

Fourth set of problems for the course on Galaxies, 2005-2006

1. Elliptical galaxies

- (a) If the luminosity density of stars in a galaxy is $j(r) = j_0(r_0/r)^\alpha$, show that the surface brightness at distance R from the center is

$$I(R) = I_0(r_0/R)^{\alpha-1} \quad (1)$$

as long as $\alpha > 1$. What happens if $\alpha < 1$? Compute the total luminosity of the system.

- (b) The fraction of galaxies with apparent axis ratios $(q_0, q_0 + dq_0)$ observed from a random direction θ and true axis ratios $q = \beta/\alpha$, is given by

$$f(q_0)dq_0 = \frac{q_0 dq_0}{\sqrt{(1-q^2)(q_0^2-q^2)}}. \quad (2)$$

Show that if we view these galaxies from random directions, the fraction of oblate elliptical galaxies with true axis ratio q that appear more flattened than q_0 is

$$F(< q_0) = \int_q^{q_0} f(q'_0)dq'_0 = \sqrt{\frac{q_0^2 - q^2}{1 - q^2}} \quad (3)$$

If these galaxies have $q = 0.8$, show that the number seen in the range $0.95 < q_0 < 1$ should be about one-third of those with $0.8 < q_0 < 0.85$. Show that for smaller values of q , an even higher proportion of the images are nearly circular, with $0.95 < q_0 < 1$. You can do the same exercise for $q = 0.6$, and compare your results to the figure shown in class, and explain why it is unlikely that all these galaxies have oblate shapes.

- (c) The virial theorem relates the internal potential energy W and the kinetic energy K of a system in equilibrium through: $2K + W = 0$. Assuming that both the velocity dispersion σ and the mass-to-light ratio M/L are constant throughout a galaxy, and that no dark matter is present, use the virial theorem to show that
- Since the potential energy $W \propto -GM^2/R_e$, where M is the total mass of the galaxy and R_e its effective radius, and the kinetic energy $K \sim M\sigma^2/2$, so the mass of the galaxy should be $M \propto \sigma^2 R_e$.
 - If the surface brightness $I(R)$ of all elliptical galaxies could be described by Sersic's law $I(R) = I_e \exp[-b(R/R_e)^{1/n} - 1]$ with the same value of n , explain why their total luminosity L should follow $L \propto I_e R_e^2$.
 - If all elliptical galaxies had the same mass-to-light ratio M/L and surface brightness at the effective radius I_e , the Faber-Jackson relation is expected.

2. Disk galaxies

In a galaxy where the potential follows the Plummer model

$$\Phi(r) = \frac{-GM}{\sqrt{r^2 + a^2}} \quad (4)$$

find the rotation curve $V(r)$. Show that $V_{max}^2 = 2GM/(3\sqrt{3}a)$. Sketch $V(r)$ for $r \leq 4a$. For inclination $i = 30^\circ$, draw a spider diagram with contours at 0.2, 0.4, 0.6, and 0.8 of V_{max} .