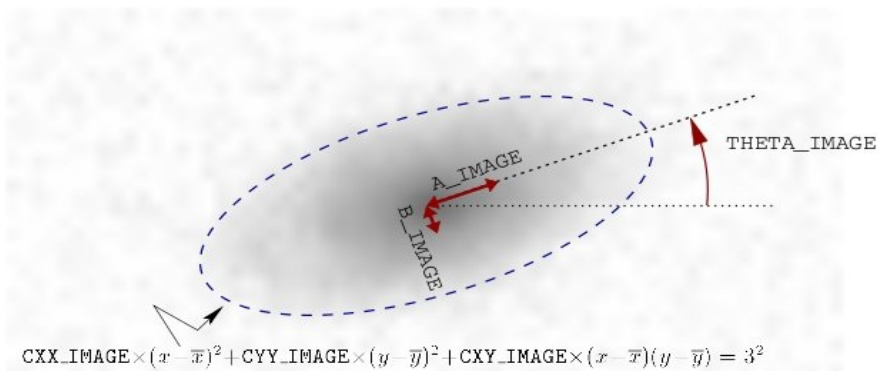
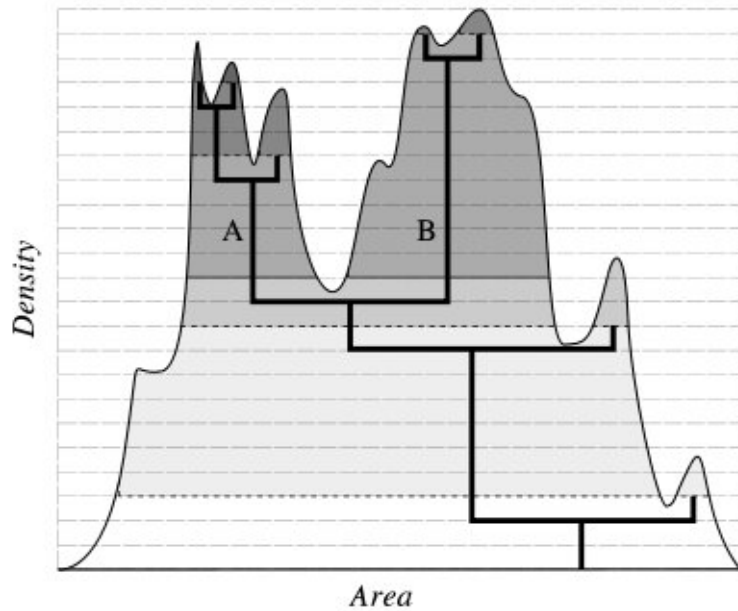


Figure 1: Layout of the main SEXTRACTOR procedures. Dashed arrows represent optional inputs.